

# **Introduction to Nuclear Fusion as An Energy Source**

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**Po-Yu Chang**

**Institute of Space and Plasma Sciences, National Cheng Kung University**

**Lecture 14**

**2024 spring semester**

**Wednesday 9:10-12:00**

**Materials:**

**<https://capst.ncku.edu.tw/PGS/index.php/teaching/>**

**Online courses:**

**<https://nckucc.webex.com/nckucc/j.php?MTID=ma76b50f97b1c6d72db61de9eaa9f0b27>**

# Course Outline

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- Inertial confinement fusion (ICF)
  - Plasma frequency and critical density
  - Direct- and indirect- drive
  - Laser generated pressure (Inverse bremsstrahlung and Ablation pressure)
  - Burning fraction, why compressing a capsule?
  - Implosion dynamics
  - Shock (Compression with different adiabat)
  - Laser pulse shape
  - Rocket model, shell velocity
  - Laser-plasma interaction (Stimulated Raman Scattering, SRS; Stimulated Brillouin Scattering, SBS; Two-plasmon decay )
  - Instabilities (Rayleigh-taylor instability, Kelvin-Helmholtz instability, Richtmeyer-Meshkov instability)



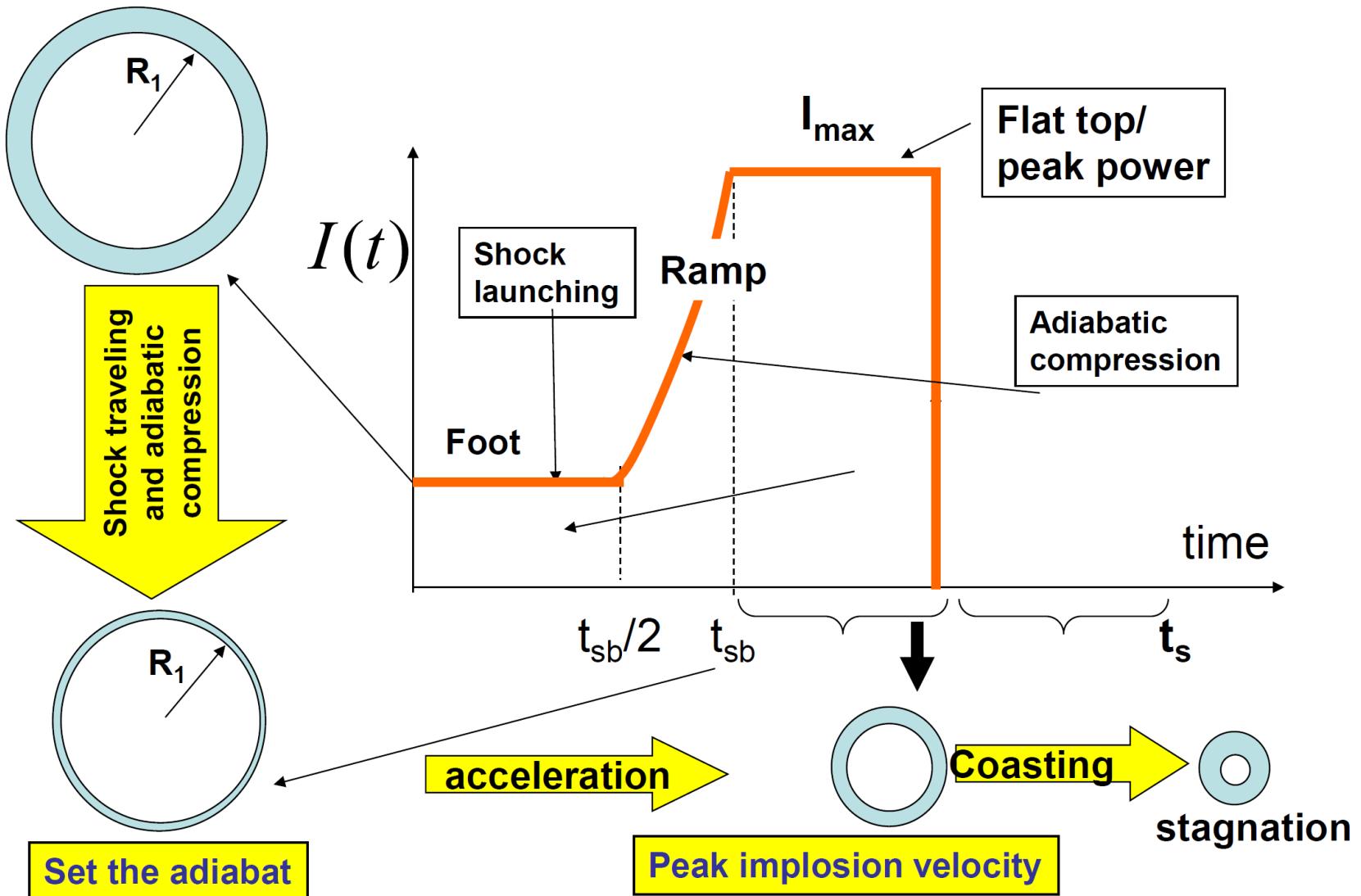
# Reference for ICF

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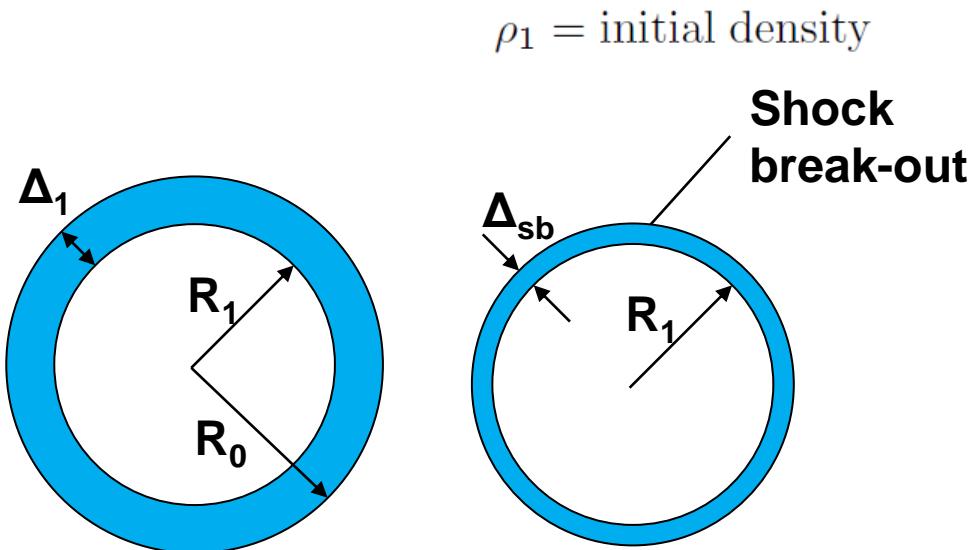


- Riccardo Betti, University of Rochester, HEDSA HEDP summer school, San Diego, CA, August 16-21, 2015
- ICF lectures for course PHY558/ME533
- The physics of inertial fusion, by S. Atzeni, J. Meyer-Ter-Vehn

# There are three stages in the laser pulse: foot, ramp, and flat top



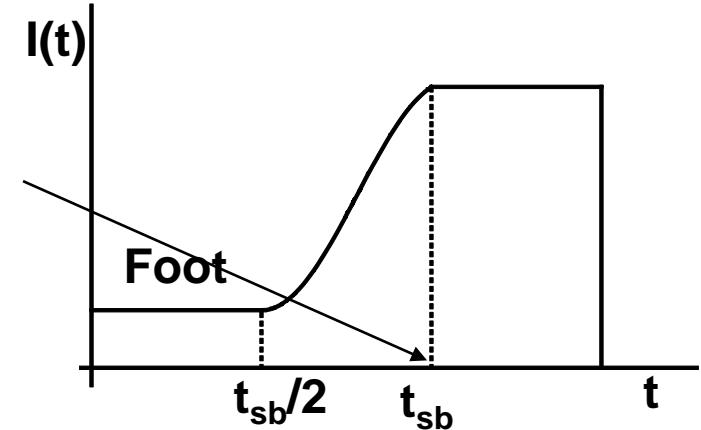
# The adiabat is set by the shock launched by the foot of the laser pulse



$$m_{\text{sb}} \sim 4\pi R_1^2 \Delta_1 \rho_1 = 4\pi R_1^2 \Delta_{\text{sb}} \rho_{\text{sb}}$$

$$\Delta_1 \rho_1 = \Delta_{\text{sb}} \rho_{\text{sb}}$$

$$\Delta_{\text{sb}} = \Delta_1 \frac{\rho_1}{\rho_{\text{sb}}} \sim \Delta_1 \frac{\rho_1}{4\rho_1 \left(\frac{p_{\text{max}}}{p_{\text{foot}}}\right)^{3/5}} = \frac{\Delta_1}{4} \left(\frac{p_{\text{foot}}}{p_{\text{max}}}\right)^{3/5}$$



$$\alpha \sim \frac{p}{\rho^{5/3}} \sim \frac{p_{\text{foot}}}{(4\rho_1)^{5/3}}$$

$$\rho_{\text{sb}} \sim \left(\frac{p_{\text{max}}}{\alpha}\right)^{5/3} \stackrel{\downarrow}{=} 4\rho_1 \left(\frac{p_{\text{max}}}{p_{\text{foot}}}\right)^{5/3}$$

# Density and thickness at shock break out time are expressed in laser intensity

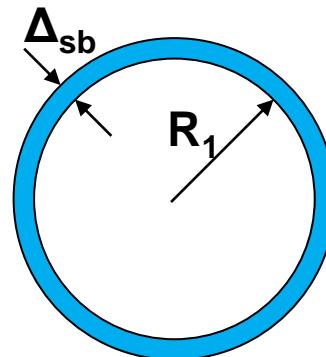


- Use  $p \sim I^{2/3}$

- Shell density  $\rho_{\text{sb}} \sim \rho_1 \left( \frac{p_{\text{max}}}{p_{\text{foot}}} \right)^{5/3} = 4\rho_1 \left( \frac{I_{\text{max}}}{I_{\text{foot}}} \right)^{2/5}$

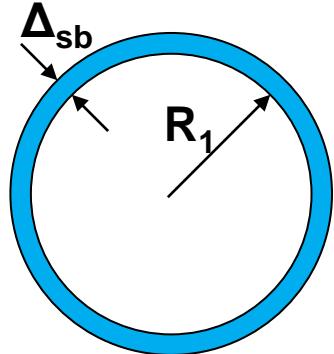
- Shell thickness  $\Delta_{\text{sb}} \sim \frac{\Delta_1}{4} \left( \frac{p_{\text{foot}}}{p_{\text{max}}} \right)^{3/5} = \frac{\Delta_1}{4} \left( \frac{I_{\text{foot}}}{I_{\text{max}}} \right)^{2/5}$

- Shell radius  $R \approx R_1$



# The aspect ratio is maximum at shock break out

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$$\text{Aspect ratio} \equiv \frac{R}{\Delta}$$

$$A_1 = \frac{R_1}{\Delta_1} = \text{initial aspect ratio}$$

$$A_{\text{sb}} = \text{IFAR} = \frac{R_1}{\Delta_{\text{sb}}} = 4A_1 \left( \frac{I_{\max}}{I_{\text{foot}}} \right)^{2/5}$$

$$A_{\text{sb}} = A_{\max}$$

**IFAR  $\equiv$  Maximum In-Flight-Aspect-Ratio = aspect ratio at shock break-out**

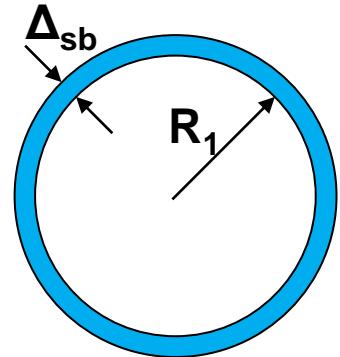
# The IFAR scales with the Mach number



- The shell kinetic energy = the work done on the shell

$$Mu_{max}^2 \sim - \int_{R_1}^R pr^2 dr \sim p(R_1^3 - R^3) \approx pR_1^3 \quad R_1^3 = \frac{Mu_{max}^2}{p}$$

$$M \sim \rho_{sb} \Delta_{sb} R_1^2 \quad \Delta_{sb} \sim \frac{M}{\rho_{sb} R_1^2} \quad R_1 \gg R$$



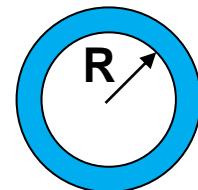
$$IFAR = \frac{R_1}{\Delta_{sb}} = \frac{R_1}{\frac{M}{\rho_{sb} R_1^2}} = \frac{\rho_{sb} R_1^3}{M} = \frac{\rho_{sb}}{M} \frac{Mu_{max}^2}{p}$$

$$= \frac{u_{max}^2}{p/\rho_{sb}} \sim Mach_{max}^2$$

$$\alpha \sim \frac{p}{\rho^{5/3}}$$

$$p \sim I^{2/3}$$

$$IFAR \sim \frac{u_{max}^2}{\alpha^{3/5} I^{4/15}}$$



# The final implosion velocity can be found using IFAR

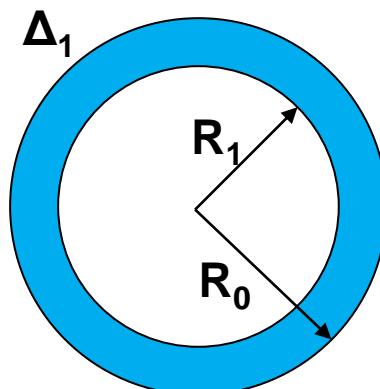
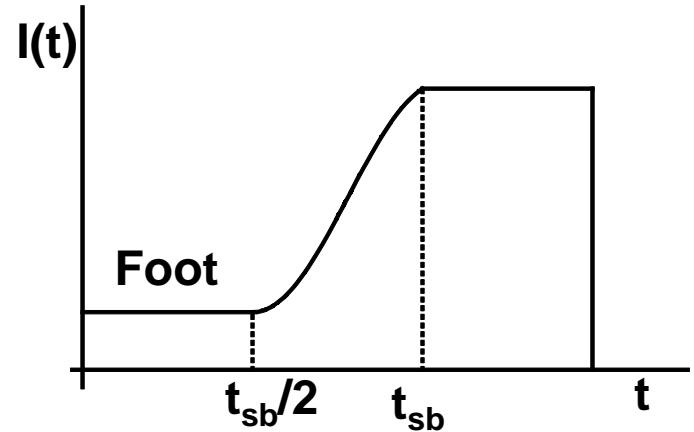


$$u_{\max}^2 \sim IFAR \times \alpha^{3/5} I^{4/15}$$

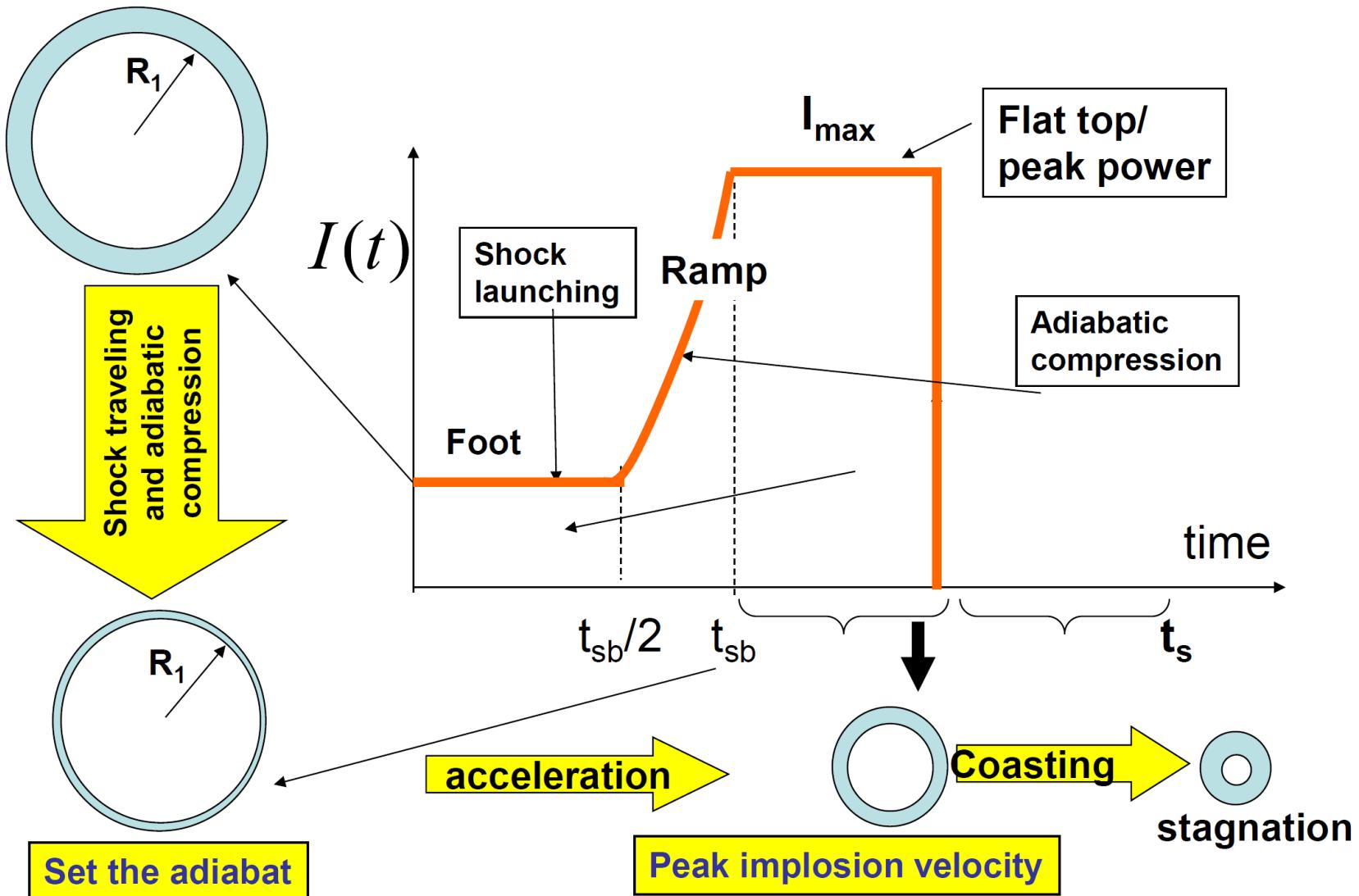
$$IFAR = 4A_1 \left( \frac{I_{\max}}{I_{\text{foot}}} \right)^{2/5}$$

$$A_1 = \frac{R_1}{\Delta_1}$$

$$u_{\max, \text{cm/s}} \approx 10^7 \sqrt{0.7 A_1 \alpha^{3/5} I_{15, \max}^{4/15} \left( \frac{I_{\max}}{I_{\text{foot}}} \right)^{2/5}}$$



# There are three stages in the laser pulse: foot, ramp, and flat top

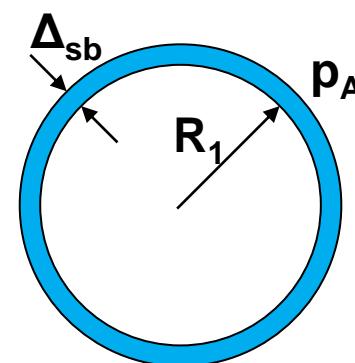


# A simple implosion theory can be derived in the limit of infinite initial aspect ratio

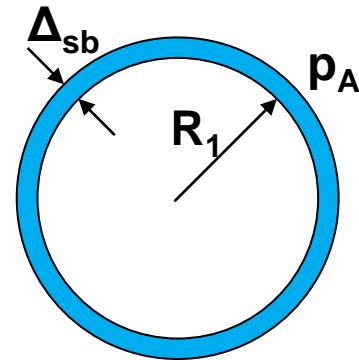
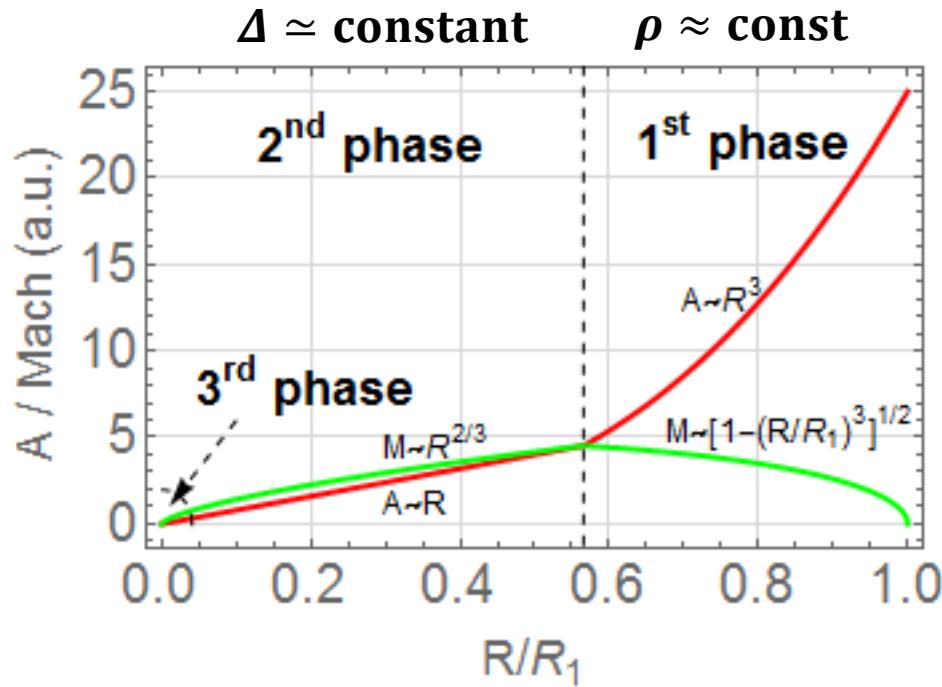


- Start from a high aspect ratio shell (thin shell) at the beginning of the acceleration phase
  - Constant ablated pressure
  - The adiabat is set and kept fixed by the first and the only shock

$$IFAR = A_{sb} = \frac{R_1}{\Delta_{sb}} \gg 1$$



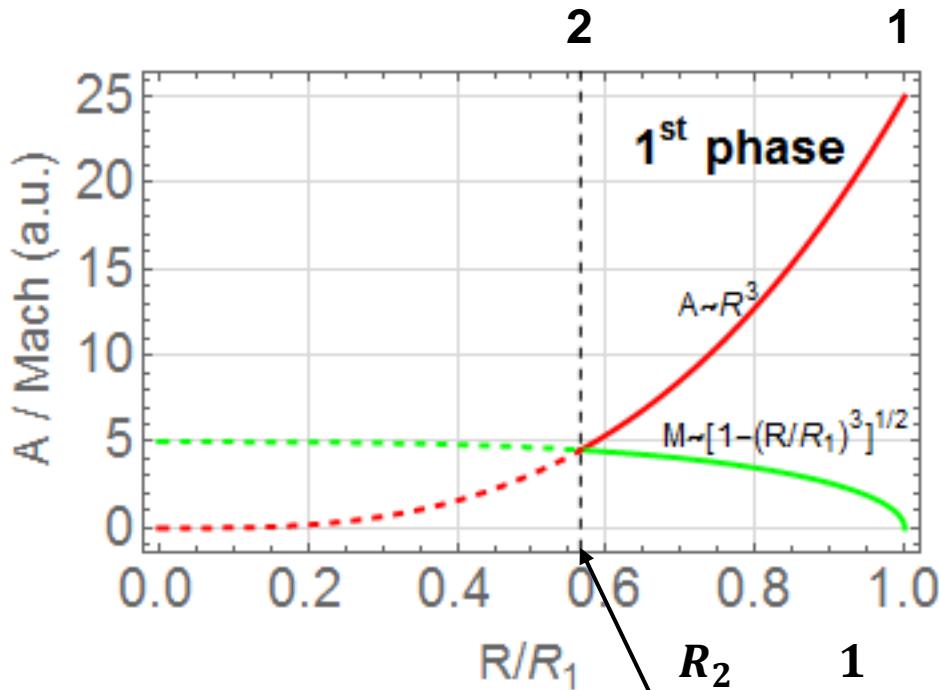
# The implosion are divided in 3 phases after the shock break out



$$\text{Aspect ratio} = A \equiv \frac{R}{\Delta}$$

- 1<sup>st</sup> phase: acceleration
- 2<sup>nd</sup> phase: coasting
- 3<sup>rd</sup> phase: stagnation

# Summary of phase 1 (acceleration phase)



$$\frac{1}{A_{\text{sb}}^{1/6}} < \frac{R}{R_1} \leq 1$$

$$A = A_{\text{sb}} \left( \frac{R}{R_1} \right)^3 = \text{IFAR} \left( \frac{R}{R_1} \right)^3$$

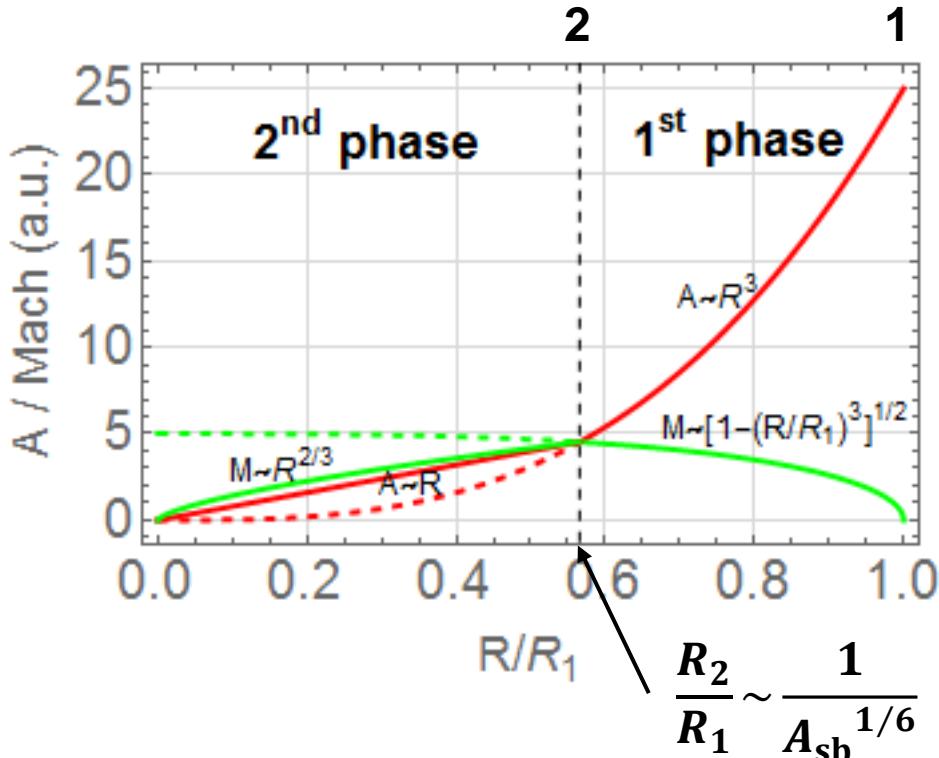
$$\text{Mach} = \text{Mach}_{\max} \left[ 1 - \left( \frac{R}{R_1} \right)^3 \right]^{1/2}$$

$$M \sim 4\pi R^2 \Delta \rho \Rightarrow \Delta \sim R^{-2}$$

$$A = \frac{R}{\Delta} \sim R^3 \Rightarrow A = A_{\text{sb}} \left( \frac{R}{R_1} \right)^3$$

$$\text{Mach}_2 \simeq \text{Mach}_{\max} \left( 1 - \frac{1}{\sqrt{A_{\text{sb}}}} \right)^{1/2} \simeq \text{Mach}_{\max} = \sqrt{A_{\text{sb}}} \quad A_2 \sim \sqrt{A_{\text{sb}}}$$

