Introduction to Nuclear Fusion as An Energy Source

Po-Yu Chang

Institute of Space and Plasma Sciences, National Cheng Kung University

Lecture 6

2025 spring semester

Tuesday 9:00-12:00

Materials:

https://capst.ncku.edu.tw/PGS/index.php/teaching/

Online courses:

https://nckucc.webex.com/nckucc/j.php?MTID=mf1a33a5dab5eb71de9da43 80ae888592





• Midterm 4/15 (One double-sided A4 cheating sheet is allowed.)

• Final exam 6/3

Charged particles drift across field lines



3

The particle drifts back to the original position if a small poloidal field is superimposed on the toroidal field





• Points with no drift

Stellarator uses twisted coil to generate poloidal magnetic field







https://www.euro-fusion.org/2011/09/tokamak-principle-2/ https://en.wikipedia.org/wiki/Stellarator

A figure-8 stellarator solved the drift issues



Introduction to Plasma Physics and Controlled Fusion 3rd Edition, by Francis F. Chen ₆

Concept of figure-8 stellarator



T. Coor, et al., Phys. Fluids 1, 411 (1958)

Model A stellarator





https://www.autoevolution.com/news/stellarator-reactors-the-onceforgotten-all-american-approach-to-nuclear-fusion-209478.html#agal_2

Different types of stellarators



Helical coils with toroidal field coils can be replaced by smoothed twisting coils

 Superposition of helical windings (blue) and the TF-coils (black) and mapped into the θ-Φ plane.



 Realization of the smoothed twisting coils



Toroidal angle



Wendelstein 7-X is a (Helias) stellarator built by Max Planck Institute for Plasma Physics (IPP)





- No need to drive plasma current. It is intrinsically steady state.
- With zero net current, one potentially dangerous class of MHD instabilities, the current-driven kink modes, is eliminated.
- Magnetic configuration is set by external coils, not by currents in the plasma. Stellarators do not suffer violent disruptions.
- Potential for greater range of designs and optimization of fusion performance.



- Complicated coil configurations. It's difficult to design. The precision requirement is high. It is expensive to build coils for stellarators.
- Achieving good particle confinement in stellarators is more difficult than that in tokamaks.
- Divertors and heat load geometry in stellarators is more complicated than those in tokamaks.

Course Outline



- Magnetic confinement fusion (MCF)
 - Gyro motion, MHD
 - 1D equilibrium (z pinch, theta pinch)
 - Drift: ExB drift, grad B drift, and curvature B drift
 - Tokamak, Stellarator (toroidal field, poloidal field)
 - Magnetic flux surface
 - 2D axisymmetric equilibrium of a torus plasma: Grad-Shafranov equation.
 - Stability (Kink instability, sausage instability, Safety factor Q)
 - Central-solenoid (CS) start-up (discharge) and current drive
 - CS-free current drive: electron cyclotron current drive, bootstrap current.
 - Auxiliary Heating: ECRH, Ohmic heating, Neutral beam injection.

General screw pinch is flexible with varies range of β



An equilibrium state may not be stable



A cylindrical plasma column may not be stable





A cylindrical plasma column is stable when the safety factor is greater than unity



MHD Safety factor: •

Kruskal–Shafranov limit

Theta pinch is stable while z pinch is unstable



$$q(r) = \frac{rB_z(r)}{R_o B_\theta(r)}$$

Course Outline



- Magnetic confinement fusion (MCF)
 - Gyro motion, MHD
 - 1D equilibrium (z pinch, theta pinch)
 - Drift: ExB drift, grad B drift, and curvature B drift
 - Tokamak, Stellarator (toroidal field, poloidal field)
 - Magnetic flux surface
 - 2D axisymmetric equilibrium of a torus plasma: Grad-Shafranov equation.
 - Stability (Kink instability, sausage instability, Safety factor Q)
 - Central-solenoid (CS) start-up (discharge) and current drive
 - CS-free current drive: electron cyclotron current drive, bootstrap current.
 - Auxiliary Heating: ECRH, Ohmic heating, Neutral beam injection.

Ideal MHD



- Continuity eq: $\frac{\partial \rho_m}{\partial t} + \nabla \cdot (\rho_m \, \vec{v}) = 0$ • Momentum eq: $\rho_m \left[\frac{\partial \, \vec{v}}{\partial t} + (\vec{v} \cdot \nabla) \, \vec{v} \right] = \vec{j} \times \vec{B} - \nabla p$
- Ohm's law: $\vec{E} + \vec{v} \times \vec{B} \approx 0$
- Equation of state:

$$\frac{d}{dt}\left(\frac{P}{\rho_{\rm m}\gamma}\right)=0$$

- Maxwell's eqs:
 - $\nabla \cdot \vec{E} \approx 0$

$$\nabla \cdot \vec{B} = 0$$
$$\nabla \times \vec{E} = -\frac{\partial \vec{B}}{\partial t}$$
$$\nabla \times \vec{B} = \mu_0 \vec{j}$$
$$\nabla \cdot \vec{i} = 0$$

- Requirement:
 - High collisionality fluid model
 - Small gyro radius low frequency
 - Small resistivity a perfect conductor

When forces are balances, the system is in the equilibrium state, or called "Magnetohydrostatics"

• Equilibrium state:

$$\rho_{m}\left[\frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \nabla) \vec{v}\right] = \vec{j} \times \vec{B} - \nabla p \equiv 0$$

$$\vec{j} \times \vec{B} = \nabla p$$

$$\vec{j} \times \vec{B} = \frac{1}{\mu_{0}} (\vec{\nabla} \times \vec{B}) \times \vec{B} = \frac{1}{\mu_{0}} \left[(\vec{B} \cdot \vec{\nabla}) \vec{B} - \frac{1}{2} \vec{\nabla} B^{2} \right] = \nabla p \checkmark \vec{X} \times \vec{B} = \mu_{0} \vec{j}$$

$$\nabla \left(P + \frac{B^{2}}{2\mu_{0}} \right) = \frac{1}{\mu_{0}} (\vec{B} \cdot \vec{\nabla}) \vec{B}$$
Magnetic Magnetic \leftarrow Forces caused by pressure tension \leftarrow forces caused by curvature of the field lines
$$\vec{j} \perp \nabla p \qquad \vec{B} \perp \nabla p \qquad \vec{j} \cdot \nabla p = 0 \qquad \vec{B} \cdot \nabla p = 0$$

 The surfaces with p = constant are both magnetic surfaces (i.e., they are made up of magnetic field lines) and current surfaces (i.e., they are made of current flow lines).

2D axisymmetric equilibrium of a torus plasma: **Grad-Shafranov equation**



The surfaces with p = constant are both magnetic surfaces (i.e., they • are made up of magnetic field lines) and current surfaces (i.e., they are made of current flow lines).

 $\nabla \cdot \vec{B} = 0$

Magnetic lines lying on pressure contour



Contours of constant pressure
 Magnetic lines lying on pressure contour



• A magnetic (or flux) surface is one that is everywhere tangential to the field, i.e., the normal to the surface is everywhere perpendicular to B.

$$\vec{J} \times \vec{B} = \nabla p \qquad \nabla \times \vec{B} = \mu_0 \vec{J}$$

$$\nabla \cdot \vec{J} = 0 \qquad \nabla \cdot \vec{B} = 0$$

$$\vec{B} = (B_R, B_{\Phi}, B_z) \qquad \text{Axisymmetric:} \frac{\partial}{\partial \phi} = 0$$

$$\nabla \cdot \vec{B} = 0$$

$$\frac{1}{R} \frac{\partial}{\partial R} (RB_R) + \frac{1}{R} \frac{\partial B_{\Phi}}{\partial \phi} + \frac{\partial B_z}{\partial z} = 0$$

$$\frac{1}{R} \frac{\partial}{\partial R} (RB_R) + \frac{\partial B_z}{\partial z} = 0$$
Represent the magnetic field using a vector potential A:
$$\vec{B} = \nabla \times \vec{A} = \hat{R} \left(\frac{1}{R} \frac{\partial A_z}{\partial \phi} - \frac{\partial A_{\Phi}}{\partial z} \right) + \hat{\phi} \left(\frac{\partial A_R}{\partial z} - \frac{\partial A_z}{\partial R} \right) + \hat{z} \left(\frac{1}{R} \frac{\partial}{\partial R} (RA_{\Phi}) - \frac{1}{R} \frac{\partial A_R}{\partial \phi} \right)$$

$$= \hat{R} \left(-\frac{\partial A_{\Phi}}{\partial z} \right) + \hat{\phi} \left(\frac{\partial A_R}{\partial z} - \frac{\partial A_z}{\partial R} \right) + \hat{z} \left(\frac{1}{R} \frac{\partial}{\partial R} (RA_{\Phi}) \right)$$

$$B_R = -\frac{\partial A_{\Phi}}{\partial z} \quad B_Z = \frac{1}{R} \frac{\partial}{\partial R} (RA_{\Phi})$$