# Summary of phase 2 (coasting phase)



$$1 < A < \sqrt{A_{\text{sb}}} \quad A < \text{Mach}$$

$$\frac{1}{\sqrt{A_{\text{sb}}}} \sim \frac{1}{A_2} < \frac{R}{R_2} < 1$$

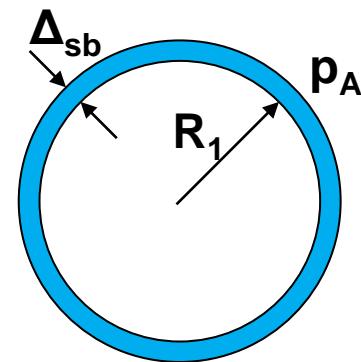
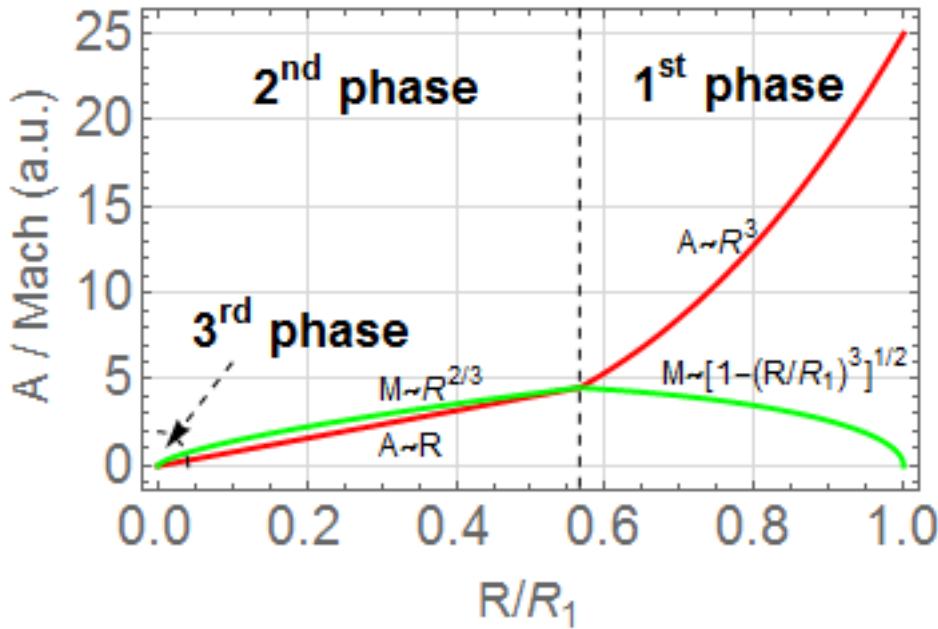
$$A = A_2 \left( \frac{R}{R_2} \right) \sim \sqrt{A_{\text{sb}}} \left( \frac{R}{R_2} \right)$$

$$\text{Mach} \sim \text{Mach}_2 \left( \frac{R}{R_2} \right)^{2/3} \sim \sqrt{A_{\text{sb}}} \left( \frac{R}{R_2} \right)^{2/3}$$

$$\text{Mach}_2 = \text{Mach}_{\max} \simeq A_2 = \sqrt{A_{\text{sb}}}$$

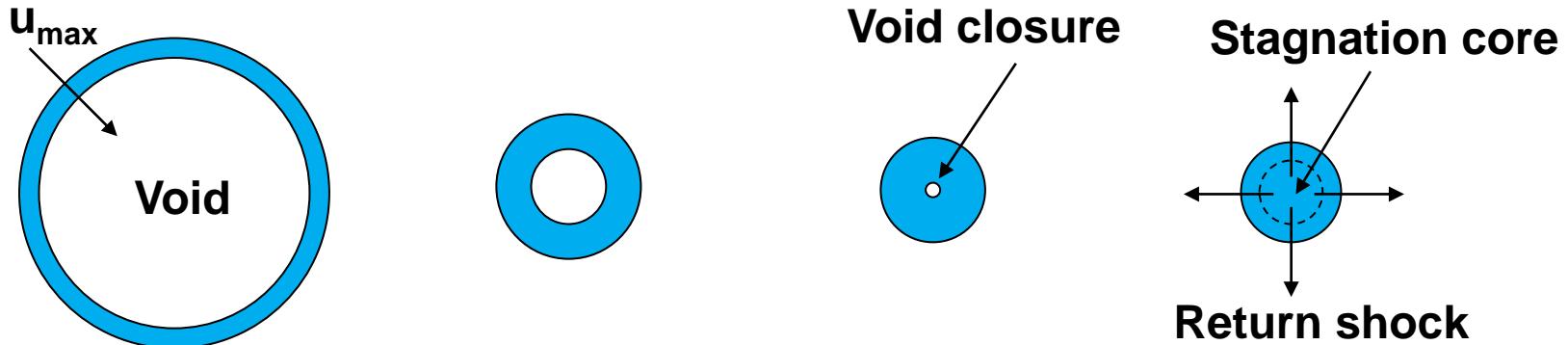
$$\Delta \simeq \text{constant} = \Delta_2 \sim \frac{R_1}{A_{\text{sb}}^{2/3}} \quad \bar{\rho} \simeq \rho_2 \left( \frac{R_2}{R} \right)^2 \sim \rho_{\text{sb}} \left( \frac{R_2}{R} \right)^2$$

# How about the 3<sup>rd</sup> phase where $A \sim 1$ ?



- 1<sup>st</sup> phase: acceleration
- 2<sup>nd</sup> phase: coasting
- 3<sup>rd</sup> phase: stagnation

# The thin shell model breaks down when $A \sim 1$



- When  $A \sim 1 \Rightarrow \Delta \sim R$ , the “void” inside the shell closes and a “return shock” propagating outward is generated due to the collision of the shell with itself
- The density is compressed by a factor no more than 4 even if the strong shock is generated

$\rho_{st} \sim 4\rho_3 \sim \rho_3$  where  $\rho_3$  is the density right before the void closure

# The stagnated density scales with square of the maximum Mach number

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$$\rho_3 \sim \rho_2 \left( \frac{R_2}{R_3} \right)^2 \sim \rho_{\text{sb}} \left( \frac{R_2}{R_3} \right)^2 \quad \bar{\rho} \simeq \rho_2 \left( \frac{R_2}{R} \right)^2$$

$$A = A_3 \sim 1 \Rightarrow \frac{R_3}{\Delta_3} \sim \frac{R_3}{\Delta_2} \sim 1 \Rightarrow R_3 \sim \Delta_2$$

$$\rho_{\text{st}} \sim \rho_3 \sim \rho_{\text{sb}} \left( \frac{R_2}{\Delta_2} \right)^2 \sim \rho_{\text{sb}} A_2^2 \sim \rho_{\text{sb}} \text{Mach}_2^2 \sim \rho_{\text{sb}} \text{Mach}_{\max}^2$$

$$\frac{\rho_{\text{st}}}{\rho_{\text{sb}}} \sim \text{Mach}_{\max}^2$$



**Density compression scaling law.**

# The stagnated pressure scales to the 4<sup>th</sup> power of the maximum Mach number



- Conservation of energy at stagnation:

$$p_{st} R_{st}^3 \sim m u_{max}^2 \quad R_{st} \sim R_3 \sim \Delta_3 \sim \Delta_2 \quad \Rightarrow \quad p_{st} \Delta_2^3 \sim \underline{m u_{max}^2} \sim \underline{\rho_2 R_2^2 \Delta_2 u_{max}^2}$$

$$\Rightarrow p_{st} \sim \rho_2 \left( \frac{R_2}{\Delta_2} \right)^2 u_{max}^2 = \rho_2 A_2^2 u_{max}^2 \sim p_2 Mach_2^2 \frac{u_{max}^2}{p_2/\rho_2} \sim p_A Mach_2^4 \sim p_A Mach_{max}^4$$

$$\boxed{\frac{p_{st}}{p_A} \sim Mach_{max}^4}$$

$$Mach_2 = Mach_{max} \simeq A_2 = \sqrt{A_{sb}}$$

$$\alpha_{st} \sim \frac{p_{st}}{\rho_{st}^{5/3}} \sim \frac{p_A Mach_{max}^4}{\rho_{sb}^{5/3} Mach_{max}^{10/3}} = \alpha_{sb} Mach_{max}^{2/3}$$

$$\boxed{\frac{\alpha_{st}}{\alpha_{sb}} \sim Mach_{max}^{2/3}}$$

$$\frac{\rho_{st}}{\rho_{sb}} \sim Mach_{max}^2$$

# Scaling of the areal density of the compressed core



$$\rho_{\text{st}} R_{\text{st}} \sim \rho_{\text{st}} A_2 \sim \left( \frac{p_{\text{st}}}{\alpha_{\text{st}}} \right)^{3/5} \frac{A_2}{R_2} \frac{R_2}{R_1} R_1 \sim \left( \frac{p_A \text{Mach}_{\max}^4}{\alpha_{\text{sb}} \text{Mach}_{\max}^{2/3}} \right)^{3/5} \frac{1}{A_2} \frac{1}{A_{\text{sb}}^{1/6}} R_1$$

$$A_2 \sim \text{Mach}_{\max} \quad A_{\text{sb}} \sim \text{Mach}_{\max}^2 \quad \frac{R_2}{R_1} \sim \frac{1}{A_{\text{sb}}^{1/6}}$$

$$\rho_{\text{st}} R_{\text{st}} \sim \left( \frac{p_A}{\alpha_{\text{sb}}} \right)^{3/5} \text{Mach}_{\max}^2 \frac{1}{\text{Mach}_{\max}} \frac{1}{\text{Mach}_{\max}^{1/3}} R_1$$

$$\sim \left( \frac{p_A}{\alpha_{\text{sb}}} \right)^{3/5} \text{Mach}_{\max}^{2/3} R_1 \sim \left( \frac{p_A}{\alpha_{\text{sb}}} \right)^{3/5} \frac{\text{u}_{\max}^{2/3}}{(p_A/\rho_{\text{sb}})^{1/3}} \frac{p_A^{1/3} R_1}{p_A^{1/3}}$$

$$\sim \left( \frac{p_A}{\alpha_{\text{sb}}} \right)^{3/5} \frac{\text{u}_{\max}^{2/3}}{(p_A^{2/5} \alpha_{\text{sb}}^{3/5})^{1/3}} \frac{(p_A R_1^3)^{1/3}}{p_A^{1/3}} \sim \frac{p_A^{2/15}}{\alpha_{\text{sb}}^{4/5}} \text{u}_{\max}^{2/3} E_k^{1/3}$$

$E_k \sim E_{\text{las}} \Rightarrow$

$$\boxed{\rho_{\text{st}} R_{\text{st}} \sim \frac{p_A^{2/15} \text{u}_{\max}^{2/3} E_{\text{las}}^{1/3}}{\alpha_{\text{sb}}^{4/5}}}$$

$$E_k \sim p_A R_1^3$$

# Amplification of areal density

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$$\rho_{\text{st}} R_{\text{st}} \sim \rho_{\text{st}}^{2/3} (\rho_{\text{st}} R_{\text{st}}^3)^{1/3} \sim \rho_{\text{sb}}^{2/3} Mach_{\max}^{4/3} \frac{Mass}{\rho_{\text{sb}}}$$

$$\sim \frac{\rho_{\text{sb}}^{2/3}}{\rho_1^{2/3}} Mach_{\max}^{4/3} \rho_1^{2/3} (\rho_1 R_1^2 \Delta_1)^{1/3}$$

$$\rho_{\text{st}} R_{\text{st}} \sim (\rho_1 \Delta_1) Mach_{\max}^{4/3} A_1^{2/3} \left( \frac{\rho_{\text{sb}}}{\rho_1} \right)^{2/3}$$

$$\frac{\rho_{\text{st}}}{\rho_{\text{sb}}} \sim Mach_{\max}^2$$

$$A_1 = \frac{R_1}{\Delta_1}$$

$$\frac{\rho_{\text{sb}}}{\rho_1} = 4 \left( \frac{I_{\max}}{I_{\text{foot}}} \right)^{2/5}$$

$$(\rho R)_{\text{st}} \sim (\rho_1 \Delta_1) IFAR^{2/3} A_1^{2/3} \left( \frac{I_{\max}}{I_{\text{foot}}} \right)^{4/15}$$

$$IFAR \sim Mach_{\max}^2$$

$$E_{\text{las}} = 4\pi R_1^2 I_{\max} t_{\text{imp}} \approx 4\pi R_1^2 I_{\max} \frac{R_1}{u_{\max}}$$

$$E_{\text{las}} \approx \frac{4\pi R_1^3 I_{\max}}{u_{\max}}$$

# Summary

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$$A_{\text{sb}} = IFAR = 4A_1 \left( \frac{I_{\text{max}}}{I_{\text{foot}}} \right)^{2/5}$$

$$u_{\text{max,cm/s}} \approx 10^7 \sqrt{0.7 A_1 \alpha^{3/5} I_{15,\text{max}}^{4/15} \left( \frac{I_{\text{max}}}{I_{\text{foot}}} \right)^{2/5}}$$

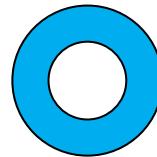
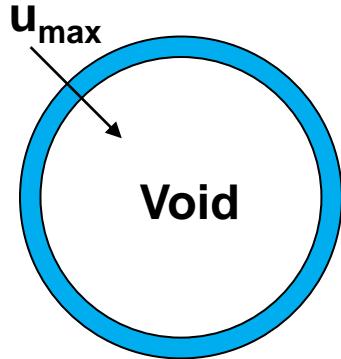
$$\rho_{\text{st}} \sim \rho_{\text{sb}} Mach_{\text{max}}^2 \sim \rho_1 IFAR \left( \frac{I_{\text{max}}}{I_{\text{foot}}} \right)^{2/5}$$

$$p_{\text{st}} \sim p_A Mach_{\text{max}}^4 \sim p_A IFAR^2$$

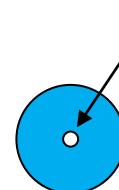
$$\alpha_{\text{st}} \sim \alpha_{\text{sb}} Mach_{\text{max}}^{2/3} \sim \alpha_{\text{sb}} IFAR^{1/3}$$

$$(\rho R)_{\text{st}} \sim (\rho_1 \Delta_1) IFAR^{2/3} A_1^{2/3} \left( \frac{I_{\text{max}}}{I_{\text{foot}}} \right)^{4/15}$$

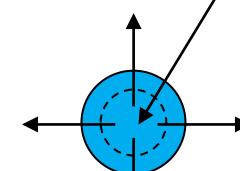
# Calculation of the burn-up fraction



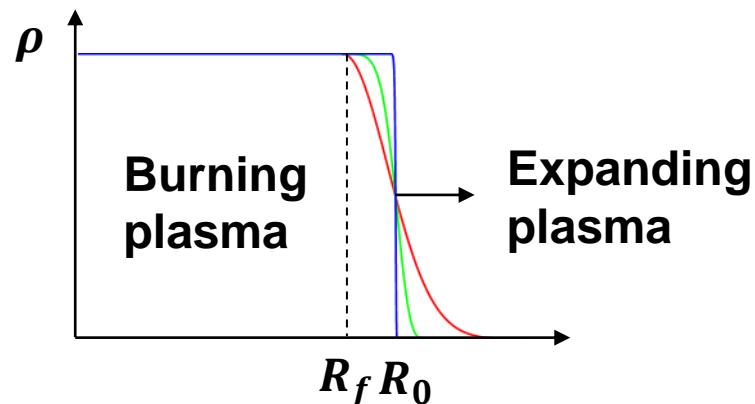
**Void closure**



**Stagnation core**



**Return shock**



$$R_f = R_0 - C_s t$$

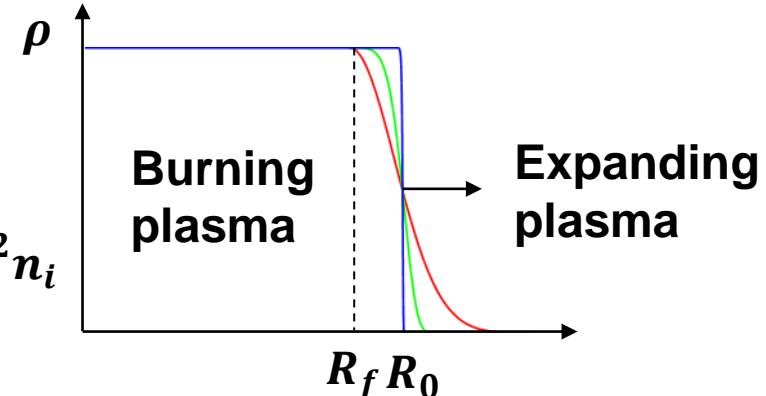
$$\frac{\partial n_i}{\partial t} = -\nabla \cdot (n_i \mathbf{v}) - \frac{n_i^2}{4} \langle \sigma \mathbf{v} \rangle \times 2$$

# Calculation of the burn-up fraction - continue



$$4\pi \int_0^{R_f} r^2 dr \left( \frac{\partial n_i}{\partial t} = -\nabla \cdot (n_i v) - \frac{n_i^2}{2} \langle \sigma v \rangle \right)$$

$$\begin{aligned} 4\pi \int_0^{R_f} r^2 \frac{\partial n_i}{\partial t} dr &= 4\pi \frac{d}{dt} \int_0^{R_f} r^2 n_i dr - 4\pi \dot{R}_f R_f^2 n_i \\ &= -n_i v 4\pi R_f^2 - \frac{n_i^2}{2} \langle \sigma v \rangle V_f \\ &\quad (\text{neglect}) \end{aligned}$$



$$\frac{d}{dt} \left( \int_{a(t)}^{b(t)} f(x, t) dx \right) = f(b(t), t) \frac{db(t)}{dt} - f(a(t), t) \frac{da(t)}{dt} + \int \frac{\partial}{\partial t} f(x, t) dx$$

(Leibniz integral rule)

$$N_f \equiv \frac{4\pi}{3} R_f^3 n_i \equiv V_f n_i$$

$$\frac{dN_f}{dt} - 3N_f \frac{\dot{R}_f}{R_f} = - \frac{N_f^2}{V_f} \frac{\langle \sigma v \rangle}{2}$$

$$\frac{d_t N_f}{N_f^2} - \frac{3\dot{R}_f}{N_f R_f} = - \frac{\langle \sigma v \rangle}{2V_f}$$

$$\frac{d}{dt} \left( \frac{1}{N_f} \right) + \frac{3\dot{R}_f}{N_f R_f} = \frac{\langle \sigma v \rangle}{2V_f}$$

$$R_f^3 \frac{d}{dt} \left( \frac{1}{N_f} \right) + 3R_f^2 \frac{\dot{R}_f}{N_f} = \frac{d}{dt} \left( \frac{R_f^3}{N_f} \right) = \frac{\langle \sigma v \rangle}{2V_f} R_f^3$$

# Calculation of the burn-up fraction - continue



$$\frac{d}{dt} \left( \frac{R_f^3}{N_f} \right) = \frac{\langle \sigma v \rangle}{2V_f} R_f^3 \quad \frac{R_f^3}{N_f} = \int_0^t \frac{\langle \sigma v \rangle}{2V_f} R_f^3 dt + \frac{R_0^3}{N_0}$$

$$R_f = R_0 - C_s t \quad dt = -\frac{dR_f}{C_s} \quad V_f = \frac{4\pi}{3} R_f^3$$

$$\frac{R_f^3}{N_f} = - \int_{R_0}^{R_f} \frac{\langle \sigma v \rangle}{2 \times 4\pi/3} \frac{dR_f}{C_s} + \frac{R_0^3}{N_0}$$

$$\frac{R_f^3}{N_f} = - \int_{R_0}^{R_f} \frac{\langle \sigma v \rangle}{2C_s} \frac{3}{4\pi} dR_f + \frac{R_0^3}{N_0}$$

$$\frac{R_f^3}{N_f} = \frac{\langle \sigma v \rangle}{2C_s} \frac{3}{4\pi} (R_0 - R_f) + \frac{R_0^3}{N_0}$$

$$\frac{V_f}{N_f} = \frac{\langle \sigma v \rangle}{2C_s} R_0 \left( 1 - \frac{R_f}{R_0} \right) + \frac{V_0}{N_0}$$

$$n_0 = \frac{N_0}{V_0}$$

$$\frac{V_f}{N_f} = \frac{V_0}{N_0} \left[ 1 + \frac{\langle \sigma v \rangle}{2C_s} n_0 R_0 \left( 1 - \frac{R_f}{R_0} \right) \right]$$

$$\xi \equiv \frac{\langle \sigma v \rangle}{2C_s} n_0 R_0$$

$$\frac{V_f}{N_f} = \frac{1}{n_0} \left[ 1 + \xi \left( 1 - \frac{R_f}{R_0} \right) \right]$$

# Calculation of the burn-up fraction - continue



$$\frac{V_f}{N_f} = \frac{1}{n_0} \left[ 1 + \xi \left( 1 - \frac{R_f}{R_0} \right) \right]$$

$$n_i = \frac{N_f}{V_f}$$

$$\begin{aligned}
 \text{\#Burned ions} &= \int_0^t \frac{\langle \sigma v \rangle}{2} n_i^2 V_f dt = \int_0^t \frac{\langle \sigma v \rangle}{2} \frac{N_f^2}{V_f} dt = - \int_{R_0}^{R_f} \frac{\langle \sigma v \rangle}{2} \left( \frac{N_f}{V_f} \right)^2 V_f \frac{dR_f}{C_s} \\
 &= \int_{R_f}^{R_0} \frac{\langle \sigma v \rangle}{2} \frac{n_0^2}{\left[ 1 + \xi \left( 1 - \frac{R_f}{R_0} \right) \right]^2} \left( \frac{R_f}{R_0} \right)^3 V_0 R_0 \frac{dR_f/R_0}{C_s} \quad V_{f,0} = \frac{4\pi}{3} R_{f,0}^3 \\
 &= \int \frac{\langle \sigma v \rangle}{2} \frac{n_0^2}{[1 + \xi(1-x)]^2} x^3 V_0 \frac{R_0}{C_s} dx = N_0 \xi \int_0^1 \frac{x^3 dx}{[1 + \xi(1-x)]^2} \\
 &= N_0 \frac{\xi[6 + \xi(9 + 2\xi)] - 6(1 + \xi)^2 \ln[1 + \xi]}{2\xi^3} \quad \xi \equiv \frac{\langle \sigma v \rangle}{2C_s} n_0 R_0
 \end{aligned}$$

#Burn-up Fraction

$$\theta(\xi) = \frac{\xi[6 + \xi(9 + 2\xi)] - 6(1 + \xi)^2 \ln[1 + \xi]}{2\xi^3}$$

# Calculation of the burn-up fraction - continue

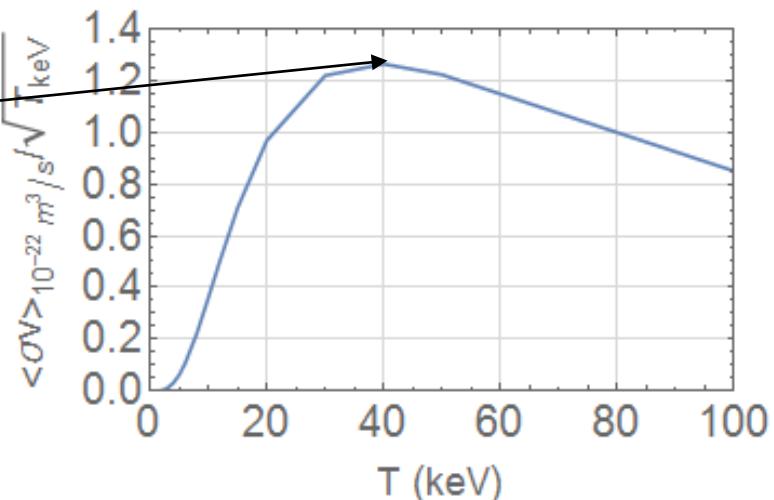


$$C_s = \sqrt{\frac{T_e + T_i}{m_i}} = \sqrt{\frac{2T}{m_i}} \quad \rho = n_0 m_i \quad m_i = \frac{m_D + m_T}{2} = 2.5 \times 1.67 \times 10^{-27} \text{ kg}$$

$$\xi = \frac{\langle \sigma v \rangle}{\sqrt{T}} (\rho R_0) \frac{1}{2\sqrt{2m_i}} = \frac{\langle \sigma v \rangle_{m^3/s}}{\sqrt{T_{keV} \times 1.6 \times 10^{-16}}} \frac{(\rho R_0)_{g/cm^2} \times 10}{2\sqrt{5 \times 1.67 \times 10^{-27}}}$$

$$\xi \approx \frac{1.25 \times 10^{-22}}{\sqrt{1.6 \times 10^{-16}}} \frac{10(\rho R_0)_{g/cm^2}}{2\sqrt{5 \times 1.67 \times 10^{-27}}} = 0.54(\rho R_0)_{g/cm^2}$$

$$\left. \frac{\langle \sigma v \rangle}{\sqrt{T_{keV}}} \right|_{\max} = 1.25 \times 10^{-22} \quad @ \quad T = 40 \text{ keV}$$



# Smallest areal density ( $\rho R$ )



#Burned-up Fraction

$$\theta(\xi) = \frac{\xi[6 + \xi(9 + 2\xi)] - 6(1 + \xi)^2 \ln[1 + \xi]}{2\xi^3}$$

$$\lim_{\xi \rightarrow 0} \theta(\xi) = \frac{\xi}{4}$$

$$\lim_{\xi \rightarrow \infty} \theta(\xi) = 1$$

$$\theta(\xi) \approx \frac{\xi}{4 + \xi}$$

$$\xi \simeq 0.54(\rho R_0)_{g/cm^2}$$

$$\theta(\xi) \approx \frac{0.54\rho R}{4 + 0.54\rho R}$$

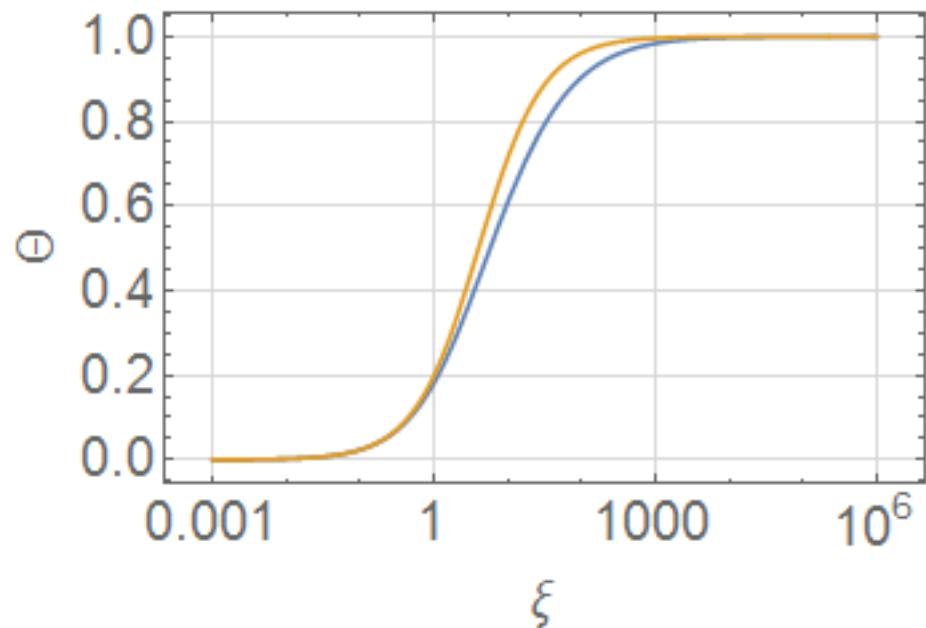
$$\theta(\xi) \approx \frac{(\rho R)_{g/cm^2}}{7 + (\rho R)_{g/cm^2}}$$

Large  $\rho R$  is needed to have high burn-up fraction.

For energy applications:

$$\theta \gtrsim 0.3$$

$$\rho R \geq 3 g/cm^2$$



# Energy gain



$$\text{Fusion energy} = \frac{M_0}{2m_i} \epsilon_f \Theta$$

$$\epsilon_f = 17.6 \text{ MeV}$$

$$\text{Energy gain} = \frac{\text{Fusion Energy}}{\text{Input Energy}}$$

Mass =  $M_0$   
Temp =  $T$   
DT  
Volume =  $V_0$

- Input energy: the sphere is heated to the temperature  $T$

$$\text{Thermal energy in sphere: } \frac{3}{2} (n_{i0} T_i + n_{e0} T_e) V_0$$

$$n_{i0} = n_{e0} \equiv n_0 \quad T_e = T_i \Rightarrow 3n_0 TV_0 = 3 \frac{M_0}{m_i} T$$

$$\text{Set heating efficiency: } \eta = \frac{\text{Thermal Energy}}{\text{Input Energy}}$$

$$\text{Gain} = \frac{\frac{M_0}{2m_i} \epsilon_f \Theta}{3 \frac{M_0}{m_i} T / \eta} = \eta \frac{M_0}{2m_i} \frac{\epsilon_f \Theta}{3 \frac{M_0}{m_i} T} = \frac{\eta}{6} \frac{\epsilon_f}{T} \Theta$$

$$\boxed{\text{Gain} = \eta 293 \left( \frac{10}{T_{\text{keV}}} \right) \Theta}$$

# The power to heat the plasma is enormous



- Consider the small T limit:

$$\theta(\xi) \approx \frac{\xi}{4 + \xi} \quad \xi \equiv \frac{\langle \sigma v \rangle}{2C_s} n_0 R_0 = \frac{\langle \sigma v \rangle}{\sqrt{T}} (\rho R_0) \frac{1}{2\sqrt{2m_i}}$$

$\langle \sigma v \rangle \sim T^4$  for  $T \rightarrow 0$ , then  $\xi \sim T^{7/2}$  and  $Gain \sim T^{5/2} \rightarrow 0$

- Required input power:

$$P_w = \frac{E_{\text{input}}}{\tau_{\text{input}}} \quad \tau_{\text{input}} \ll \tau_{\text{burn}} = \frac{R}{C_s} \quad (\text{Heat out before it runs away})$$

$$P_w = \frac{E_{\text{input}}}{\mu R/C_s} = \frac{E_{\text{thermal}}}{\eta \mu R/C_s} = 3 \frac{M_0}{m_i} T \frac{1}{R} \frac{C_s}{\eta \mu} \quad \tau_{\text{input}} = \mu \frac{R}{C_s} \quad \text{Ex: } \mu \sim 0.1$$

$$\frac{P_w}{M_0} = \frac{3}{m_i} \frac{T}{R} \frac{C_s}{\eta \mu} = \frac{3}{m_i} \frac{T}{R} \sqrt{\frac{2T}{m_i}} \frac{1}{\eta \mu}$$

$$\boxed{\frac{P_w}{M_0} = 10^{18} \left( \frac{T_{\text{keV}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{1}{R_{\text{cm}}} \frac{1}{\eta} \text{ Watts/g}}$$

# A clever way is needed to ignite a target



- For  $T = 10 \text{ keV}$

$$\xi \approx 0.18(\rho R)$$

$$Gain|_{10\text{keV}} \approx 293\eta \frac{0.18\rho R}{4 + 0.18\rho R} \approx 293\eta \frac{\rho R_{g/\text{cm}^2}}{22 + \rho R_{g/\text{cm}^2}}$$

- For  $T=40 \text{ keV}$

$$\xi \approx 0.54(\rho R)$$

$$Gain|_{40\text{keV}} \approx 73\eta \frac{\rho R_{g/\text{cm}^2}}{7 + \rho R_{g/\text{cm}^2}}$$

- For Gains  $\gtrsim 100$

- $T = 10 \text{ keV}$

$$\rho R_{g/\text{cm}^2} \gtrsim \frac{22}{2.93\eta - 1}$$

$$\rho R \gtrsim 22 \text{ g/cm}^2$$

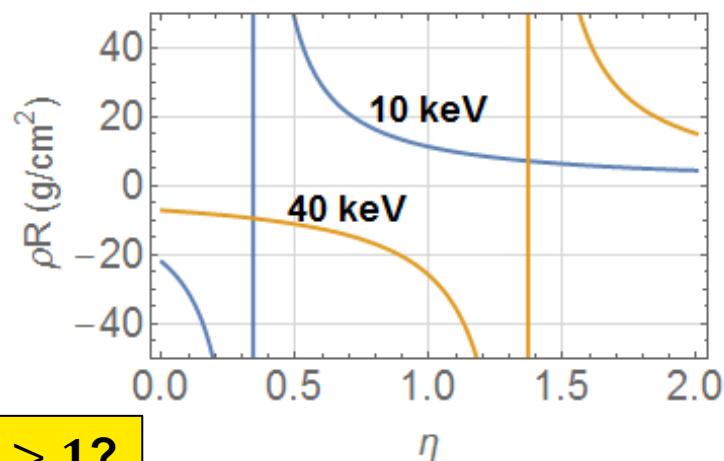
$$\eta > 1$$

- $T = 40 \text{ keV}$

$$\rho R_{g/\text{cm}^2} \gtrsim \frac{7}{0.73\eta - 1}$$

$$\rho R \gtrsim 7 \text{ g/cm}^2$$

$$\eta > 1$$



How do we get  $\eta > 1$ ?

# Requirement to ignite a target



- For  $T=10 \text{ keV}$  and  $\rho R \gtrsim 22 \text{ g/cm}^2$

$$\rho R = \frac{4\pi}{3} \frac{\rho R^3}{4\pi R^2/3} = \frac{M_0}{\frac{4\pi}{3} R^2} = \frac{3}{4\pi} \frac{M_0}{R^2} \gtrsim 22 \text{ g/cm}^2$$

$$\frac{M_0}{R^2} \gtrsim 92 \text{ g/cm}^2$$

$$P_w \Big|_{10keV} = 10^{18} \left( \frac{T_{\text{keV}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{M_0}{R_{\text{cm}}} \frac{1}{\eta} = 10^{18} \frac{0.1}{\mu} \frac{1}{\eta} 92 R_{\text{cm}} \quad \text{Watts}$$

$$P_w \Big|_{10keV} \approx 10^{20} \frac{0.1}{\mu} \frac{R_{\text{cm}}}{\eta} \text{ Watts}$$

- For  $T=40 \text{ keV}$

$$\rho R \gtrsim 7 \implies \frac{M_0}{R^2} \gtrsim 30 \text{ g/cm}^2$$

$$P_w \Big|_{40keV} \approx 2.4 \times 10^{20} \frac{0.1}{\mu} \frac{R_{\text{cm}}}{\eta} \text{ Watts}$$

- Needed:

$$R_{\text{cm}} \ll 1$$

$$\eta \gg 1$$

$$\mu \gg 0.1$$

# Requirements to ignite a target

---



$$P_w \Big|_{10keV} \approx 10^{20} \frac{0.1}{\mu} \frac{R_{cm}}{\eta} \text{ Watts}$$

- $R_{cm} \ll 1$  : sphere size in the order of 100's um
- $\eta \gg 1$  : input energy amplification
- $\mu \gg 0.1$  : energy delivery time decoupled from burn time. Need longer energy delivery time. Need to bring down power to  $\sim 10^{15}$  W

# Math....#!@%\$#\$#&^%\$#

---



$$P_w = 10^{18} \frac{M_{0,g}}{\eta} \left( \frac{T_{\text{keV}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{1}{R_{\text{cm}}} \text{ Watts/g}$$

$$\tau_{\text{input}} = \mu \frac{R}{C_s} \quad \text{Ex: } \mu \sim 0.1 \quad \eta = \frac{\text{Thermal Energy}}{\text{Input Energy}}$$

$$\text{Gain} = 293\eta \left( \frac{10}{T_{\text{keV}}} \right) \theta(\xi) \quad \theta(\xi) \approx \frac{\xi}{4 + \xi} \quad \xi = \frac{\langle \sigma v \rangle}{2m_i C_s} (\rho R_0)$$

$$G_{\max} \equiv 293\eta \left( \frac{10}{T_{\text{keV}}} \right) \quad G = G_{\max} \frac{\xi}{4 + \xi} \Rightarrow \xi = \frac{4G}{G_{\max} - G}$$

$$P_w = \frac{10^{18}}{\eta} \left( \frac{T_{\text{kev}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{4\pi}{3} \frac{\rho R_0^3}{R_0} = \frac{10^{18}}{\eta} \left( \frac{T_{\text{kev}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{4\pi}{3} (\rho R_0) R_0$$

# More math...!#\$%%^&\*&^(\*&%)(#%!@\$#%^\*&%()



$$P_w = \frac{10^{18}}{\eta} \left( \frac{T_{\text{kev}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{4\pi}{3} \frac{\rho R_0^3}{R_0} = \frac{10^{18}}{\eta} \left( \frac{T_{\text{kev}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{4\pi}{3} (\rho R_0) R_0$$

$$= \frac{10^{18}}{\eta} \left( \frac{T_{\text{kev}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{4\pi}{3} R_0 \frac{2m_i C_s}{\langle \sigma v \rangle} \xi \quad \text{where } \xi \equiv \frac{\langle \sigma v \rangle}{2m_i C_s} (\rho R)$$

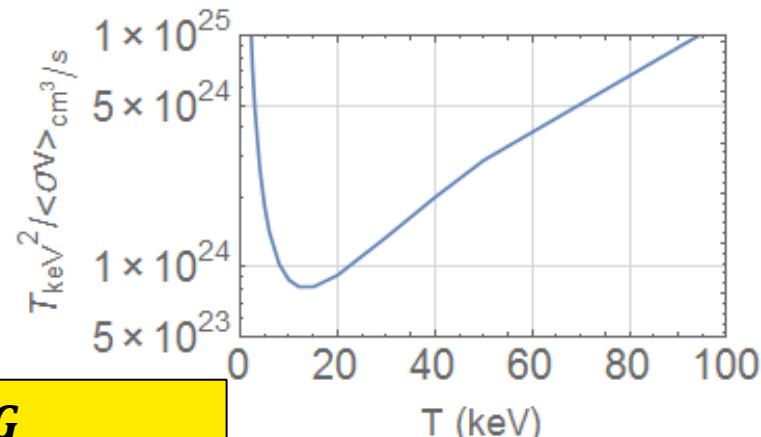
$$= \frac{10^{18}}{\eta} \left( \frac{T_{\text{kev}}}{10} \right)^{3/2} \frac{0.1}{\mu} \frac{32\pi}{3} R_{0,\text{cm}} \frac{\sqrt{T m_i}}{\langle \sigma v \rangle} \frac{G}{G_{\max} - G} \quad \text{where } C_s = \sqrt{\frac{2T}{m_i}}$$

$$P_w = \frac{10^{18}}{\eta} \frac{T_{\text{kev}}^2}{\langle \sigma v \rangle_{\text{cm}^2/\text{s}}} \frac{0.1}{\mu} R_{0,\text{cm}} \frac{G}{G_{\max} - G} \text{ Watts}$$

$$\left. \frac{T_{kev}^2}{\langle \sigma v \rangle_{cm^2/s}} \right|_{min} = 8 \times 10^{23} \quad \text{for } T = 14 \text{ keV}$$

$$\frac{G}{G_{\max} - G} \approx \frac{G}{G_{\max}}$$

$$P_w \approx \frac{7 \times 10^{19}}{\eta} \frac{0.1}{\mu} R_{0,\text{cm}} \frac{G}{G_{\max}} \text{ Watts}$$



# Need to lower the power by 5 orders of magnitude

---



$$P_w \approx \frac{7 \times 10^{19}}{\eta} \frac{0.1}{\mu} R_{0,\text{cm}} \frac{G}{G_{\max}} \text{ Watts}$$

- $\mu \uparrow$  :
- $\eta \uparrow$  : require the fuel ignition from a “spark.” Ignite only a small portion of the DT plasma, i.e.,  $M_h \ll M_0$
- $R_0 \downarrow$  : smaller system size

$$P_w = P_w(M_0) \frac{M_h}{M_0}$$

$$P_w^{\min} = \frac{7 \times 10^{15}}{\eta_h} \left( \frac{M_h/M_0}{0.01} \right) \left( \frac{R_{0,\mu\text{m}}}{100} \right) \left( \frac{0.1}{\mu} \right) \left( \frac{G}{G_{\max}} \right) \text{ Watts}$$



Effective increase in  $\eta$

# Target design using an 1MJ laser

---



$$P_w^{\min} = \frac{7 \times 10^{15}}{\eta_h} \left( \frac{M_h/M_0}{0.01} \right) \left( \frac{R_{0,\mu\text{m}}}{100} \right) \left( \frac{0.1}{\mu} \right) \left( \frac{G}{G_{\max}} \right) \text{Watts}$$

- For the case of using a huge laser, ex: 1MJ.
- The ignition requires temperatures  $T \gtrsim 5\text{keV}$ , then the energy required for ignition is

$$E_{\text{ign}} \approx 3 \frac{M_h}{m_i} \frac{T}{\eta_h}$$

$$M_h \approx \frac{m_i}{3} \frac{\eta_h E_{\text{ign}}}{T}$$

$$M_{h,\mu\text{g}} \approx 17 \left( \frac{5}{T_{\text{keV}}} \right) E_{\text{igm,MJ}} \left( \frac{\eta_h}{0.01} \right) \quad M_h \approx 20\mu\text{g}$$

# Target design using an 1MJ laser - continue



- For “inefficient” heating mechanism ( $\eta_h \approx 1\%$ ), the mass that can be heated to  $T \approx 5$  keV is in the order of  $M_h \approx 20 \mu\text{g}$ .
- If  $M_h/M_0 \approx 0.01$ , then  $M_0 \approx 2 \text{ mg}$ .
- Assuming that the burned-up fraction  $\theta \approx \frac{\rho R}{7 + \rho R}$   
for  $\theta \approx 30\% \rightarrow \rho R \approx 3 \text{ g/cm}^2$

$$M_0 = \frac{4\pi}{3} \rho R^3 = \frac{4\pi}{3} R^2 (\rho R)$$

$$R = \sqrt{\frac{4\pi}{3} \frac{M_0}{\rho R}} = 126 \sqrt{\frac{M_{0,\text{mg}}}{2}} \sqrt{\frac{3}{\rho R}} \mu\text{m}$$

$$\rho = \frac{3M_0}{4\pi R^3} = 240 \sqrt{\frac{M_{0,\text{mg}}}{2}} \left(\frac{126}{R_{\mu\text{m}}}\right)^3 \text{ g/cm}^3 \xleftarrow[\times 1000]{\quad} \rho_{\text{DT}} = 0.25 \text{ g/cm}^3$$

- DT must be compressed  $\sim 1000$  times
- The initial radius of a 2 mg sphere of DT is  $R_{\text{init}} \simeq 2.6 \text{ mm}$  while the final radius  $R_{\text{final}} \simeq 100 \mu\text{m}$ , the convergence ratios of 30 ~ 40 are required.

# Requirements of the density and size of the ignition mass

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$$M_h \approx 20\mu\text{g}$$

$$\rho_h R_h \approx 0.3 \text{ g/cm}^2 \leftarrow \text{To stop 3.5 MeV } \alpha \text{ particles}$$

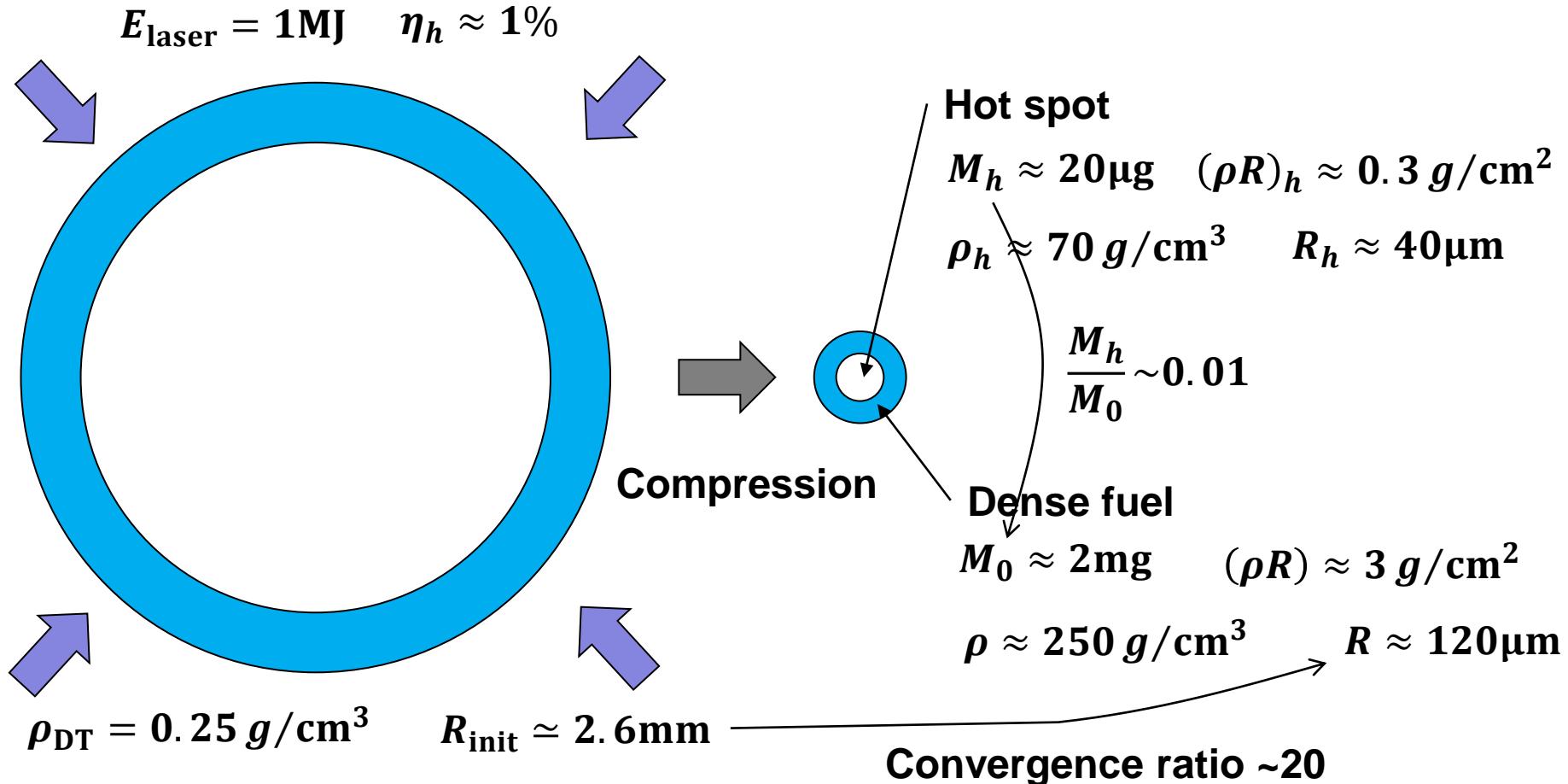
$$R_h \simeq \sqrt{\frac{3}{4\pi} \frac{M_h}{\rho_h R_h}} \approx 40\mu\text{m}$$

$$\rho_h \approx \frac{(\rho_h R_h)}{R_h} = \frac{0.3}{40 * 10^{-4}} = 75 \text{ g/cm}^3$$

# Summary



- Possible fuel assembly for 1MJ ICF driver



# There are alternative

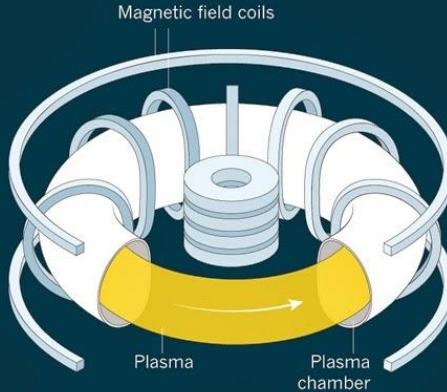


## TRAPPING FUSION FIRE

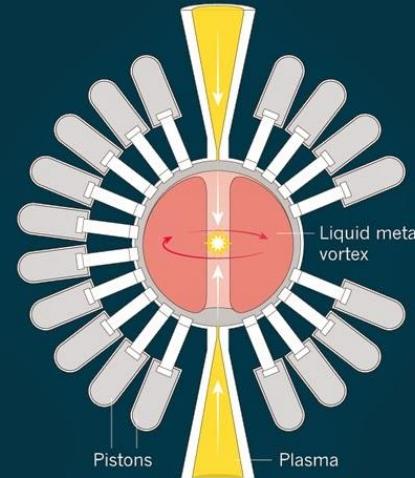
When a superhot, ionized plasma is trapped in a magnetic field, it will fight to escape. Reactors are designed to keep it confined for long enough for the nuclei to fuse and produce energy.

### A CHOICE OF FUELS

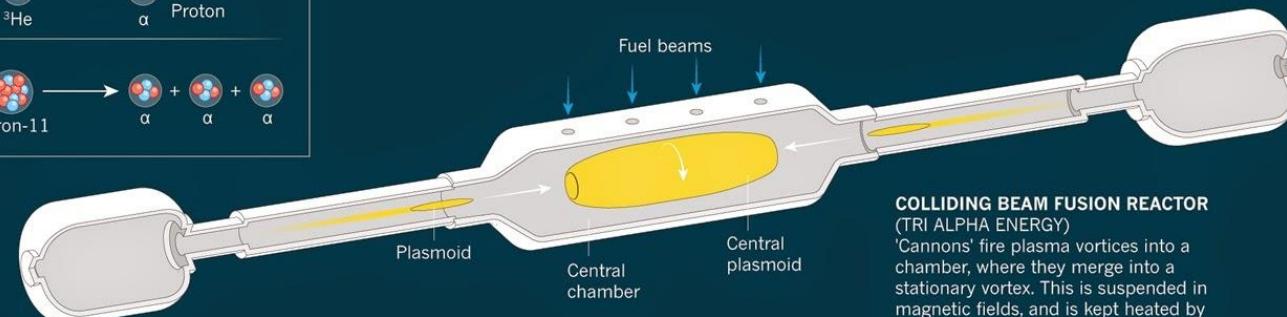
Many light isotopes will fuse to release energy. A deuterium-tritium mix ignites at the lowest temperature, roughly 100 million kelvin, but produces neutrons that make the reactor radioactive. Other fuels avoid that, but ignite at much higher temperatures.



**TOKAMAK**  
(ITER AND MANY OTHERS)  
Multiple coils produce magnetic fields that hold the plasma in the chamber. A coil through the centre drives a current through the plasma to keep it hot.