25

Pressure can be written as a function of flux

$$-\frac{1}{R}\frac{\partial}{\partial R}(RB_{R}) + \frac{\partial B_{z}}{\partial z} = 0$$

$$B_{R} = -\frac{\partial A_{\Phi}}{\partial z}$$

$$B_{z} = \frac{1}{R}\frac{\partial}{\partial R}(RA_{\Phi})$$

$$\psi = \frac{1}{2\pi}\int \vec{B} \cdot d\vec{S} = \frac{1}{2\pi}\int (\vec{\nabla} \times \vec{A}) \cdot d\vec{S}$$

$$= \frac{1}{2\pi}\int \vec{A} 2\pi R \cdot d\vec{l} = \int \vec{A} \cdot \hat{\phi}Rdl = RA_{\Phi}$$

$$B_{R} = -\frac{1}{R}\frac{\partial\psi}{\partial z}$$

$$B_{z} = \frac{1}{R}\frac{\partial\psi}{\partial R}$$

$$\Rightarrow \vec{B} \cdot \nabla\psi = B_{R}\frac{\partial\psi}{\partial R} + B_{\Phi}\frac{1}{R}\frac{\partial\psi}{\partial \Phi} + B_{z}\frac{\partial\psi}{\partial z} = B_{R}\frac{\partial\psi}{\partial R} + B_{z}\frac{\partial\psi}{\partial z} = 0$$

$$for \nabla p = 0$$

$$for \nabla p \neq 0$$

$$p = p(\psi)$$

JG KUA

Pressure can be written as a function of flux







• Let's see the $\widehat{\phi}$ component of the force-balance equation:

$$\left(\vec{j}\times\vec{B}=\nabla p\right)_{\phi}$$
 $j_{z}B_{R}-j_{R}B_{z}=\frac{1}{R}\frac{\partial p}{\partial \phi}\equiv 0$

Ampére's law:

$$\nabla \times \vec{B} = \mu_{0} \vec{j}$$

$$\nabla \times \vec{B} = \hat{R} \left(\frac{1}{R} \frac{\partial B_{z}}{\partial \phi} - \frac{\partial B_{\phi}}{\partial z} \right) + \hat{\phi} \left(\frac{\partial B_{R}}{\partial z} - \frac{\partial B_{z}}{\partial R} \right) + \hat{z} \left(\frac{1}{R} \frac{\partial}{\partial R} (RB_{\phi}) - \frac{1}{R} \frac{\partial B_{R}}{\partial \phi} \right)$$

$$= \hat{R} \left(-\frac{\partial B_{\phi}}{\partial z} \right) + \hat{\phi} \left(\frac{\partial B_{R}}{\partial z} - \frac{\partial B_{z}}{\partial R} \right) + \hat{z} \left(\frac{1}{R} \frac{\partial}{\partial R} (RB_{\phi}) \right)$$

$$= \hat{R} \mu_{0} j_{R} + \hat{\phi} \mu_{0} j_{\phi} + \hat{z} \mu_{0} j_{z}$$

$$j_{R} = -\frac{1}{\mu_{0}} \frac{\partial B_{\phi}}{\partial z} \qquad j_{z} = \frac{1}{\mu_{0}} \frac{1}{R} \frac{\partial}{\partial R} (RB_{\phi})$$

$$= \frac{B_{R}}{\partial R} (RB_{\phi}) + B_{z} \frac{\partial B_{\phi}}{\partial z} = 0$$

Magnetic field can be decomposed into the poloidal component and the toroidal component

$$\frac{B_{R}}{R}\frac{\partial}{\partial R}(RB_{\phi})+B_{z}\frac{\partial B_{\phi}}{\partial z}=0 \qquad \Rightarrow \qquad B_{R}\frac{\partial}{\partial R}(RB_{\phi})+B_{z}\frac{\partial}{\partial z}(RB_{\phi})=0$$

$$F \equiv RB_{\phi} \qquad \Rightarrow \qquad B_{R}\frac{\partial F}{\partial R}+B_{z}\frac{\partial F}{\partial z}=0 \qquad \Rightarrow \qquad \overrightarrow{B} \cdot \nabla F=0$$

$$\begin{pmatrix} \frac{\partial}{\partial \phi}=0 \end{pmatrix} \qquad \overrightarrow{B} \cdot \nabla F=0$$

$$\begin{pmatrix} \frac{\partial}{\partial \phi}=0 \end{pmatrix} \qquad \overrightarrow{B} \cdot \nabla F=0$$

$$\begin{pmatrix} \frac{\partial}{\partial \phi}=0 \end{pmatrix} \qquad \overrightarrow{F}=F(\psi)$$

$$B_{R}=-\frac{\partial A_{\phi}}{\partial z}=-\frac{1}{R}\frac{\partial \psi}{\partial z}$$

$$F = F(\psi)$$

$$B_{z}=\frac{1}{R}\frac{\partial}{\partial R}(RA_{\phi})=\frac{1}{R}\frac{\partial \psi}{\partial R}$$

$$(\psi = RA_{\phi})$$

$$B_{\phi}=\frac{F(\psi)}{R}$$

$$\overrightarrow{B}=\widehat{R}B_{R}+\widehat{\phi}B_{\phi}+\widehat{z}B_{z}=\widehat{R}\left(-\frac{1}{R}\frac{\partial \psi}{\partial z}\right)+\widehat{\phi}\left(\frac{F(\psi)}{R}\right)+\widehat{z}\left(\frac{1}{R}\frac{\partial \psi}{\partial R}\right)$$

$$= \left(\frac{\nabla \psi}{R}\right)\times\widehat{\phi}+\frac{F(\psi)}{R}\widehat{\phi}$$
Poloidal component \overrightarrow{B}_{T}

Arbitrary integration constant associated with flux can be chosen such that flux equals to zero on the field axis

• The poloidal flux of the area of a washershaped surface lying in the z = 0 plane from $R = R_a$ to an arbitrary ψ contour defined by $\psi = \psi(R_b, 0)$:

$$\psi_{\rm P} \equiv \frac{1}{2\pi} \int \vec{B}_{\rm P} \cdot d\vec{S}$$
$$= \frac{1}{2\pi} \int_{0}^{2\pi} d\phi \int_{R_{\rm a}}^{R_{\rm b}} dRRB_{z}(R,0)$$
$$= \psi(R_{\rm b},0) - \psi(R_{\rm a},0)$$
$$\equiv \psi(R_{\rm b},0)$$

where $\psi(R_a, 0) \equiv 0$ is chosen.



Let's see the \hat{R} component of the force-balance equation: $B_{\rm R} = -\frac{1}{R}\frac{\partial\psi}{\partial z}$ $j_{\phi}B_z - j_z B_{\phi} = \frac{\partial p}{\partial P}$ $(\overrightarrow{j} \times \overrightarrow{B} = \nabla p)_{\mathbf{n}}$ $B_{\rm z} = \frac{1}{P} \frac{\partial \psi}{\partial P}$ Ampére's law: $B_{\phi} = \frac{F(\psi)}{P}$ $\nabla \times \vec{B} = \mu_0 \vec{i}$ $\nabla \times \vec{B} = \hat{R} \left(-\frac{\partial B_{\phi}}{\partial z} \right) + \hat{\phi} \left(\frac{\partial B_{R}}{\partial z} - \frac{\partial B_{z}}{\partial R} \right) + \hat{z} \left(\frac{1}{R} \frac{\partial}{\partial R} (RB_{\phi}) \right) = \hat{R} \mu_{o} j_{R} + \hat{\phi} \mu_{o} j_{\phi} + \hat{z} \mu_{o} j_{z}$ $\mu_{0}j_{\phi} = \frac{\partial B_{R}}{\partial z} - \frac{\partial B_{z}}{\partial R} = \frac{\partial}{\partial z} \left(-\frac{1}{R} \frac{\partial \psi}{\partial z} \right) - \frac{\partial}{\partial R} \left(\frac{1}{R} \frac{\partial \psi}{\partial R} \right) = -\frac{1}{R} \frac{\partial^{2} \psi}{\partial z^{2}} - \frac{1}{R} \frac{\partial^{2} \psi}{\partial R^{2}} + \frac{1}{R^{2}} \frac{\partial \psi}{\partial R}$ $\equiv -\frac{1}{R}\Delta^*\psi \qquad \text{where } \Delta^*\psi \equiv \frac{\partial^2\psi}{\partial z^2} + R\frac{\partial}{\partial R}\left(\frac{1}{R}\frac{\partial\psi}{\partial R}\right) = R^2\nabla\cdot\left(\frac{\nabla\psi}{R^2}\right)$ $\mu_{0}j_{z} = \frac{1}{R}\frac{\partial}{\partial R}(RB_{\phi}) = \frac{1}{R}\frac{\partial F}{\partial R} = \frac{1}{R}\frac{dF}{dy}\frac{\partial \psi}{\partial R}$