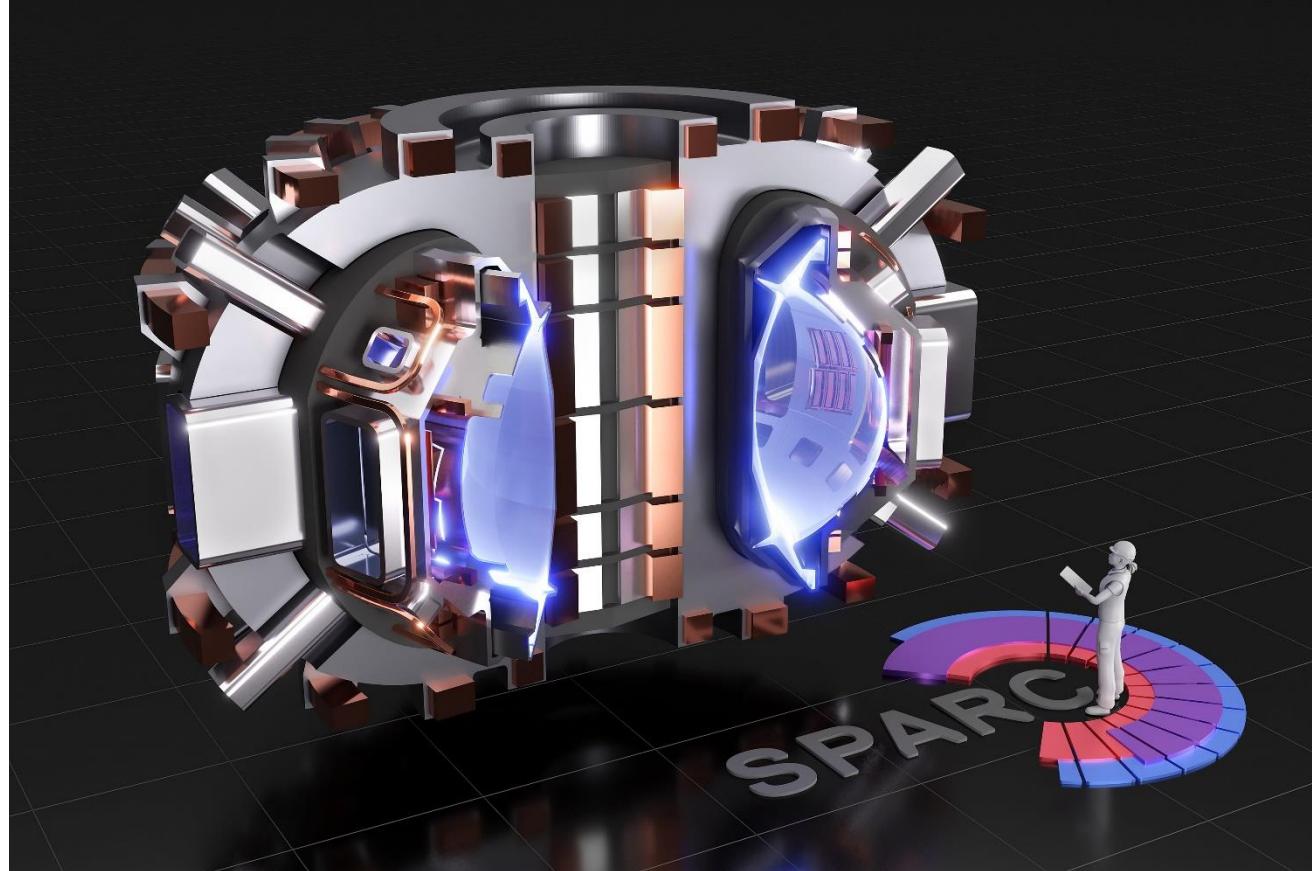
**MAGNETIZED TARGET REACTOR**  
(GENERAL FUSION)  
Magnetized rings of plasma are injected into a vortex of liquid metal. Pistons punch the metal inwards, compressing the plasma to ignite fusion.



**COLLIDING BEAM FUSION REACTOR**  
(TRI ALPHA ENERGY)  
'Cannons' fire plasma vortices into a chamber, where they merge into a stationary vortex. This is suspended in magnetic fields, and is kept heated by beams of fresh fuel.

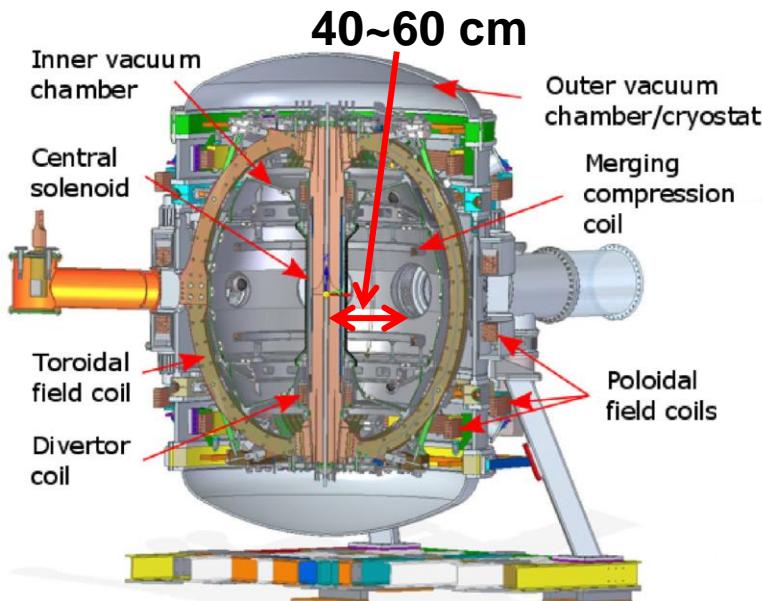
# Commonwealth Fusion Systems, a MIT spin-out company, is building a high-magnetic field tokamak

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- Fusion power  $\propto B^4$ .
- The fusion gain  $Q > 2$  is expected for SPARC tokamak.

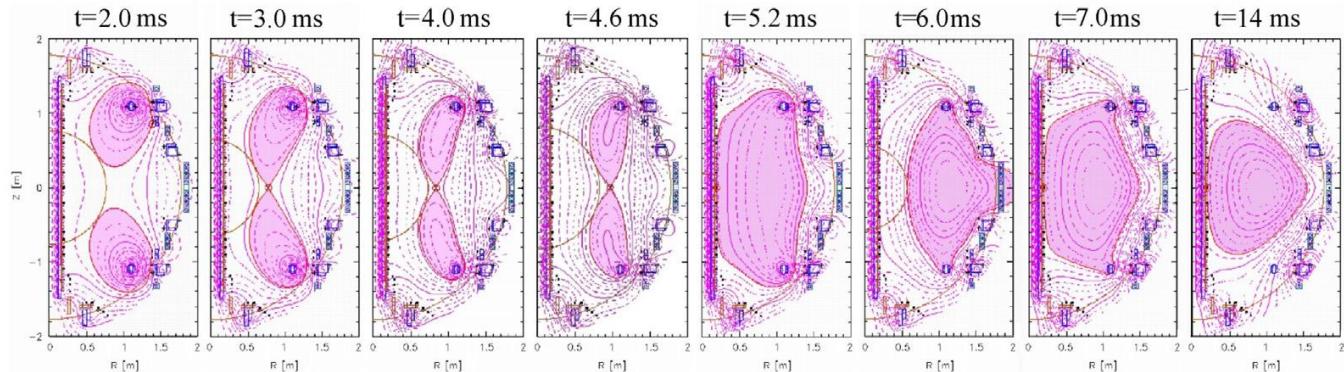
# Merging compression is used to heat the tokamak at the start-up process in ST40 Tokamak at Tokamak Energy Ltd



- High temperature superconductors are used.
- $B_T \sim 3\text{ T}$



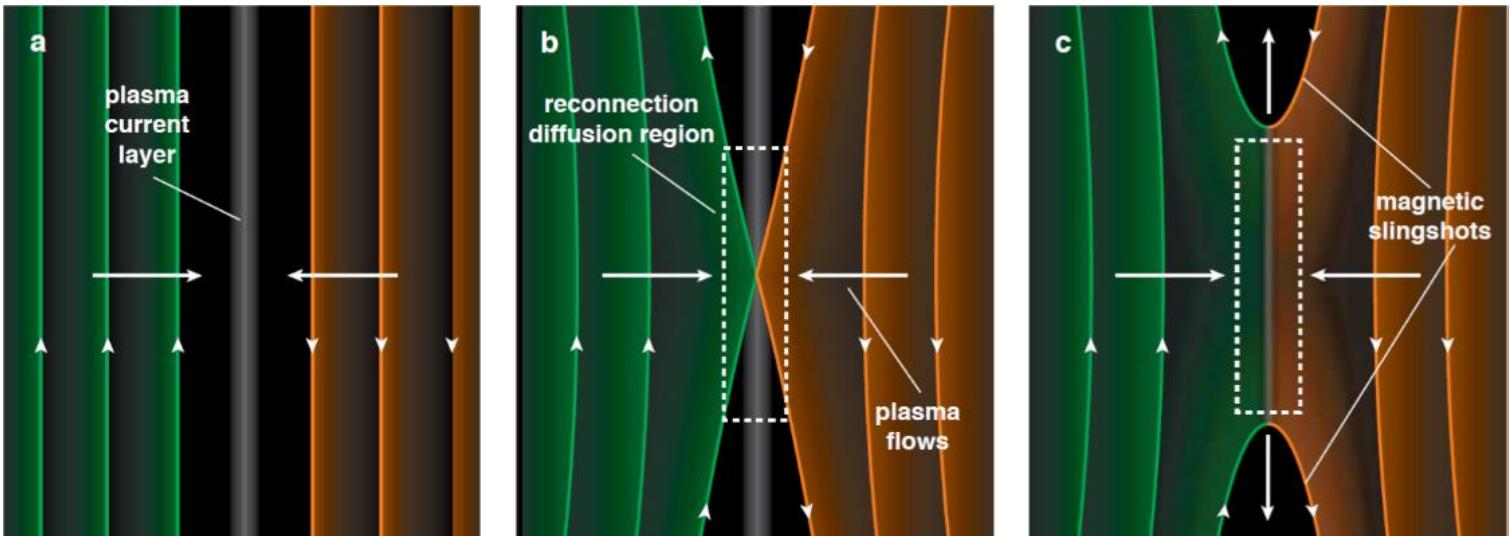
- **Merging compression**



M. Gryaznevich, etc., Fusion Eng. Design, 123,177 (2017)  
<https://www.tokamakenergy.co.uk/>

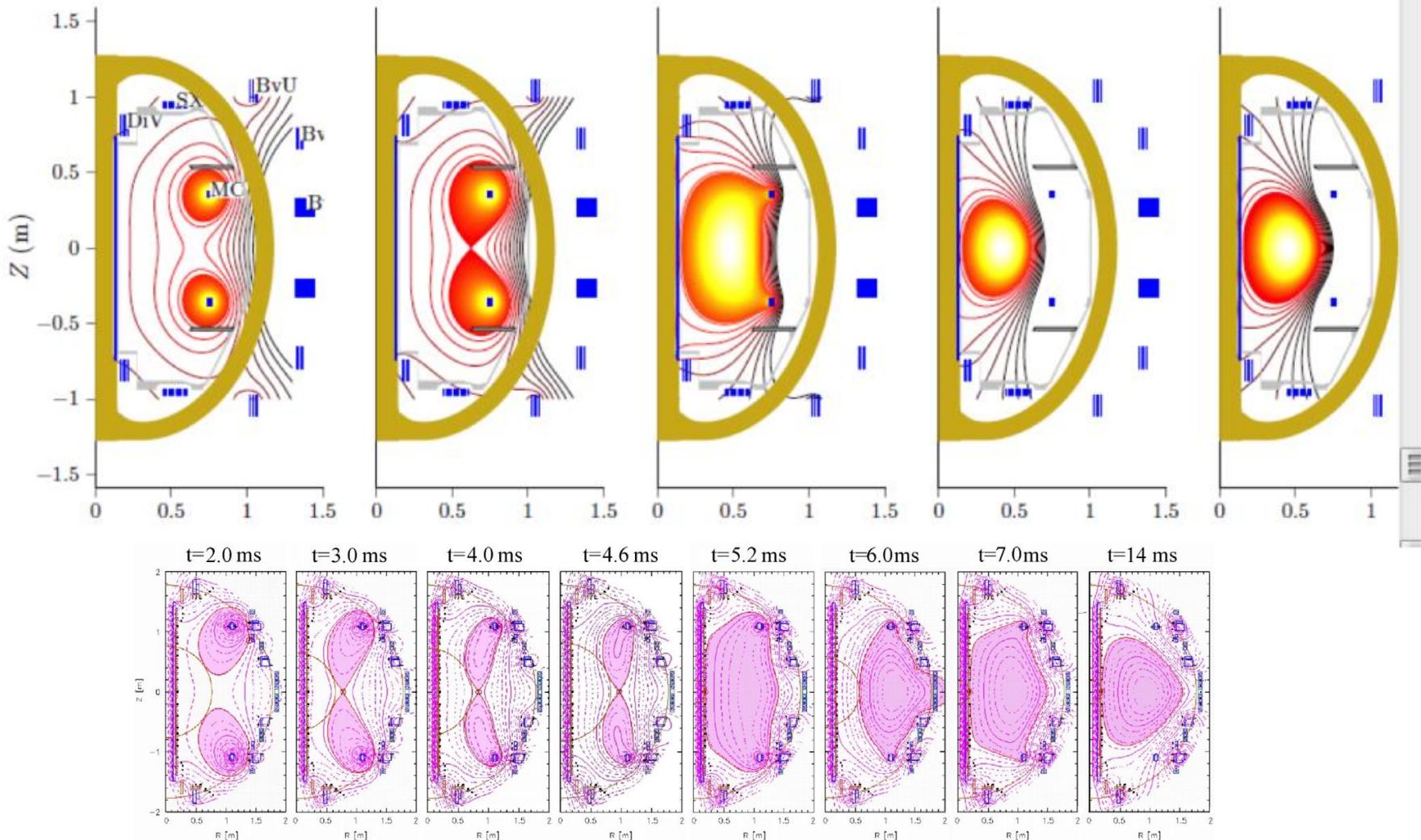
P. F. Buxton, etc., Fusion Eng. Design, 123, 551 (2017)

# Reconnection

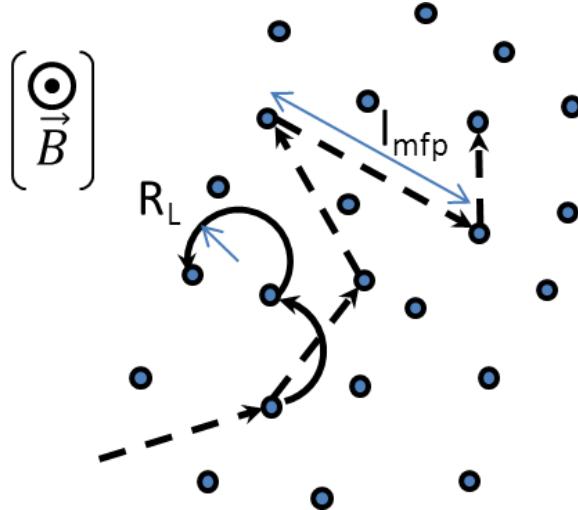


<https://www.youtube.com/watch?v=7sS3Lpzh0Zw>

# Merging compression is used to heat the plasma



# A strong magnetic field reduces the heat flux



$$\mathbf{q}_T = -\kappa_{||} \nabla_{||} T - \kappa_{\perp} \nabla_{\perp} T$$

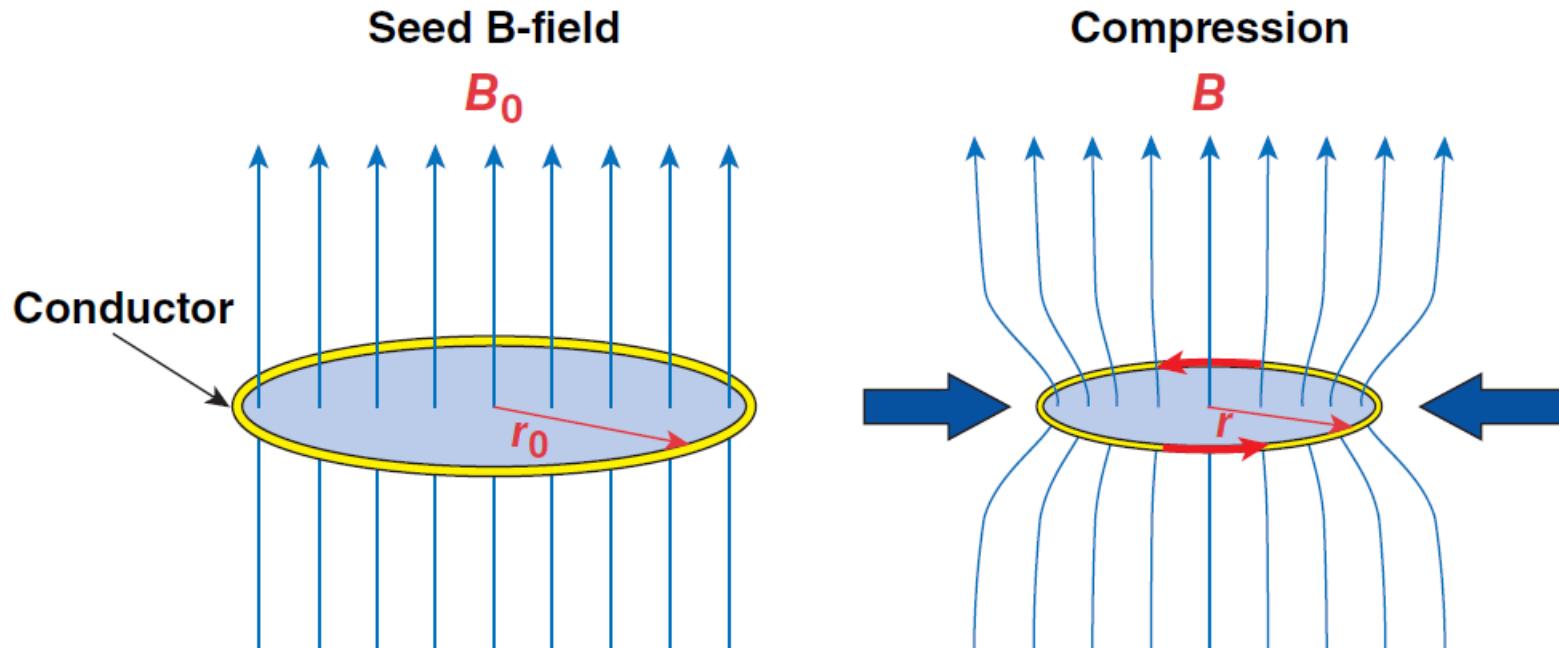
$$\kappa_{||} = \kappa_0 T^{5/2}$$

$$\kappa_{\perp} = \frac{\kappa_{||}}{\chi^2} \quad \text{for large Hall parameter } \chi \propto \frac{l_{\text{mfp}}}{R_L} \gg 1$$

- Typical hot spot conditions:  
 $R_{hs} \sim 40 \mu\text{m}$ ,  $\rho \sim 20 \text{ g/cm}^3$ ,  $T \sim 5 \text{ keV}$ :  
 $B > 10 \text{ MG}$  is needed for  $\chi > 1$

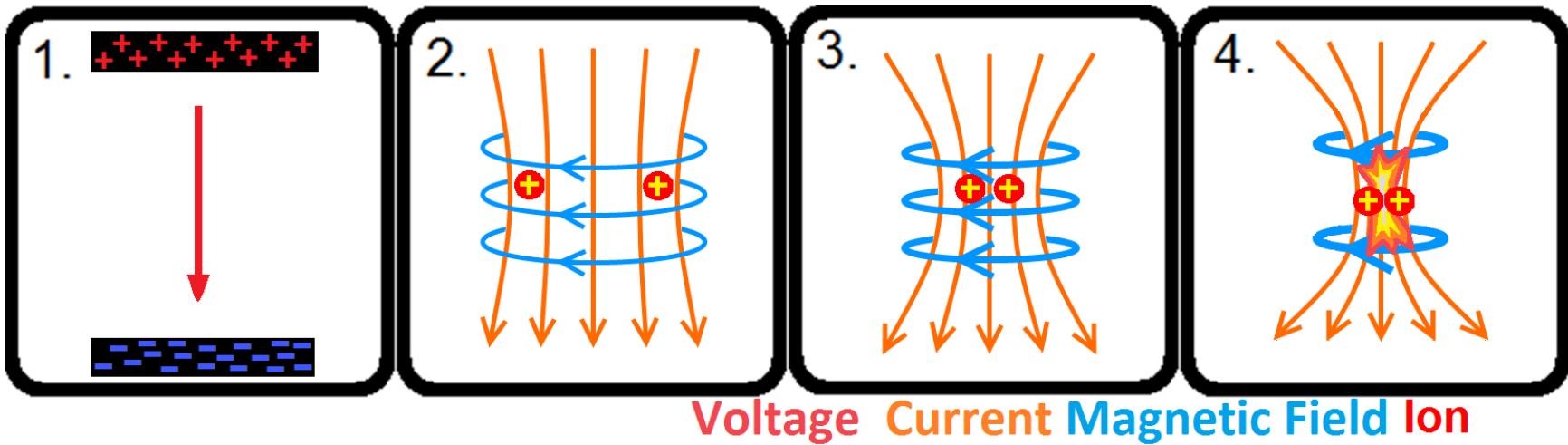
**Magnetic-flux compression can be used to provide the needed magnetic field.**

# Principle of frozen magnetic flux in a good conductor is used to compress fields



$$\Phi = \pi r_0^2 B_0 = \pi r^2 B$$

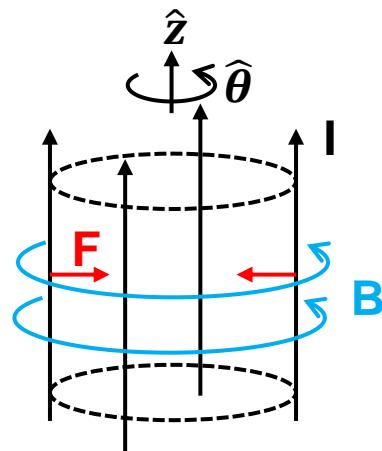
# Plasma can be pinched by parallel propagating plasmas



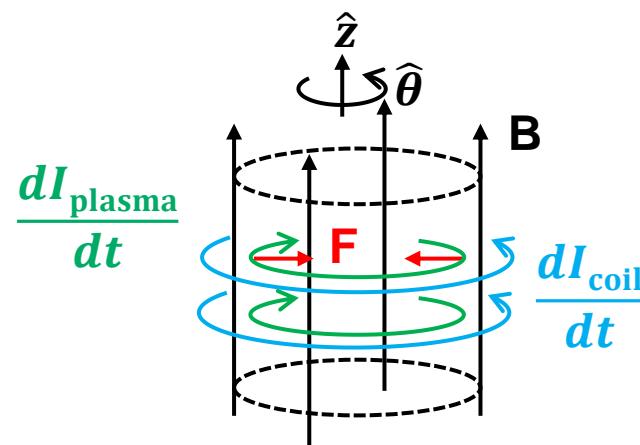
# Plasma can be heated via pinches



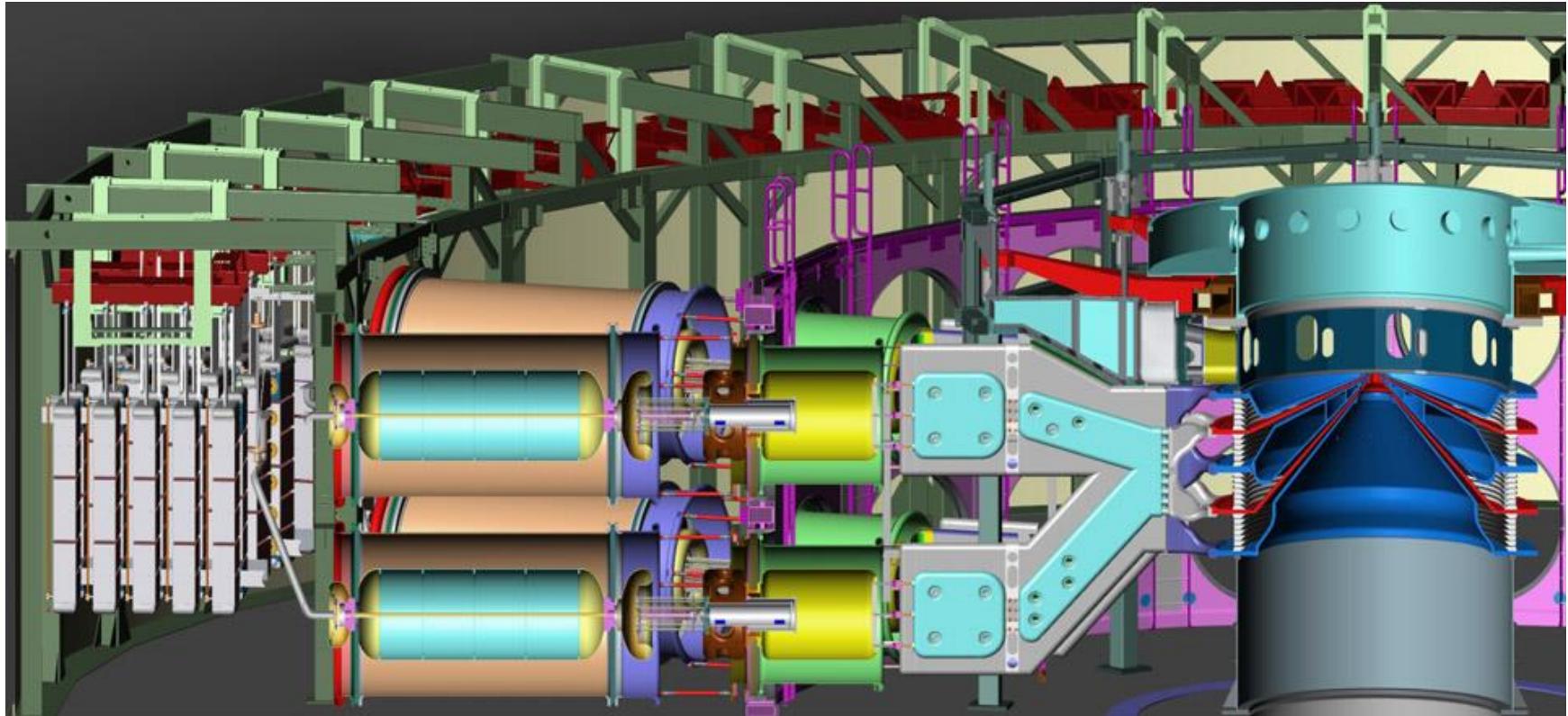
Z pinch



Theta pinch

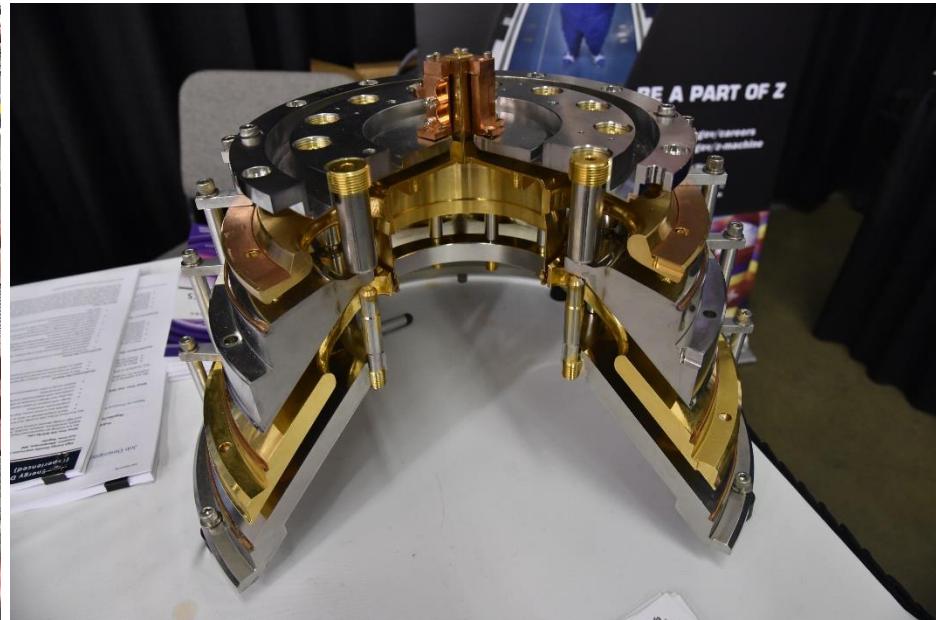


# Sandia's Z machine is the world's most powerful and efficient laboratory radiation source

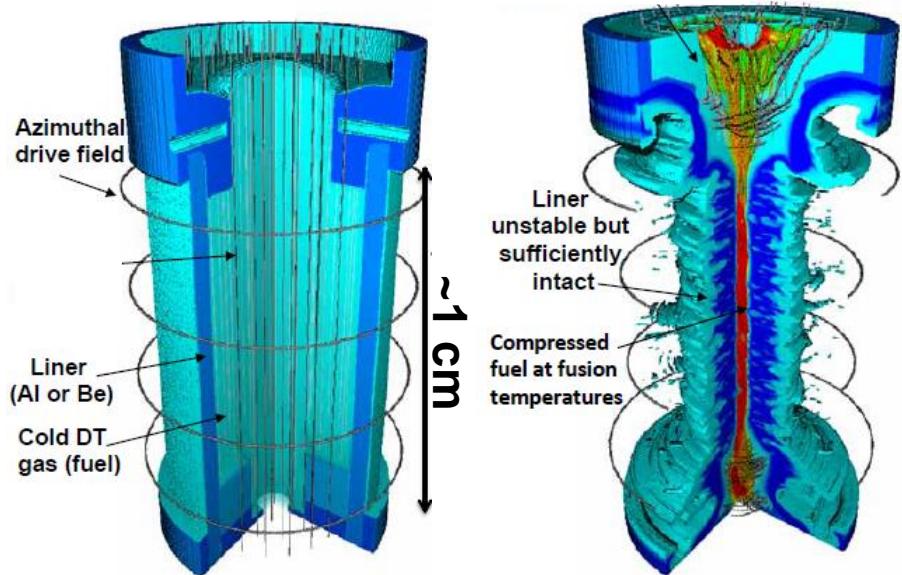


- Stored energy: 20 MJ
- Marx charge voltage: 85 kV
- Peak electrical power: 85 TW
- Peak current: 26 MA
- Rise time: 100 ns
- Peak X-ray emissions: 350 TW
- Peak X-ray output: 2.7 MJ