$$j_{\phi}B_{z} - j_{z}B_{\phi} = \frac{\partial p}{\partial R}$$

$$B_{\phi} = \frac{F}{R}$$

$$j_{\phi} = -\frac{1}{\mu_{o}R}\Delta^{*}\psi$$

$$B_{z} = \frac{1}{R}\frac{\partial\psi}{\partial R}$$

$$j_{z} = \frac{1}{\mu_{o}R}\frac{dF}{d\psi}\frac{\partial\psi}{\partial R}$$

$$-\frac{1}{\mu_{o}R}\Delta^{*}\psi\frac{1}{R}\frac{\partial\psi}{\partial R} - \frac{1}{\mu_{o}R}\frac{dF}{d\psi}\frac{\partial\psi}{\partial R}\frac{F}{R} = \frac{dp}{d\psi}\frac{\partial\psi}{\partial R}$$

$$-\Delta^{*}\psi\frac{1}{R}\frac{1}{R^{2}} - \frac{1}{\mu_{o}}\frac{F}{R^{2}}\frac{dF}{d\psi} = \frac{dp}{d\psi}$$
Grad - Shafranov equation: $\Delta^{*}\psi = -\mu_{o}R^{2}\frac{dp}{d\psi} - \frac{1}{2}\frac{dF^{2}}{d\psi}$
where $\Delta^{*}\psi = R^{2}\nabla \cdot \left(\frac{\nabla\psi}{R^{2}}\right)$

$$\overline{B} = \left(\frac{\nabla\psi}{R}\right) \times \widehat{\phi} + \frac{F(\psi)}{R}\widehat{\phi}$$

$$\Delta^{*}\psi = -\mu_{0}R^{2}\frac{dp}{d\psi} - \frac{1}{2}\frac{dF^{2}}{d\psi} \text{ where } \Delta^{*}\psi = R^{2}\nabla \cdot \left(\frac{\nabla\psi}{R^{2}}\right) \quad \vec{B} = \left(\frac{\nabla\psi}{R}\right) \times \hat{\phi} + \frac{F(\psi)}{R}\hat{\phi}$$

$$\mu_{0}j_{\phi} = -\frac{1}{R}\Delta^{*}\psi \qquad \mu_{0}j_{z} = \frac{1}{R}\frac{\partial F}{\partial R} \qquad F \equiv RB_{\phi}$$

$$\mu_{0}j_{R} = -\frac{\partial B_{\phi}}{\partial z} = -\frac{1}{R}\frac{\partial}{\partial z}(RB_{\phi}) = -\frac{1}{R}\frac{\partial F}{\partial z}$$

$$\mu_{0}\vec{j} = \hat{R}\mu_{0}j_{R} + \hat{\phi}\mu_{0}j_{\phi} + \hat{z}\mu_{0}j_{z} = \hat{R}\left(-\frac{1}{R}\frac{\partial F}{\partial z}\right) + \hat{\phi}\left(-\frac{1}{R}\Delta^{*}\psi\right) + \hat{z}\left(\frac{1}{R}\frac{\partial F}{\partial R}\right)$$

$$= \left(\frac{\nabla F}{R}\right) \times \hat{\phi} + \left(-\frac{1}{R}\Delta^{*}\psi\right)\hat{\phi}$$

$$I_{P} = \int \vec{j}_{P} \cdot d\vec{S} = -\int_{0}^{2\pi} d\phi \int_{0}^{R_{b}} dRRj_{z}(R, 0)$$

$$= -2\pi \int_{0}^{R_{b}} dRR\frac{1}{R}\frac{\partial F(R, 0)}{\partial R} = -2\pi F(\psi)$$

Plasma condition can be obtained by solving Grad-Shafranov equation

$$\Delta^* \psi = -\mu_0 R^2 \frac{dp}{d\psi} - \frac{1}{2} \frac{dF^2}{d\psi} \qquad \text{where } \Delta^* \psi = R^2 \nabla \cdot \left(\frac{\nabla \psi}{R^2}\right) \qquad F = RB_{\phi} = -\frac{I_{\rm P}}{2\pi}$$

- The usual strategy to solve the Grad-Shafranov equation:
 - 1. Specify two free functions, the plasma pressure $p = p(\psi)$ and the toroidal field function $F = F(\psi)$.
 - 2. Solve the equation with specified boundary conditions to determine the flux function $\psi(R, z)$.
 - 3. Calculation the magnetic field using the following equations:

$$B_{\rm R} = -\frac{1}{R} \frac{\partial \psi}{\partial z}$$
 $B_{\phi} = \frac{F(\psi)}{R}$ $B_{\rm z} = \frac{1}{R} \frac{\partial \psi}{\partial R}$

4. The pressure profile can then be obtained from $p = p(\psi(R, z))$.

Example of the analytical solution of the Grad-Shafranov equation



35

Magnetically confined toroidal equilibrium



- 1. Radial pressure balance in the poloidal plan needs to be provided so that the pressure contours form closed nested surfaces. Both toroidal and poloidal fields can readily accomplish this task.
- 2. The radially outward expansion force inherent in all toroidal geometries needs to be balanced without sacrificing stability.
- Forces associated with toroidal force balance are usually than those corresponding to radial pressure balance. However, they are more difficult to compensate.



Toroidal configuration with a purely poloidal magnetic field



The outward force can be compensated by either a perfectly conducting shell or externally applied vertical field

Externally applied vertical field

Perfectly conducting shell



- With a finite conductivity wall, flux can only remain compressed for about a skin time.
- This configuration develops disastrous MHD instabilities (z pinch).

Toroidal configuration with a purely toroidal magnetic field, stable but NOT balanced



Coils in a tokamak



- Toroidal field coils (in poloidal direction) generate toroidal field for confinement.
- Poloidal field coils generate vertical field for plasma positioning and shaping.
- Central solenoid for breakdown and generating plasma current (in toroidal direction) and thus generating poloidal field for confinement.

Plasma condition can be obtained by solving Grad-Shafranov equation

$$\Delta^* \psi = -\mu_0 R^2 \frac{dp}{d\psi} - \frac{1}{2} \frac{dF^2}{d\psi} \qquad \text{where } \Delta^* \psi = R^2 \nabla \cdot \left(\frac{\nabla \psi}{R^2}\right) \qquad F = RB_{\phi} = -\frac{I_{\rm P}}{2\pi}$$

- The usual strategy to solve the Grad-Shafranov equation:
 - 1. Specify two free functions, the plasma pressure $p = p(\psi)$ and the toroidal field function $F = F(\psi)$.
 - 2. Solve the equation with specified boundary conditions to determine the flux function $\psi(R, z)$.
 - 3. Calculation the magnetic field using the following equations:

$$B_{\rm R} = -\frac{1}{R} \frac{\partial \psi}{\partial z}$$
 $B_{\phi} = \frac{F(\psi)}{R}$ $B_{\rm z} = \frac{1}{R} \frac{\partial \psi}{\partial R}$

4. The pressure profile can then be obtained from $p = p(\psi(R, z))$.

Application of solving Grad-Shafranov equation for designing a tokamak



- Given I_{plasma} , $p(\psi)$, $I(\psi)$, I_{coils} , free boundary of plasma, perfect conductor as the chamber.
- Given I_{plasma} , $p(\psi)$, $I(\psi)$, I_{coils} , free boundary of plasma, insulator chamber.
- Given I_{plasma} , $p(\psi)$, $I(\psi)$, I_{coils} , free boundary of plasma, chamber with eddy current.
- Given I_{plasma} , $p(\psi)$, $I(\psi)$, fixed boundary of plasma. Then, use I_{coils} , free boundary of plasma and match the plasma shape calculated in the fixed boundary condition.

$$\Delta^* \psi = -\mu_0 R^2 \frac{dp}{d\psi} - \frac{1}{2} \frac{dF^2}{d\psi} \quad \text{where } \Delta^* \psi = R^2 \nabla \cdot \left(\frac{\nabla \psi}{R^2}\right)$$
$$I_P = -2\pi F(\psi)$$
$$\mu_0 \overrightarrow{j} = \left(\frac{\nabla F}{R}\right) \times \widehat{\phi} + \left(-\frac{1}{R}\Delta^* \psi\right) \widehat{\phi} \quad \overrightarrow{B} = \left(\frac{\nabla \psi}{R}\right) \times \widehat{\phi} + \frac{F(\psi)}{R} \widehat{\phi}$$