# Z machine



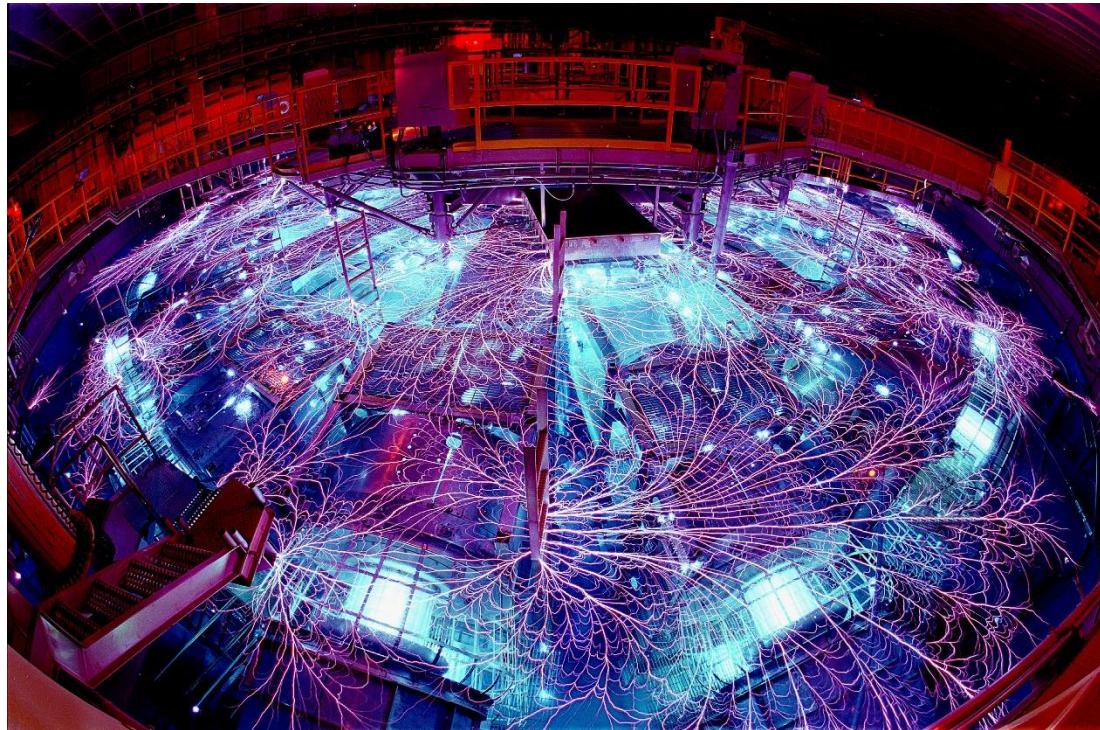
# Z machine



- **Stored energy: 20 MJ**
- **Peak electrical power: 85 TW**
- **Peak current: 26 MA**
- **Rise time: 100 ns**
- **Peak X-ray output: 2.7 MJ**

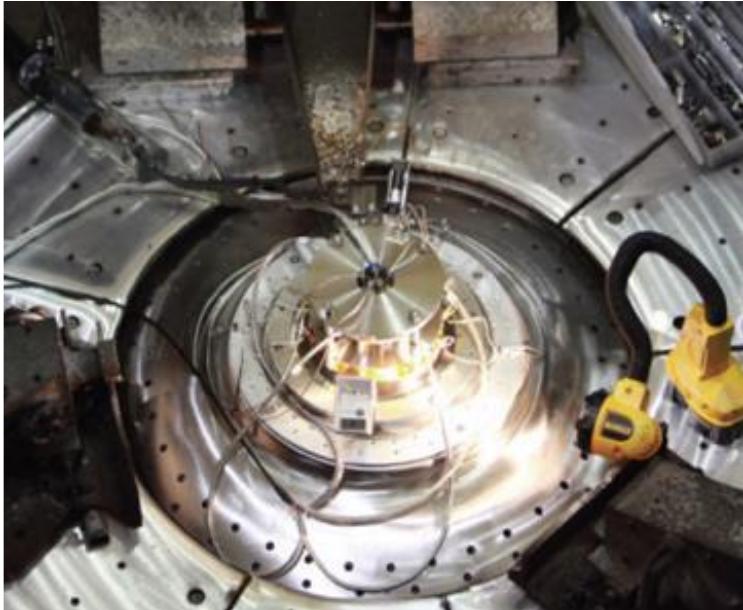
# Z machine discharge

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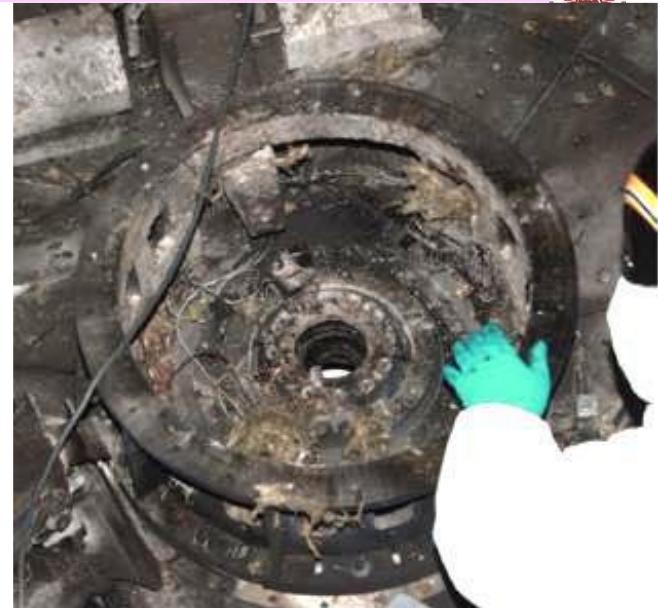


# Before and after shots

- Before shots

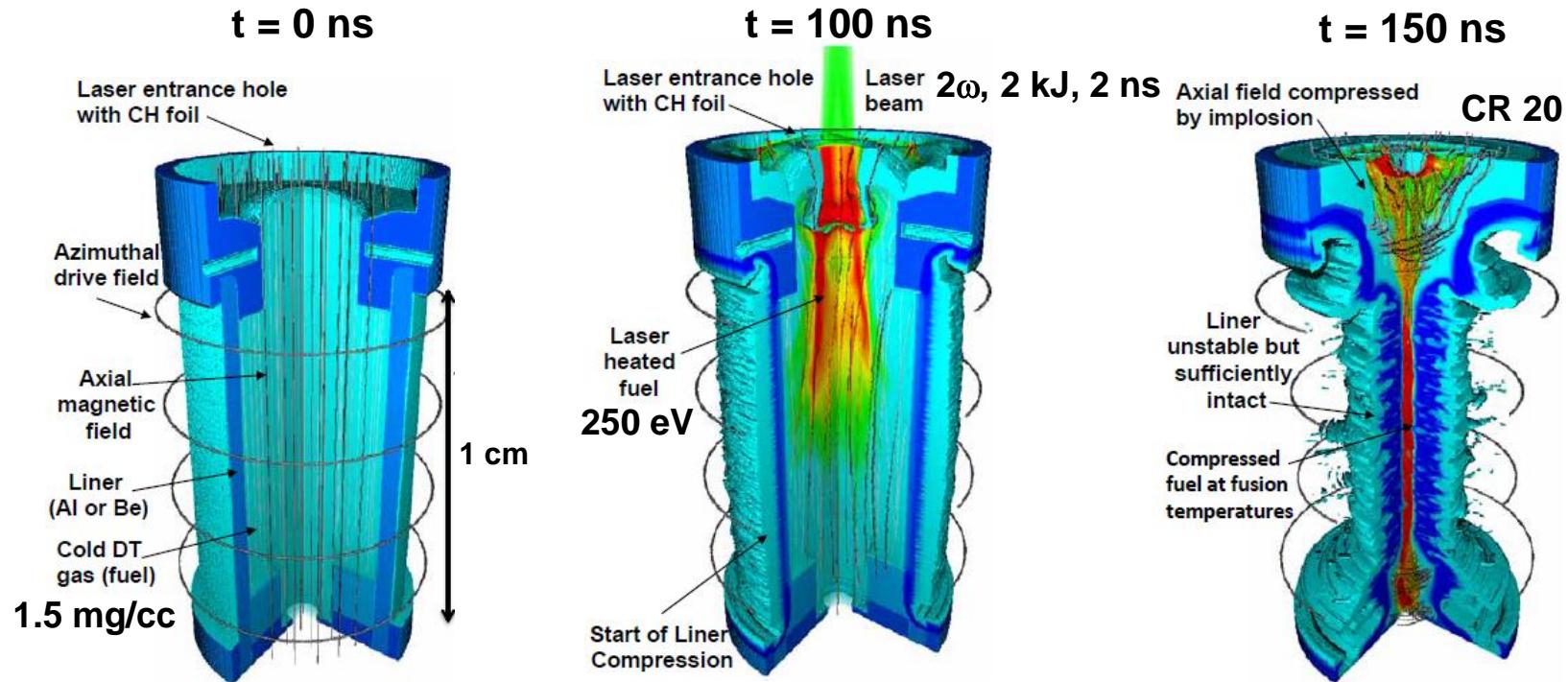


- After shots



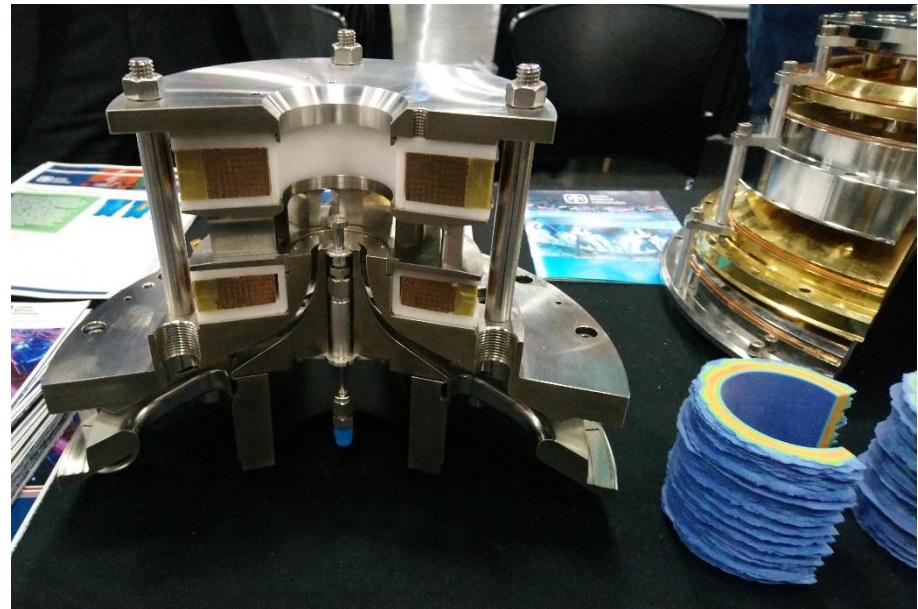
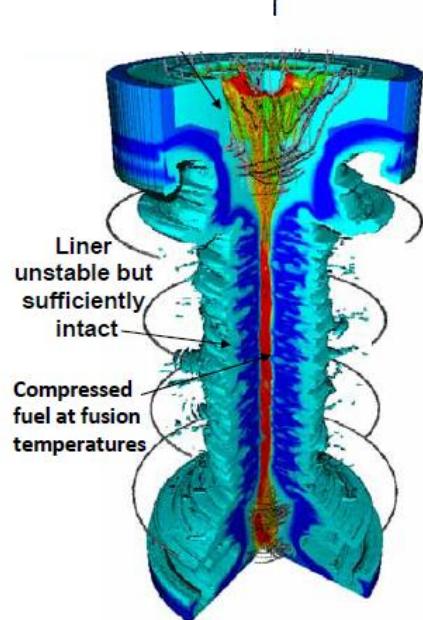
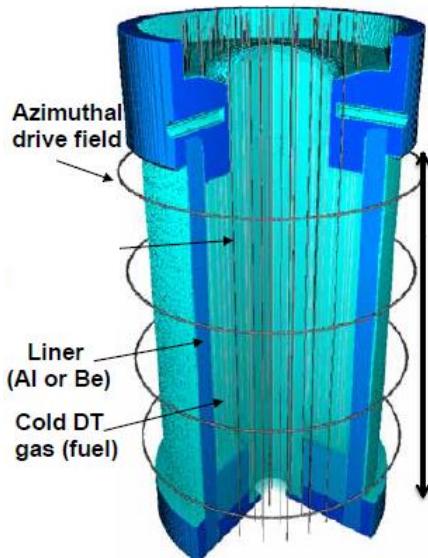
SAND2017-0900PE\_The sandia z machine - an overview of the world's most powerful pulsed power facility.pdf

# Promising results were shown in MagLIF concept conducted at the Sandia National Laboratories

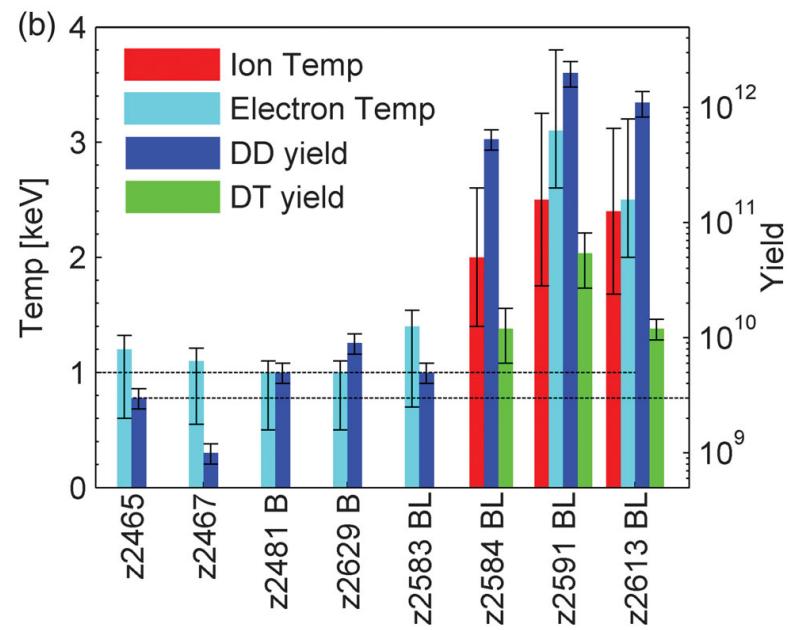
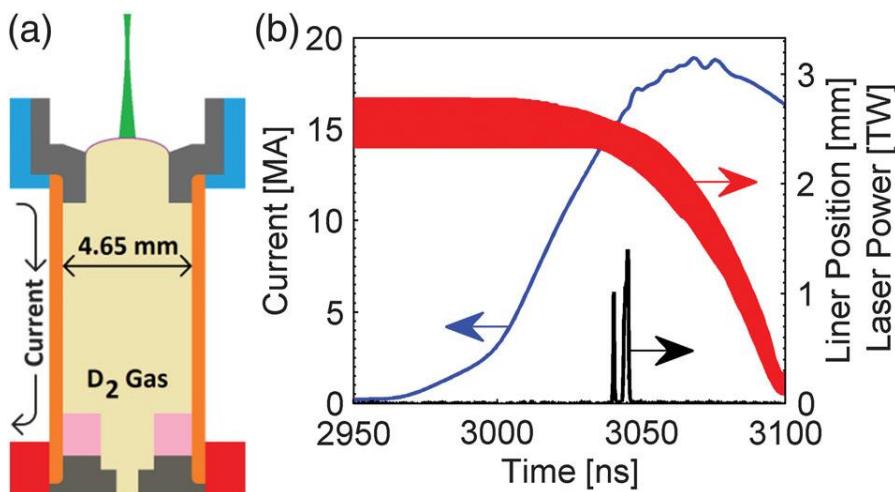


The stagnation plasma reached fusion-relevant temperatures with a 70 km/s implosion velocity

# MagLIF target



# Neutron yield increased by 100x with preheat and external magnetic field.

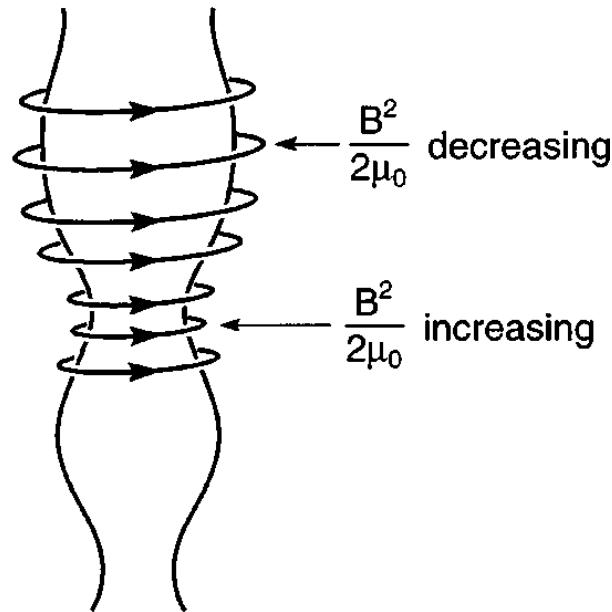


# Sheared flow stabilizes MHD instabilities

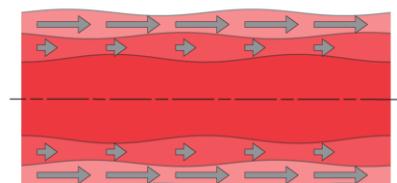
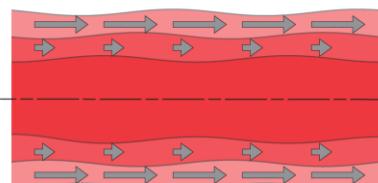
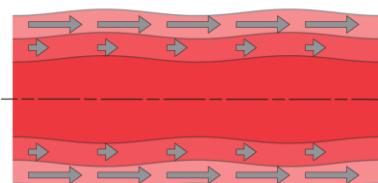
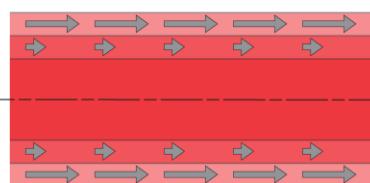
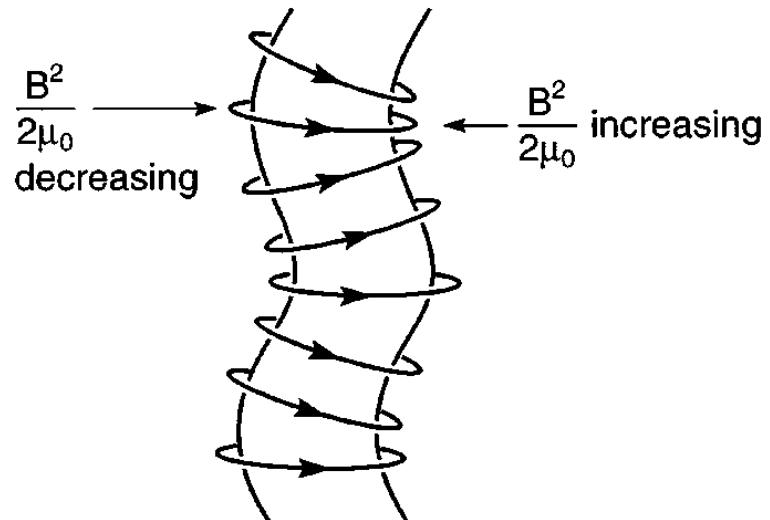


$m = 0$  (sausage)

Perturbation  $\propto e^{(im\theta + ikz + \gamma t)}$



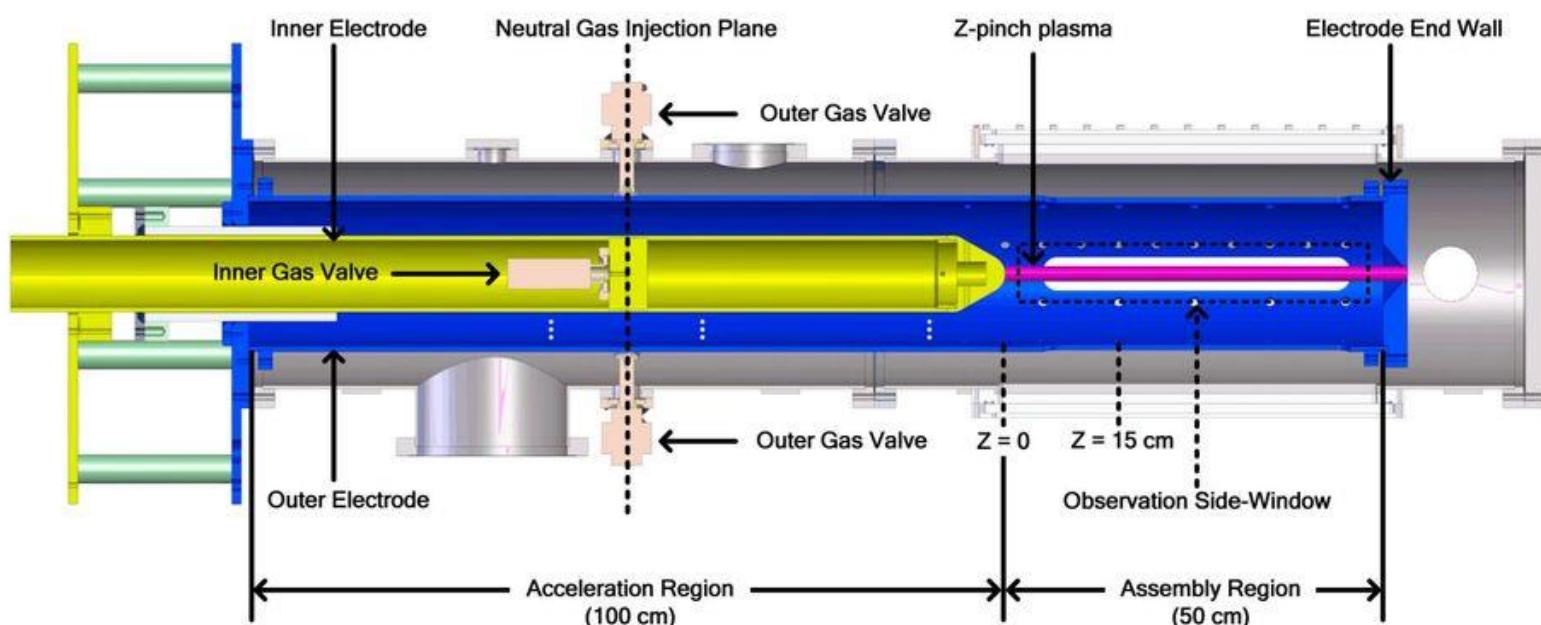
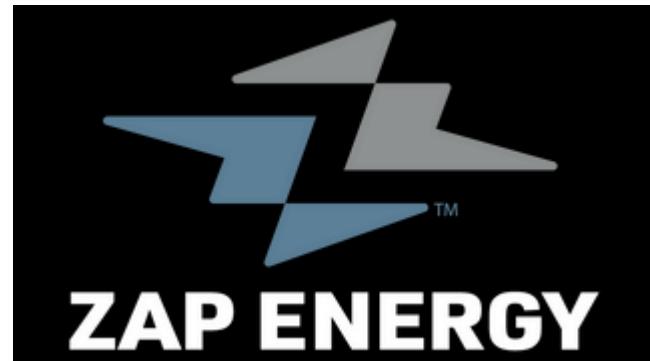
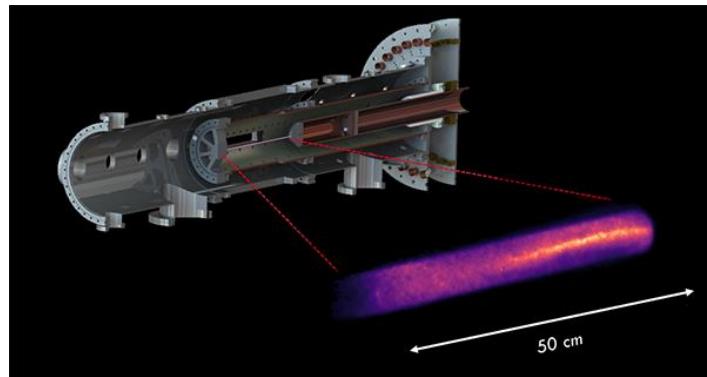
$m = 1$  (kink)



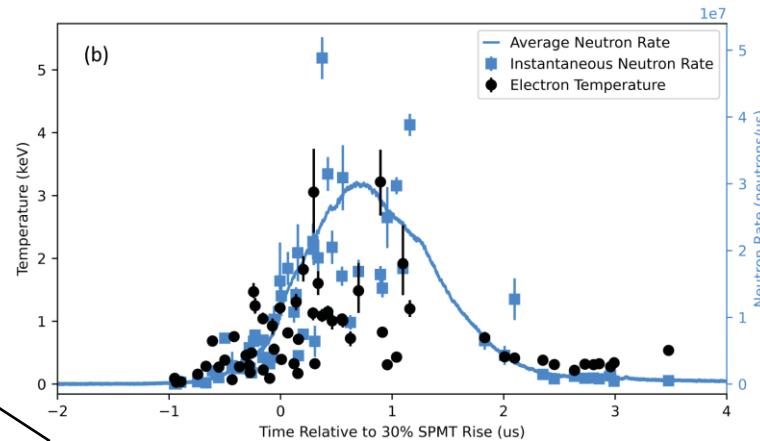
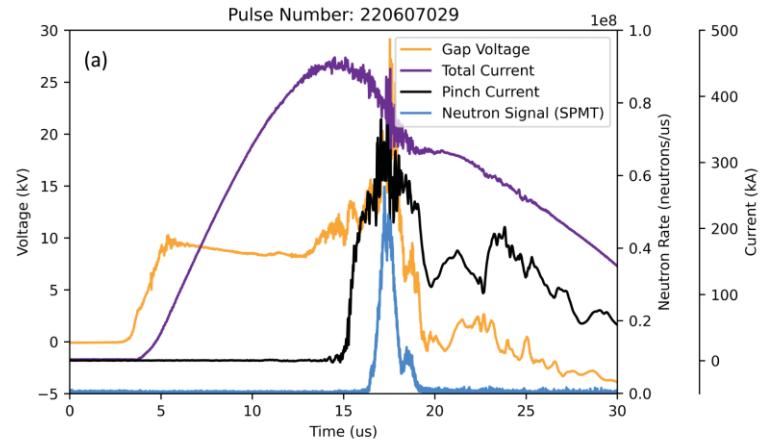
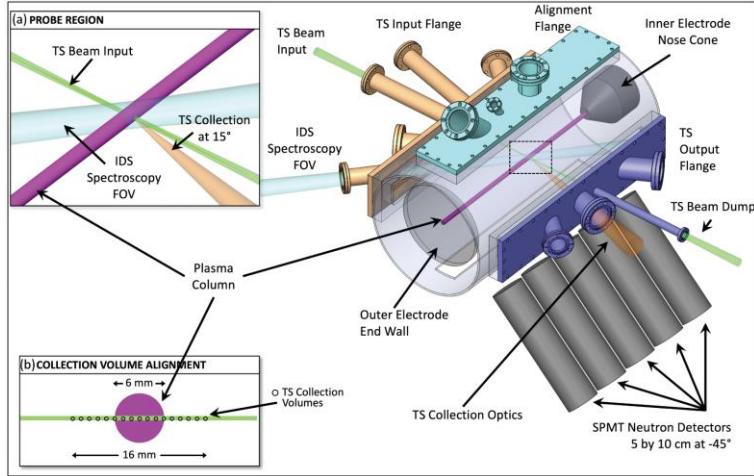
$$\frac{dV_z}{dr} \neq 0$$

- M. G. Haines, etc., Phys. Plasmas 7, 1672 (2000)  
U. Shumlak, etc., Physical Rev. Lett. 75, 3285 (1995)  
U. Shumlak, etc., ALPHA Annual Review Meeting 2017

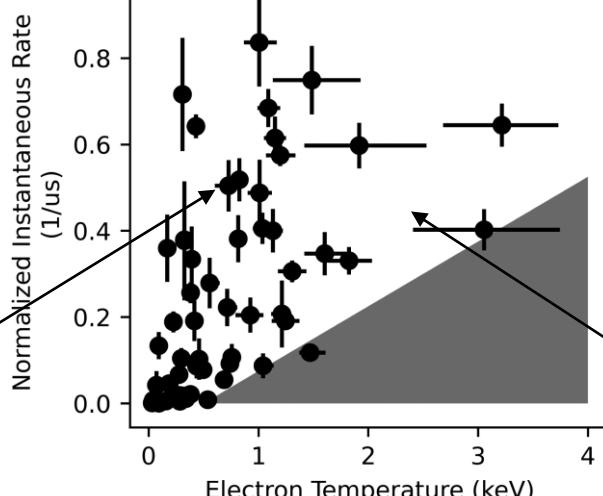
# A z-pinch plasma can be stabilized by sheared flows



# Elevated electron temperature coincident with observed fusion reactions in a sheared-flow-stabilized z pinch

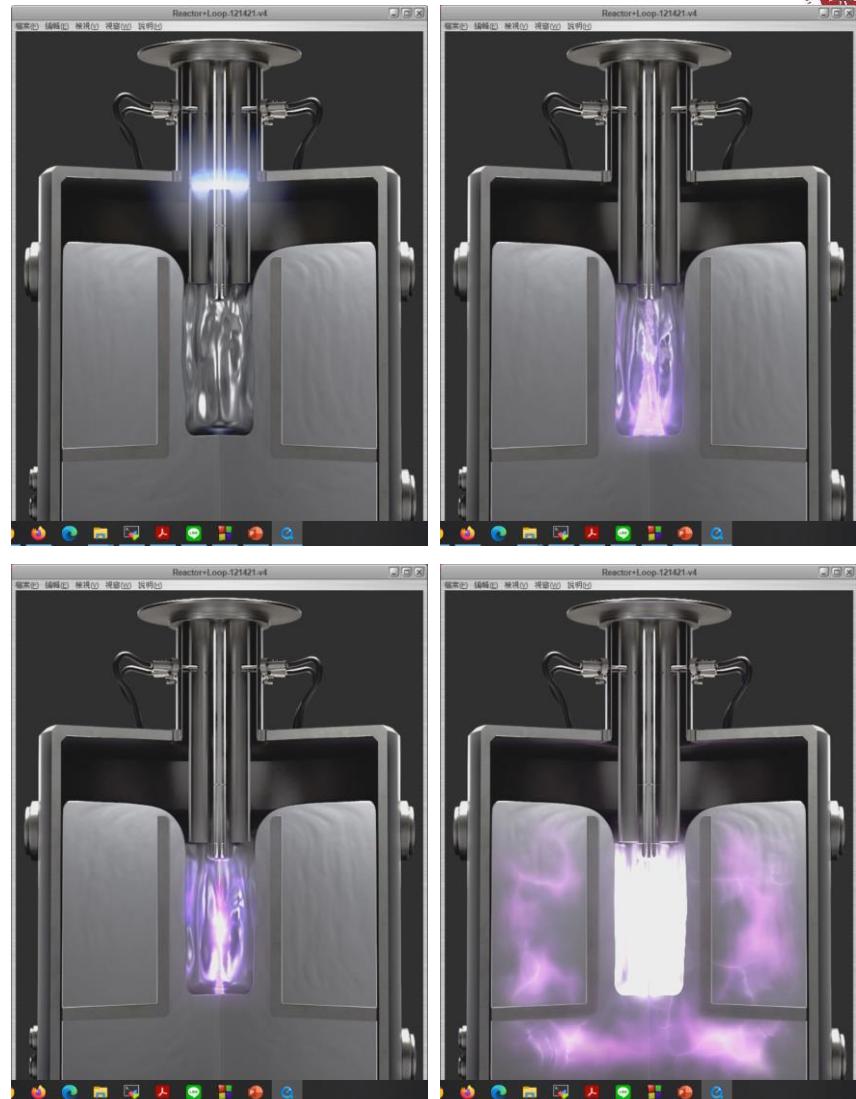
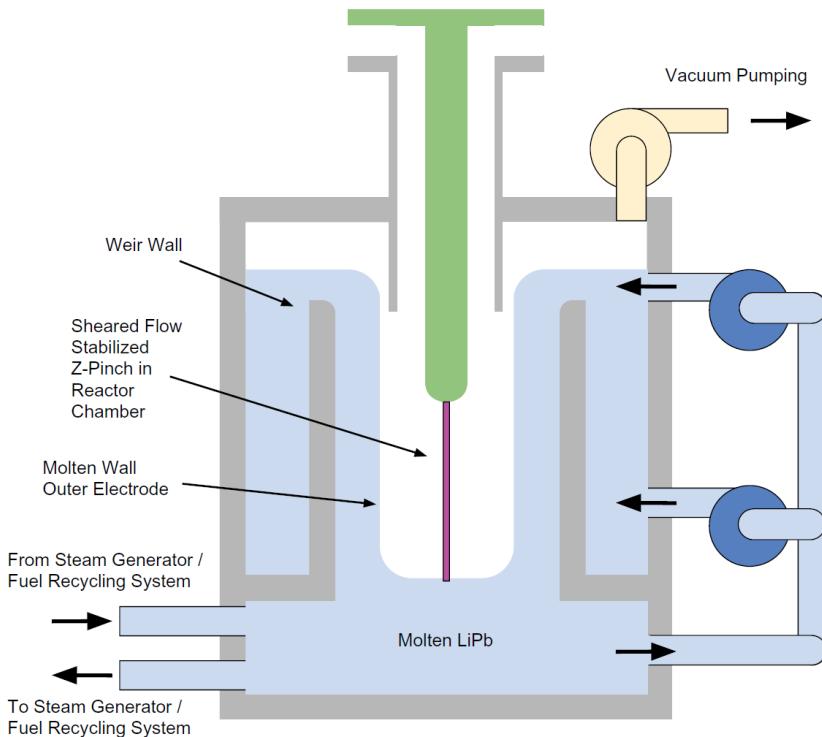


- High temperature coincide with high rate.



- Thomson scattering measurement missed the high temperature region.

# Fusion reactor concept by ZAP energy

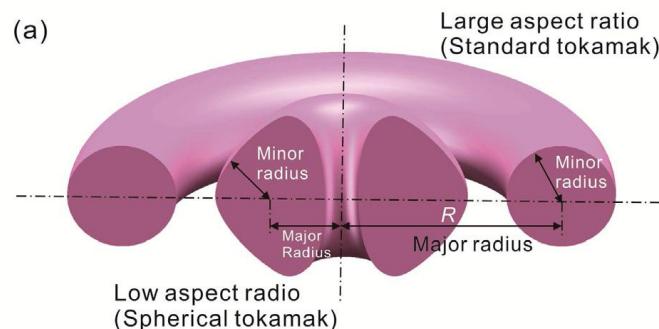


<https://www.zapenergyinc.com/about>  
E. G. Forbes, etc., Fusion Sci. Tech. 75, 599 (2019)

# Spherical torus (ST) and compact torus (CT)

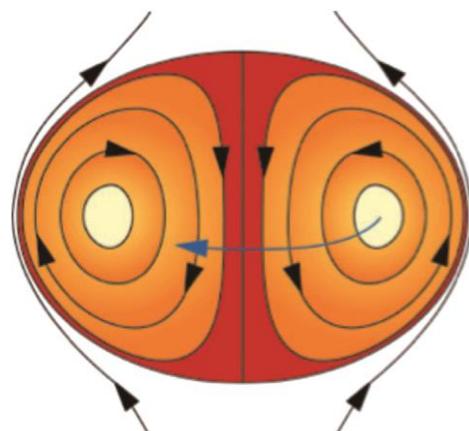


- **Spherical torus (ST)**

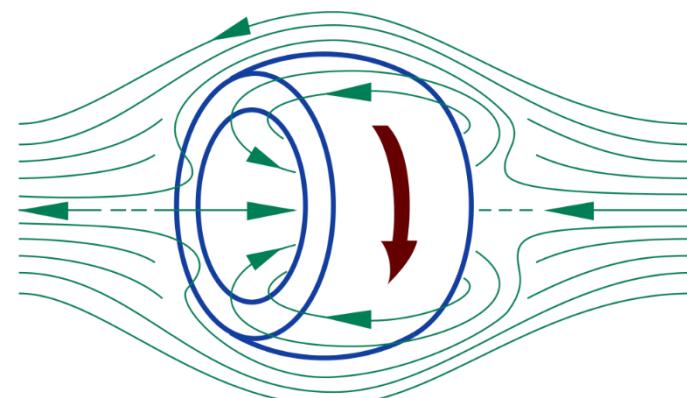


- **Compact torus (CT)**

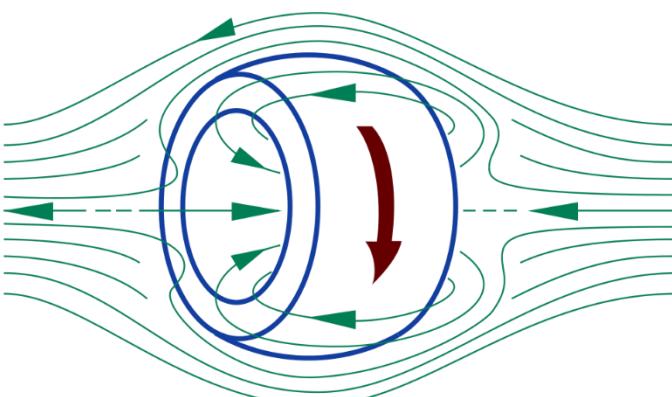
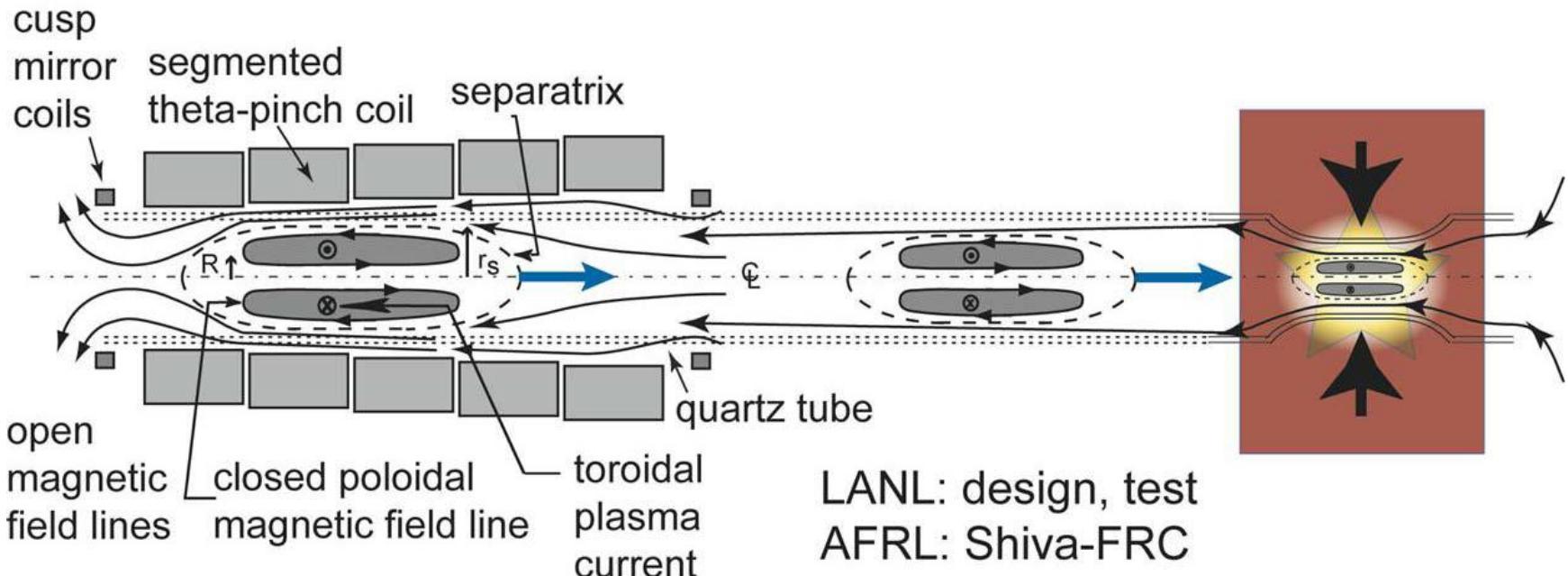
- **Spheromak**



- **Field reversed configuration (FRC)**

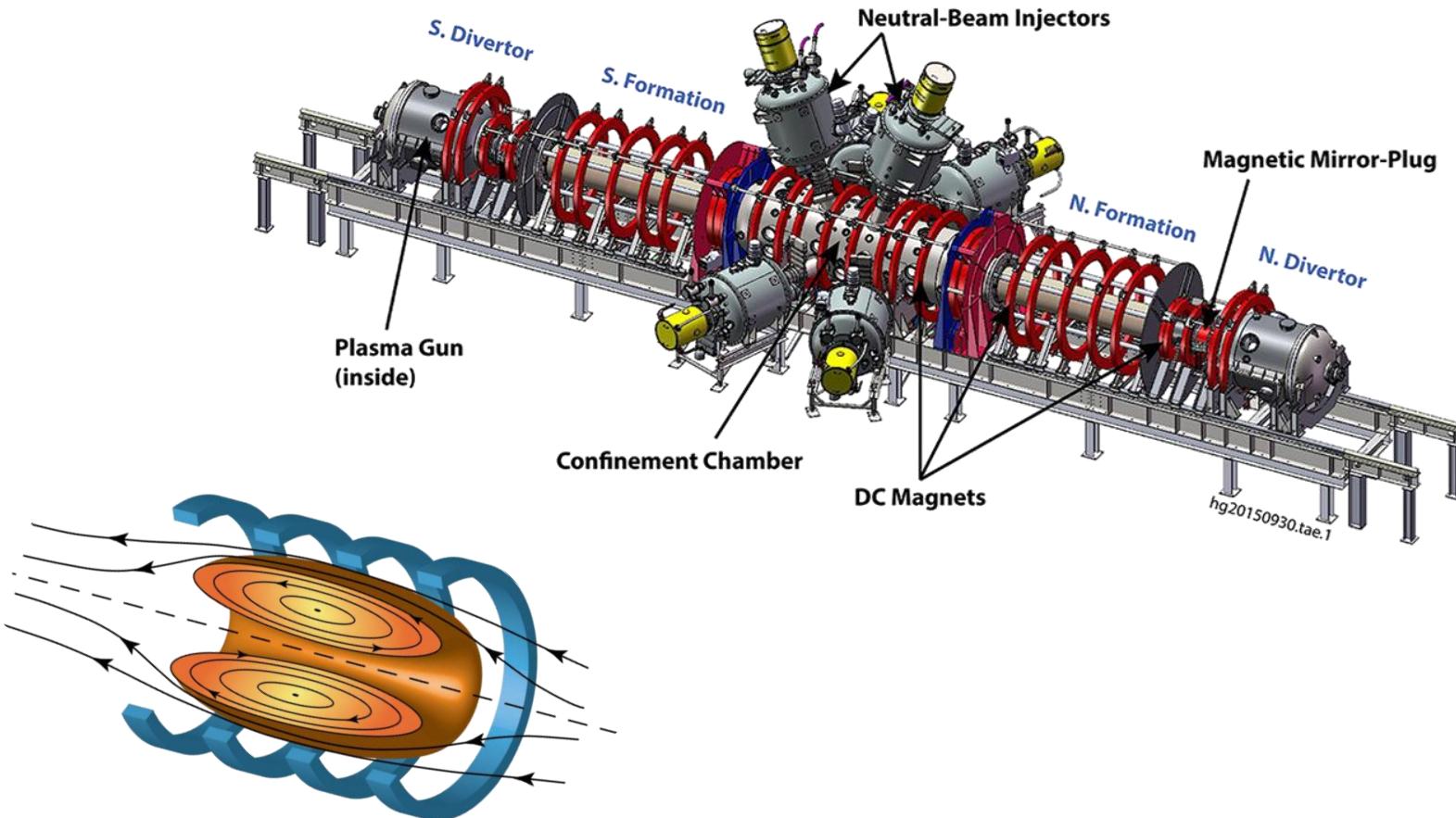


# Field reverse configuration is used in Tri-alpha energy



\*Magneto-Inertial Fusion & Magnetized HED Physics by Bruno S. Bauer, UNR & Magneto-Inertial Fusion Community  
\*\*[https://en.wikipedia.org/wiki/Field-reversed\\_configuration](https://en.wikipedia.org/wiki/Field-reversed_configuration)