Application of solving Grad-Shafranov equation for reconstructing a tokamak equilibrium state

- Measure
 - boundary conditions, including ψ , *B*, etc., on the wall (using flux loop and B-dot probe).
 - Pressure.
 - Plasma current (using Rogowski coil).
- Reconstruct $\psi(r,z)$, j, $p(\psi)$, $l(\psi)$, etc.

$$\Delta^* \psi = -\mu_0 R^2 \frac{dp}{d\psi} - \frac{1}{2} \frac{dF^2}{d\psi} \quad \text{where } \Delta^* \psi = R^2 \nabla \cdot \left(\frac{\nabla \psi}{R^2}\right)$$
$$I_P = -2\pi F(\psi)$$
$$\mu_0 \overrightarrow{j} = \left(\frac{\nabla F}{R}\right) \times \widehat{\phi} + \left(-\frac{1}{R}\Delta^* \psi\right) \widehat{\phi} \quad \overrightarrow{B} = \left(\frac{\nabla \psi}{R}\right) \times \widehat{\phi} + \frac{F(\psi)}{R} \widehat{\phi}$$

Fluxes and currents





Normalized plasma pressure, β



Different poloidal shapes



Safety factor

A LAND ROAD

• Kink Safety Factor:

$$q^*(r) = \frac{aB_o}{R_oB_p} = \frac{2\pi a^2 \kappa B_o}{\mu_o R_o I_o}$$





$$q(r) = \frac{rB_z(r)}{R_o B_\theta(r)}$$

Safety factor



C. C. Baker, et al, Nuclear Technology/Fusion, 1, 5 (1981)

Volume of the constant flux



Magnetic well

$$\widehat{W} = 2 rac{V}{\langle B^2
angle} rac{d}{dV} \left\langle rac{B^2}{2}
ight
angle$$
 $= 2 rac{V}{\langle B^2
angle} rac{d}{dV} \left\langle \mu_o p + rac{B^2}{2}
ight
angle$



EQUILIBRIUM WITH LINEAR

STABILITY AND NONLINEAR

INSTABILITY

 A magnetic well is a quantity that measures plasma stability against short perpendicular wavelength modes driven by the plasma pressure gradient.

EQUILIBRIUM WITH

LINEAR INSTABILITY AND NONLINEAR

STABILITY

The Spherical tokamak





51

The Spherical tokamak



- Aspect ratio R_o/a ~ 1.6
- Advantages:
 - Higher β_t limit.
 - A compact design almost spherical in appearance.
- Challenges:
 - Minimum space is given in the center of the torus to accommodate the toroidal field coils.
 - With a very compact design the technology associated with the construction and maintenance of the device may be more difficult than for a "normal" tokamak.
 - Large currents will have to be driven noninductively, a costly and physically difficult requirement.

Limiter protects the vacuum chamber from plasma bombardment and defines the edge of the plasma



Limiter protects the vacuum chamber from plasma bombardment and defines the edge of the plasma

- A mechanical limiter is a robust piece of material, often made of tungsten, molybdenum, or graphite placed inside the vacuum chamber.
- Some of the particles of the limiter surface may escape. Neutral particles can penetrate some distance into the plasma before being ionized.
- The high-z impurities can lead to significant additional energy loss in the plasma through radiation.
- In ignition experiments and fusion reactors, the bombardment is more intense and extends over longer periods of time. In addition, if the impurity level is too high, it may not be possible to achieve a high enough temperature to ignite.



The magnetic divertor – guide a narrower layer of magnetic lines away from the edge of the plasma



Pros and cons of a divertor



- Advantages:
 - The collector plate is remote from the plasma. There is space available to spread out the magnetic lines.
 - A lower intensity of particles and energy bombard the collector plate leading to a longer replacement time.
 - It is more difficult for impurities to migrate into the plasma.
 - There are longer distance distances to travel and if a neutral particle becomes ionized before or during the time it crosses the divertor layer on its way toward the plasma, its parallel motion then carries it back to the collector plate.
 - The larger divertor chamber provides more access to pump out impurities.
 - The plasma edge is not in direct contact with a solid material such as a limiter.
- Disadvantages: larger and more complex system and more expensive.