# Field reverse configuration is used in Tri-alpha energy



# NBI for Tri-Alpha Energy Technologies

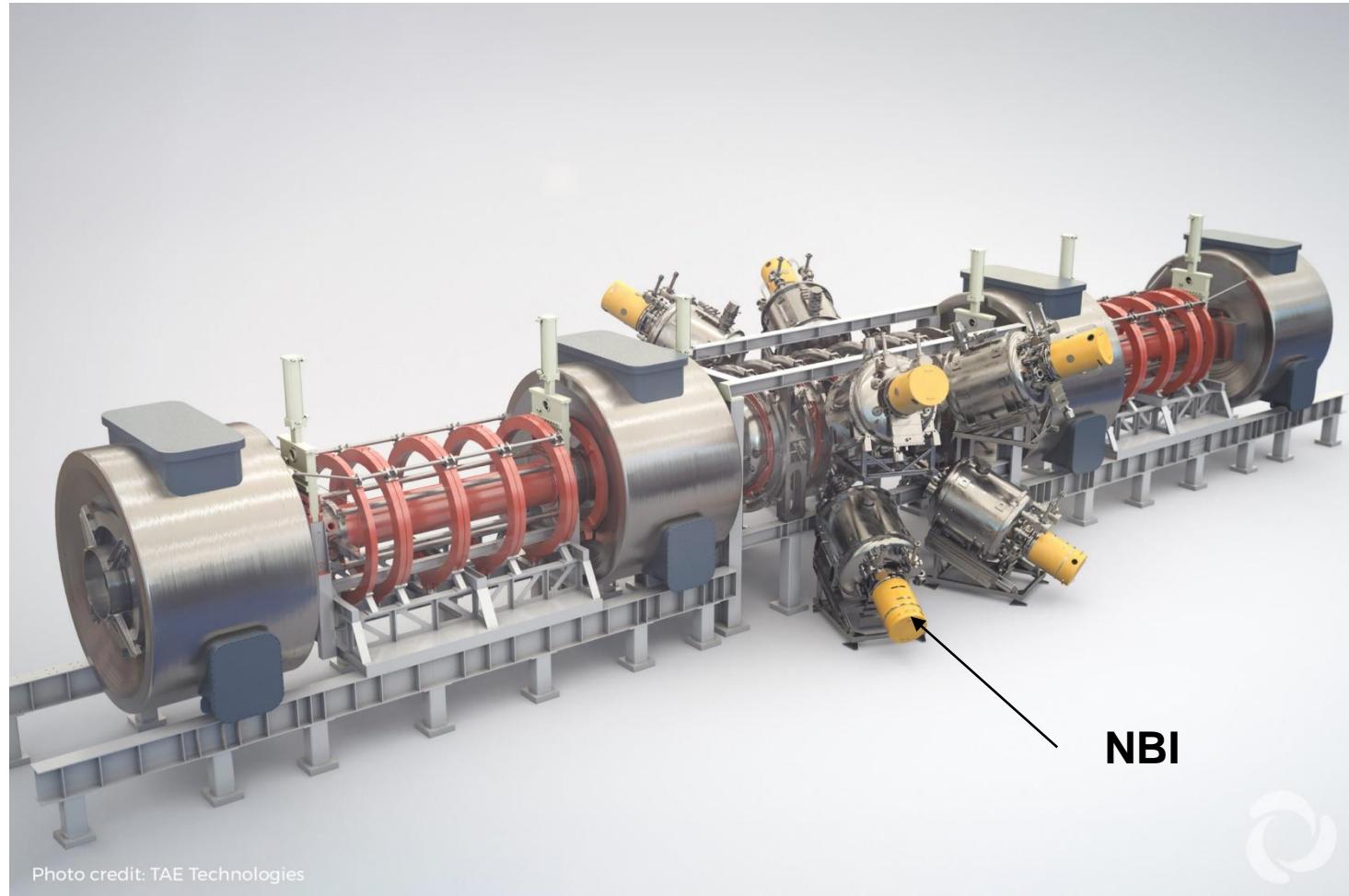


Photo credit: TAE Technologies

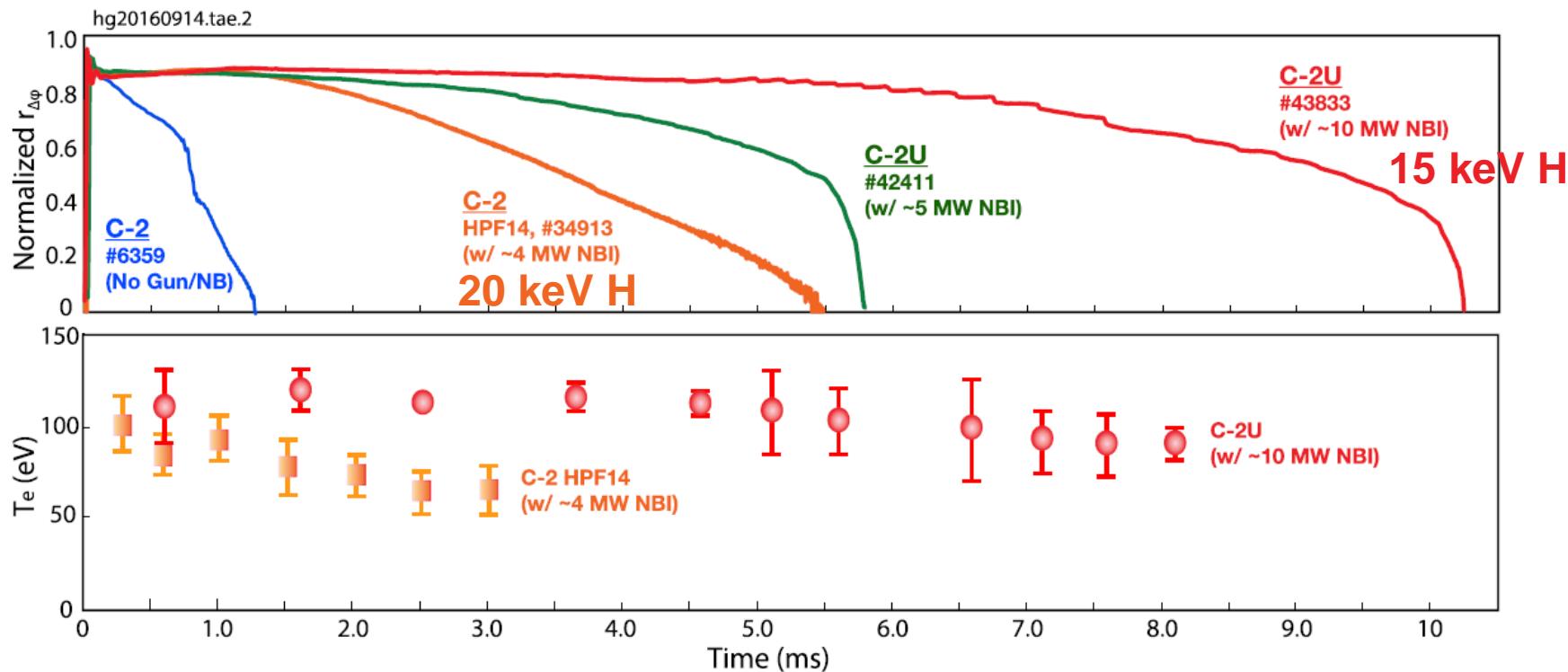
**NBI**



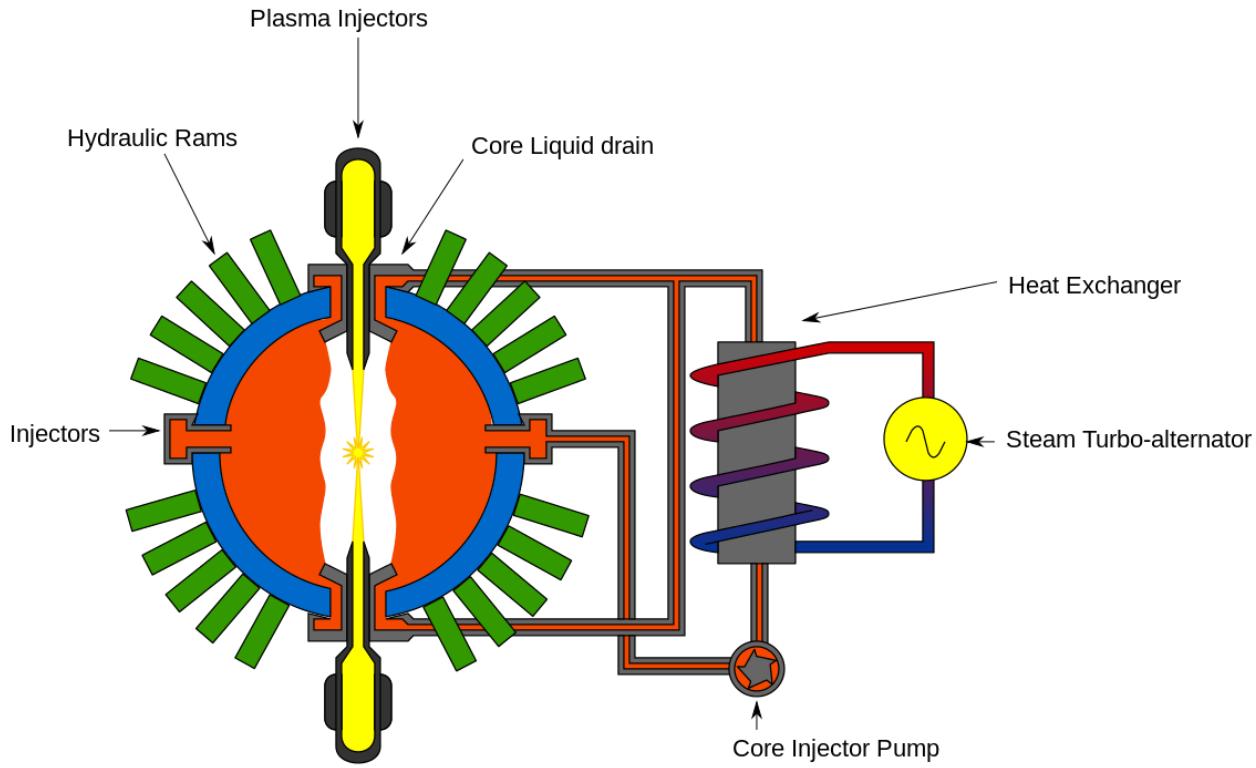
# Neutral beams are injected in to the chamber for spinning the FRC



# FRC sustain longer with neutral beam injection



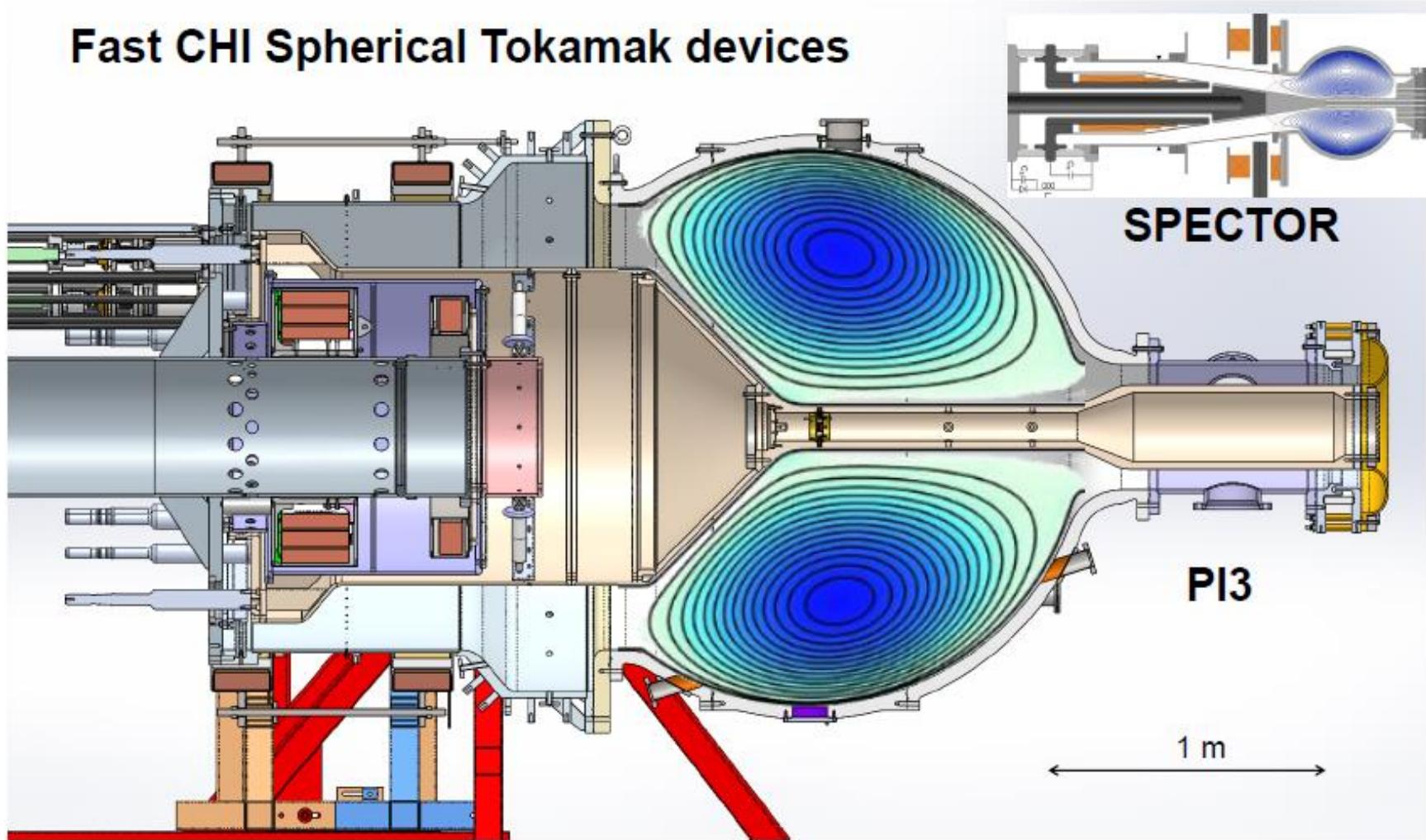
# General fusion is a design ready to be migrated to a power plant



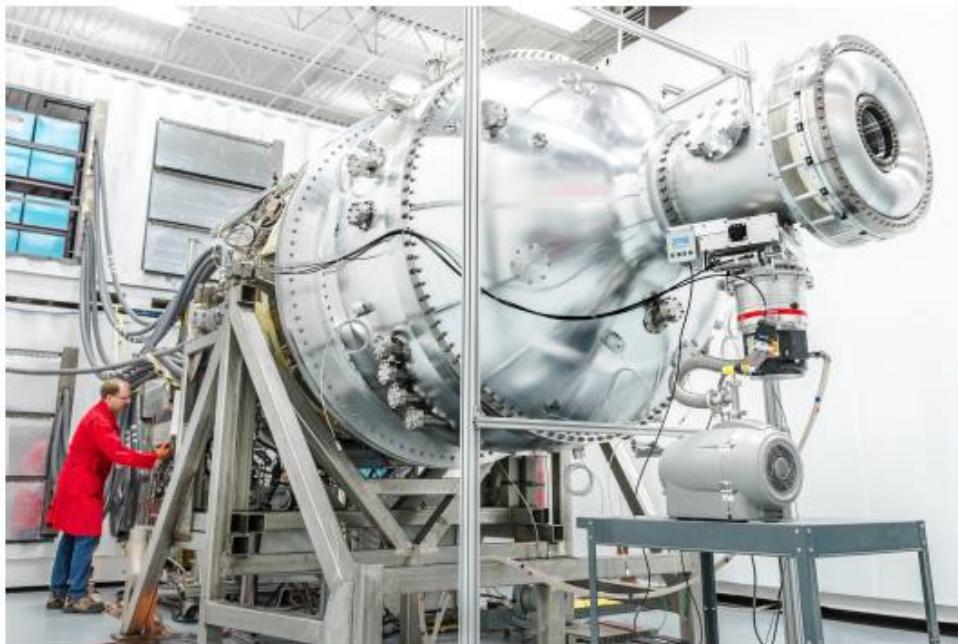
# A spherical tokamak is first generated



## Fast CHI Spherical Tokamak devices

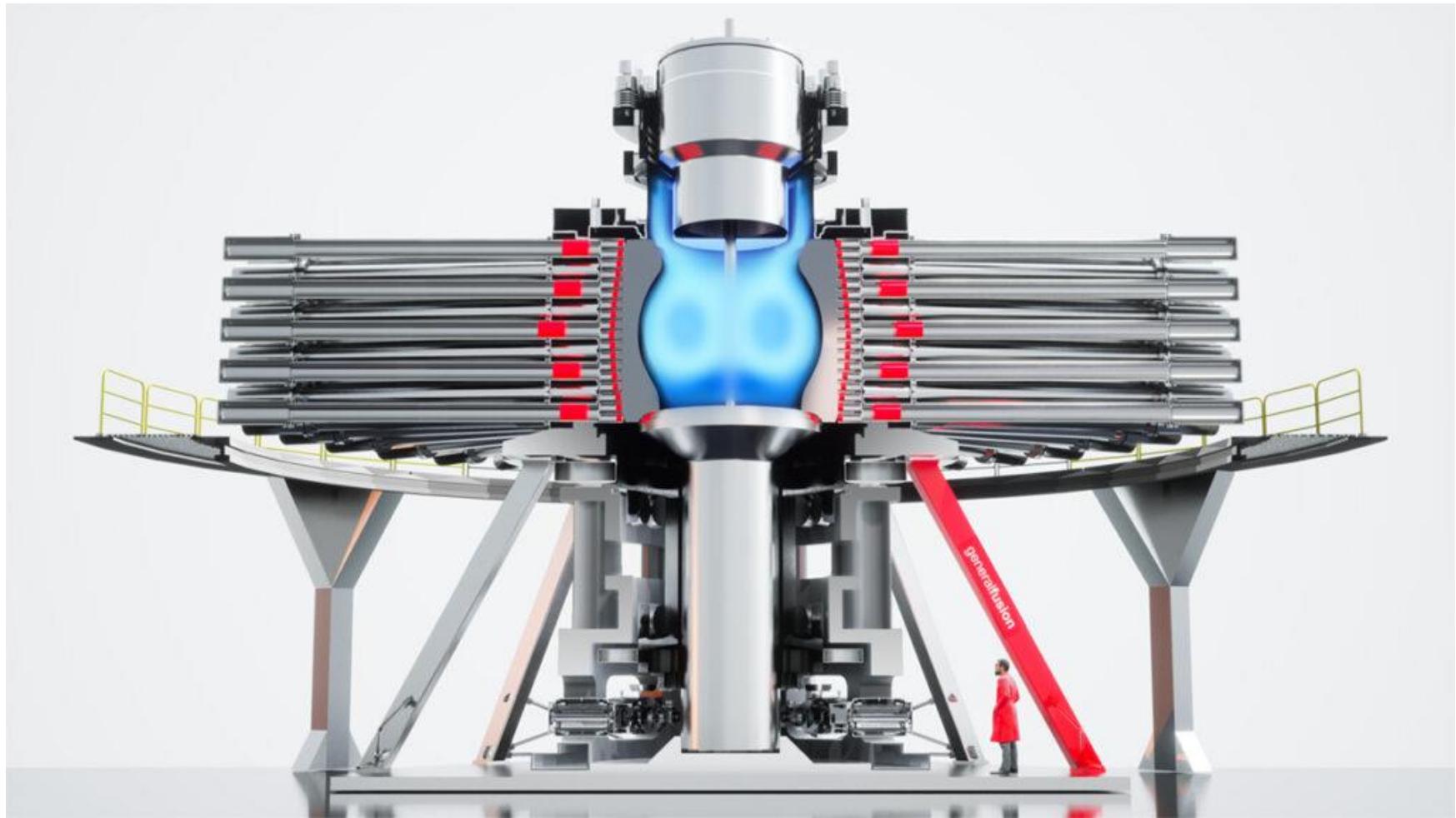


# Plasma injector for the spherical tokamak



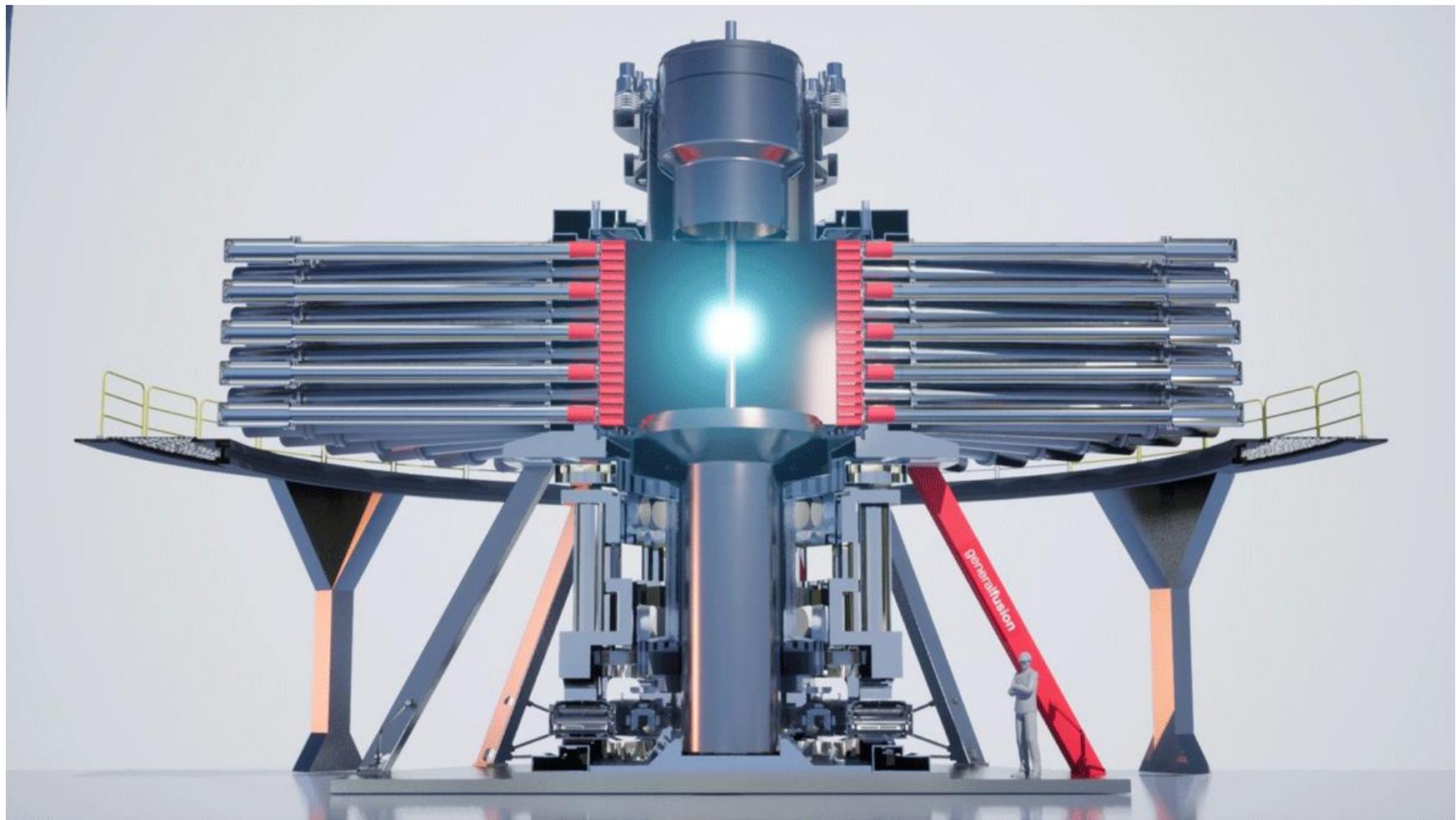
# A spherical tokamak is generated in a liquid metal vortex

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# The spherical tokamak is compressed by the pressure provided by the surrounding hydraulic pistons

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# BBC: General Fusion to build its Fusion Demonstration Plant in the UK, at the UKAEA Culham Campus



## Nuclear energy: Fusion plant backed by Jeff Bezos to be built in UK

By Matt McGrath  
Environment correspondent

17 June

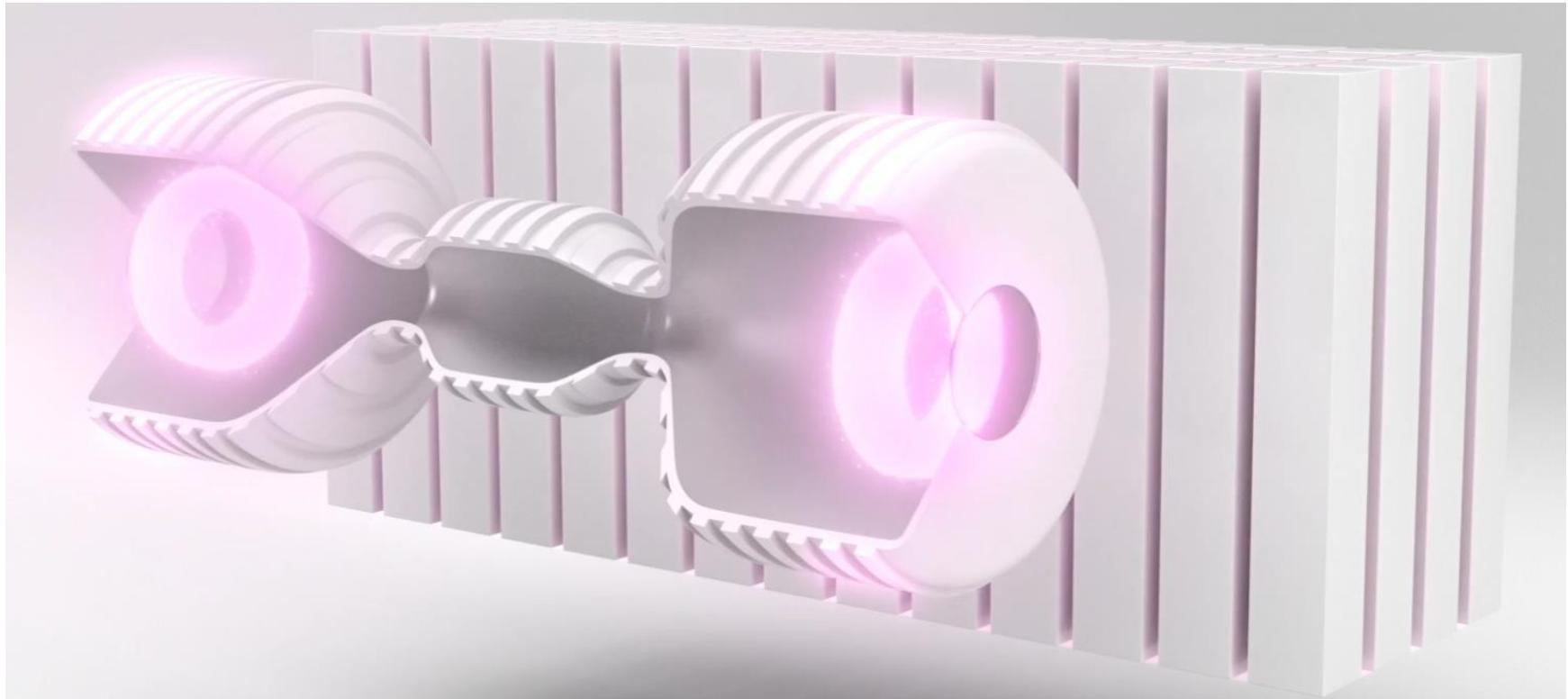


An artist's impression of what the new demonstration plant might look like

A company backed by Amazon's Jeff Bezos is set to build a large-scale nuclear fusion demonstration plant in Oxfordshire.

Canada's General Fusion is one of the leading private firms aiming to turn the

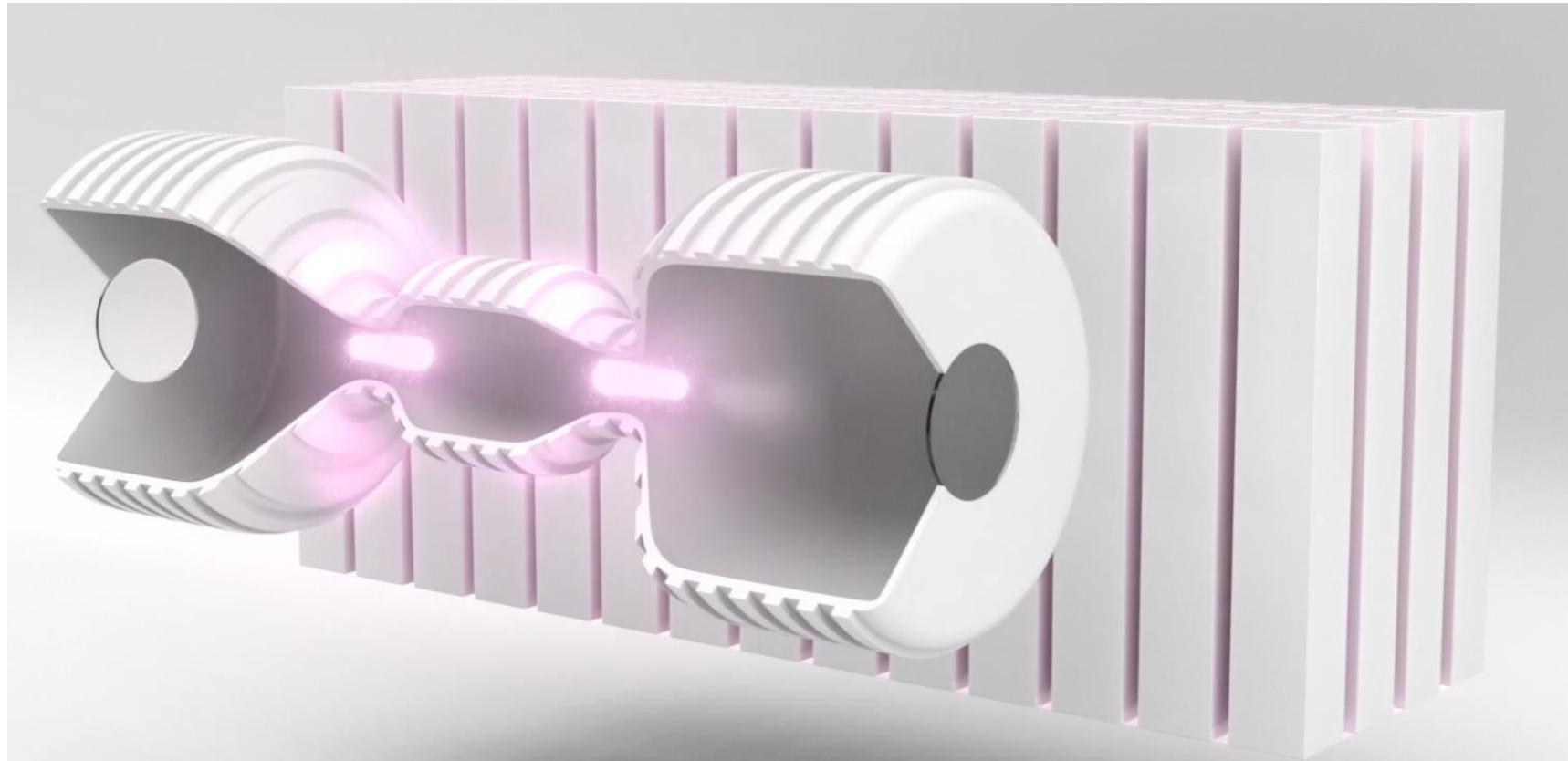
# Helion energy is compressing the two merging FRCs



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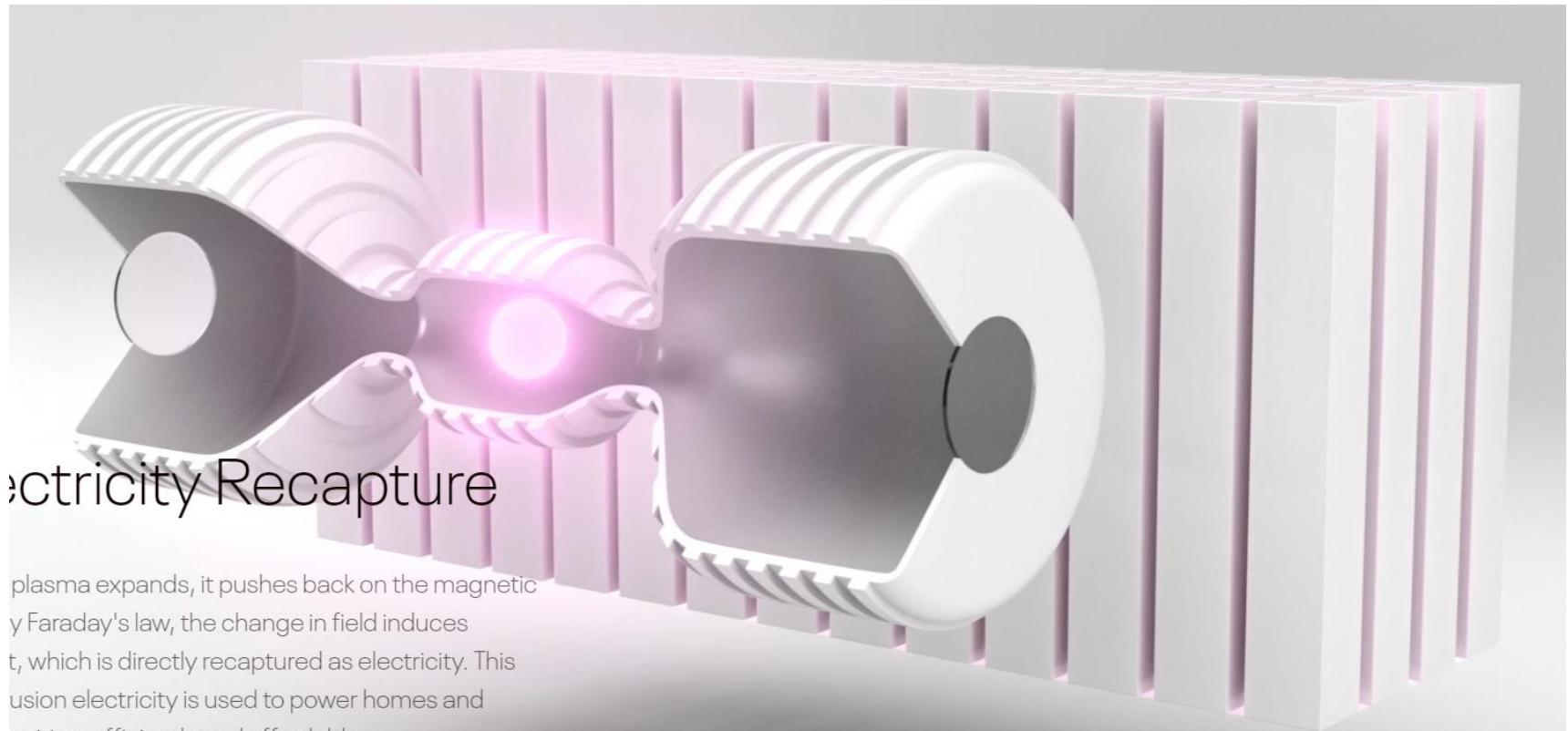
# Two FRCs are accelerated toward each other



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# Two FRCs merge with each other

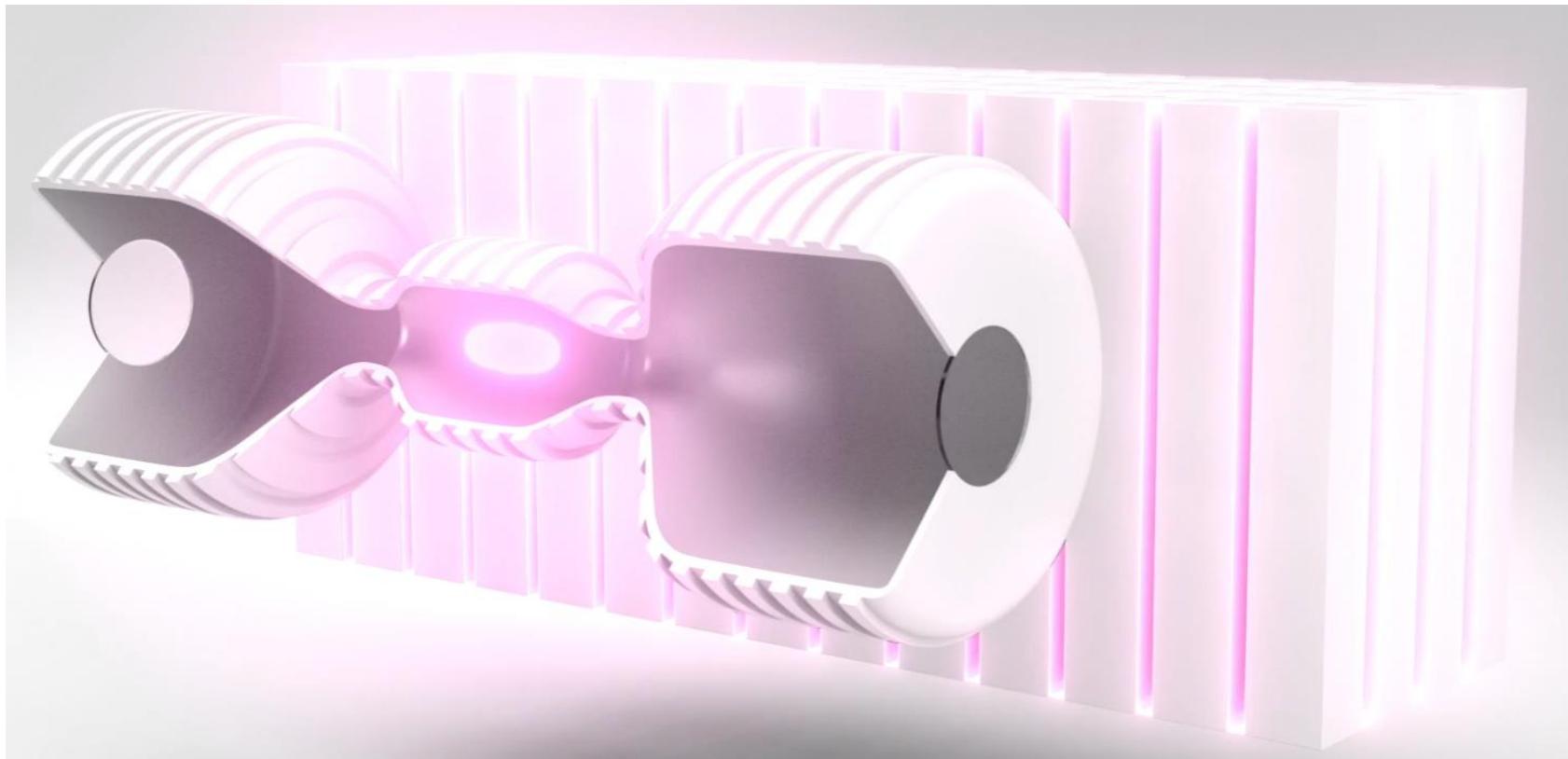


plasma expands, it pushes back on the magnetic field. By Faraday's law, the change in field induces a current, which is directly recaptured as electricity. This fusion electricity is used to power homes and businesses, efficiently and affordably.

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# The merged FRC is compressed electrically to high temperature



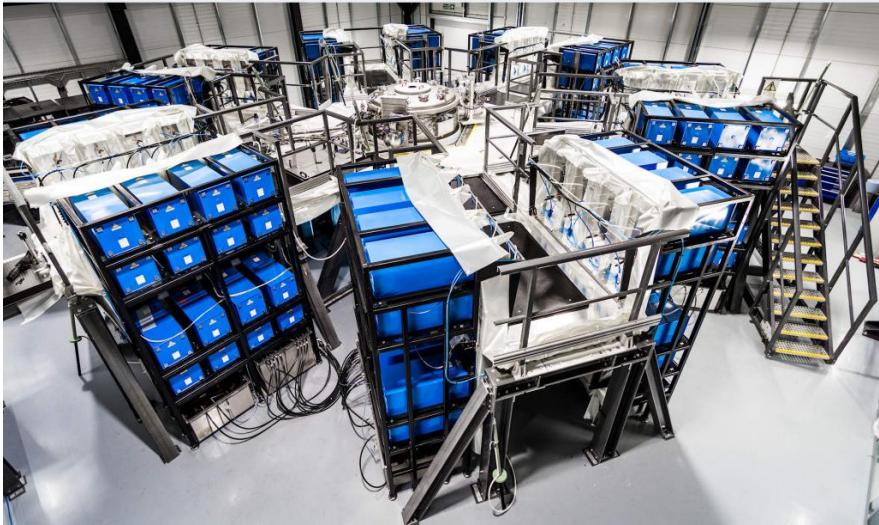
uses cookies. Read more about our privacy policy & terms of use.

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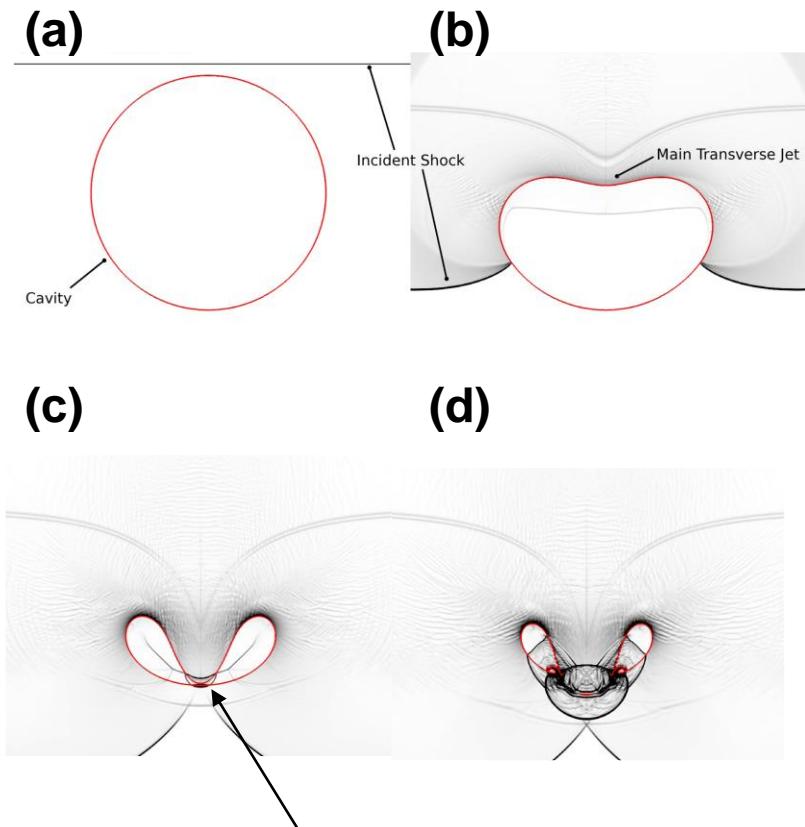
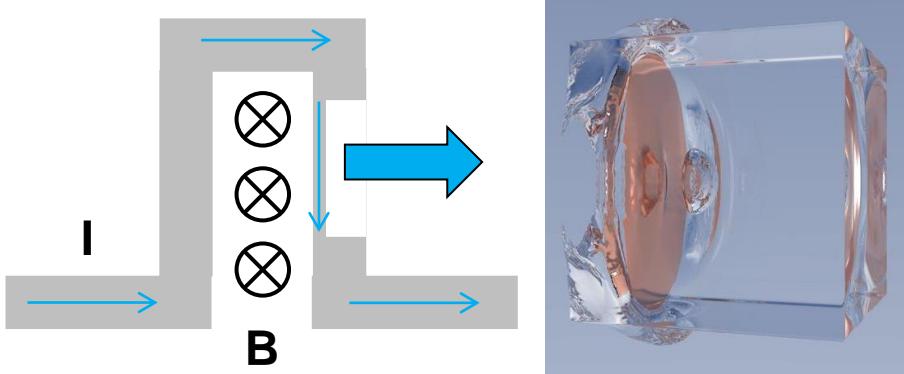
- Similar concept will be studied in our laboratory.

<https://www.helionenergy.com/>

# Projectile Fusion is being established at First Light Fusion Ltd, UK



- **Stored energy: 2.5 MJ @ 200 kV  
( $C_{tot}=125 \mu F$ )**
- $I_{peak}=14 \text{ MA w/ } T_{rise}\sim 2\text{us.}$



- **High pressure is generated by the colliding shock.**

<https://firstlightfusion.com/>

B. Tully and N. Hawker, Phys. Rev. E93, 053105 (2016)

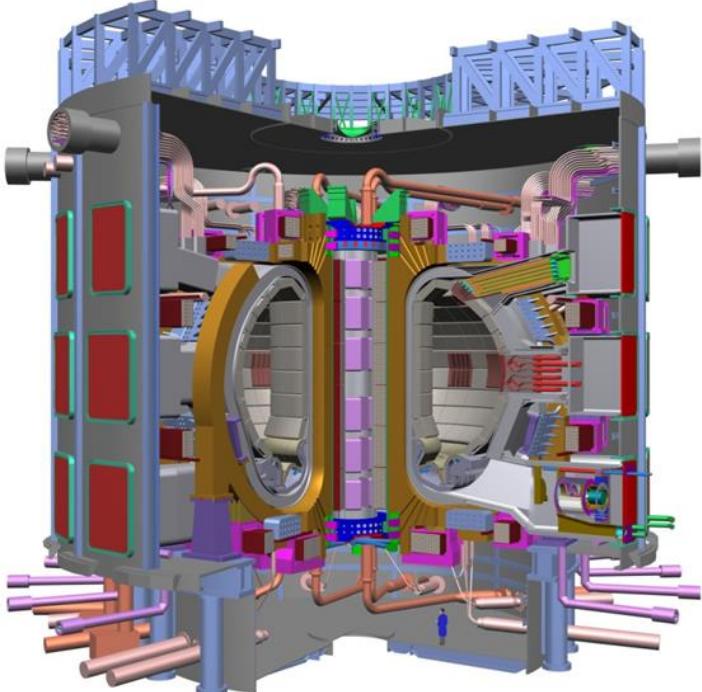
# A gas gun is used to eject the projectile



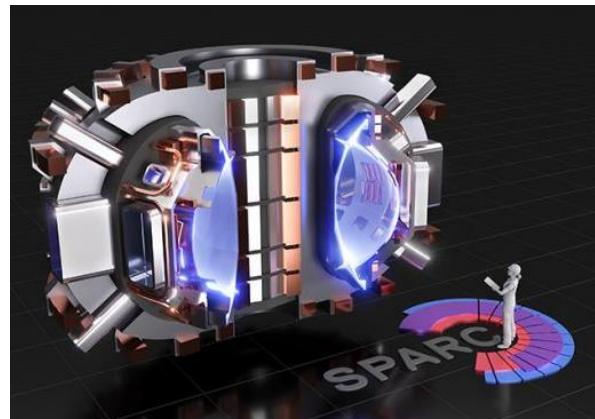
<https://www.youtube.com/watch?v=JN7lyxC11n0>  
<https://www.youtube.com/watch?v=aW4eufacf-8>

# Many groups aim to achieve ignition in the MCF regime in the near future

- ITER – 2025 First Plasma  
2035 D-T Exps  
2050 DEMO
- Tokamak energy, UK
  - 2025 Gain
  - 2030 to power grid



- Commonwealth Fusion Systems, USA – 2025 Gain



<https://www.iter.org>

<https://www.tokamakenergy.co.uk/>

<https://www.psfc.mit.edu/sparc>



# Fusion is blooming



## FIA Members

FUSION  
INDUSTRY  
ASSOCIATION

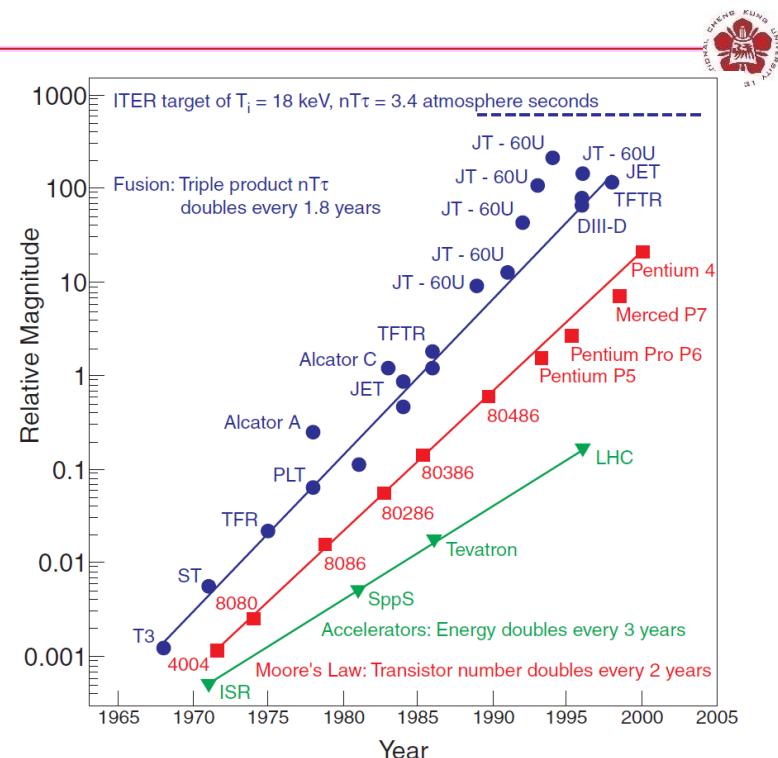
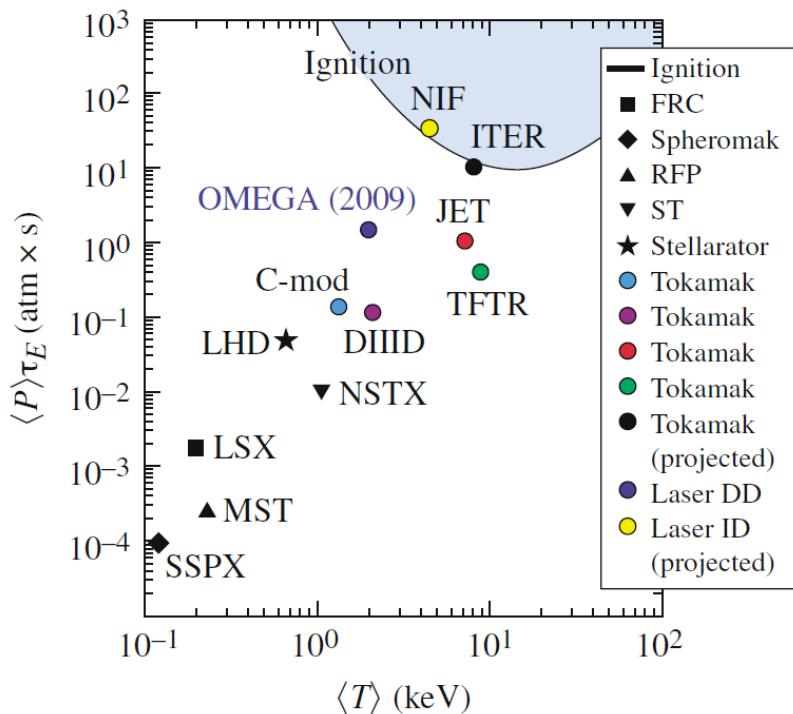


generalfusion™



:

# We are closed to ignition!



- **Other private companies:**



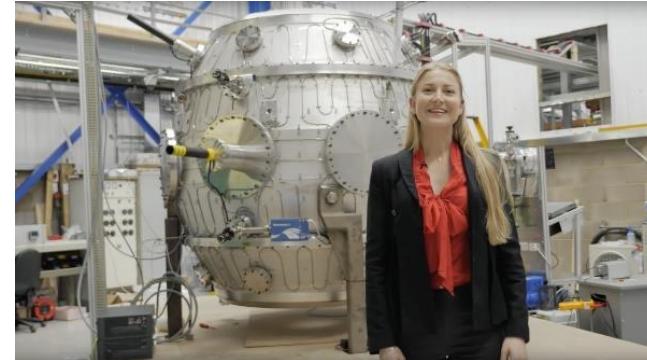
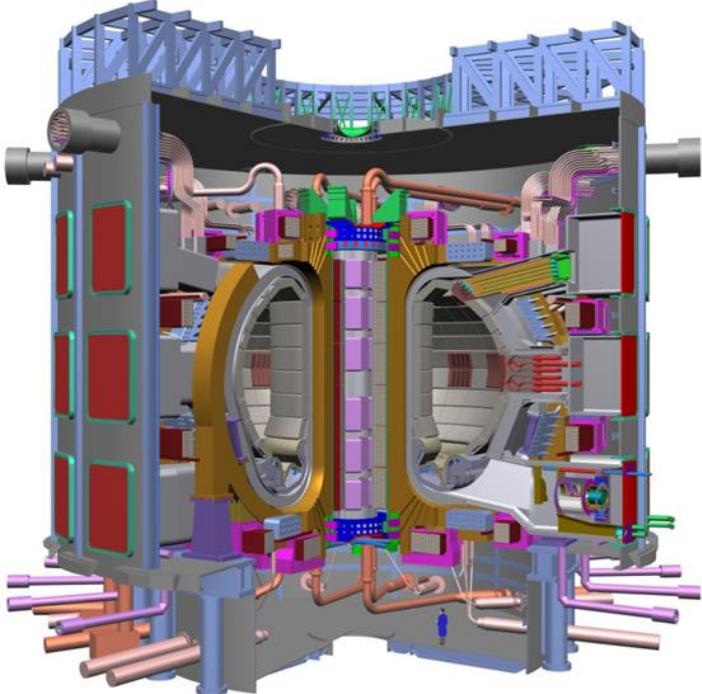
# Tokamak

MTF

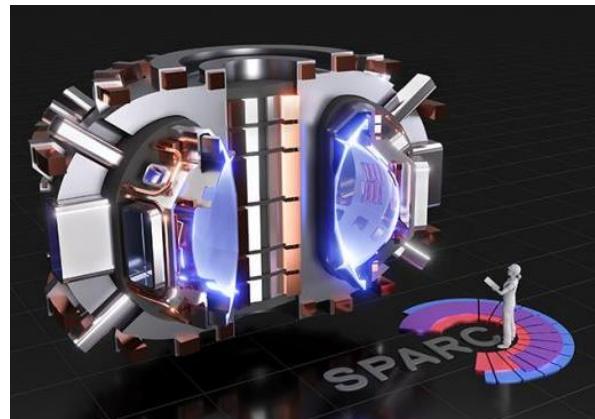
ICF

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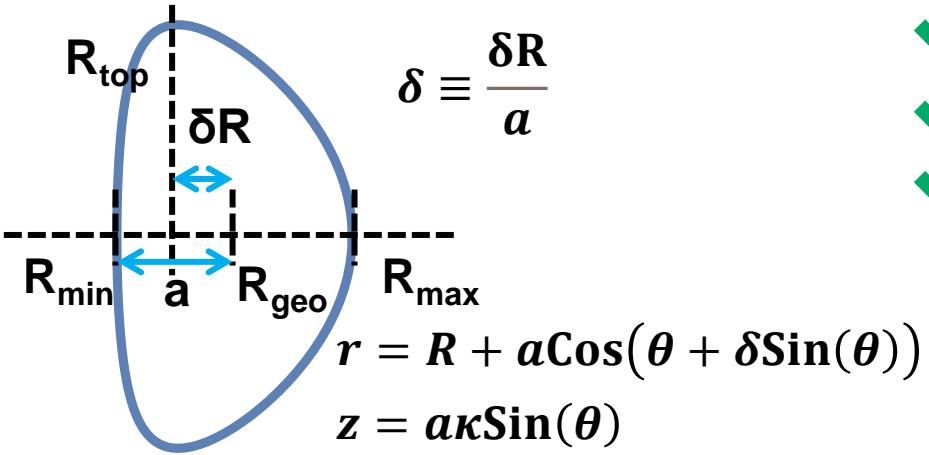
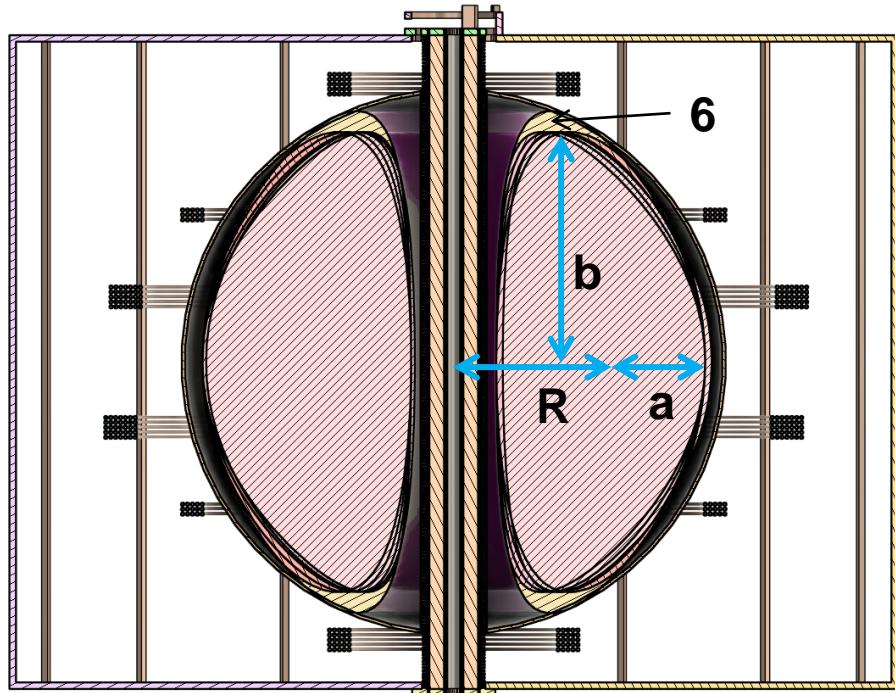
<https://www.iter.org>

<https://www.tokamakenergy.co.uk/>

<https://www.psfc.mit.edu/sparc>



# A new design using a spherical chamber can tolerate several potential shapes and sides calculated by the theory group



- Parameters:
  - Elongation  $\kappa=b/a$
  - Triangularity  $\delta$
  - $T \sim 100$  eV
  - $B_T \sim 0.5$  T
  - $I_p \sim 100$  kA

|     | R (cm) | a (cm) | R/a  | $\kappa$ | $\delta$ |
|-----|--------|--------|------|----------|----------|
| ✓ 1 | 45     | 32     | 1.41 | 2.2      | 0.5      |
| ✓ 2 | 45     | 32     | 1.41 | 2.2      | 0.3      |
| ✓ 3 | 45     | 32     | 1.41 | 2.2      | 0.4      |
| ✓ 4 | 45     | 32     | 1.41 | 2.2      | 0.6      |
| ✓ 5 | 47     | 32     | 1.47 | 2.2      | 0.5      |

• We welcome anyone interested in fusion research to join us!