

3.3 Plasma as Fluids

p5:

- * In previous chapter: \vec{E} & \vec{B} are not prescribed but are determined by the positions & the motions of the charge themselves
- * A self-consistent problem needs to be considered
 - particles generates field by their charges (\vec{E})
 - and motions ($\vec{v} \rightarrow \vec{j}$)
 - a time-varying situation.
- * Typical plasma density $\sim 10^{12} / \text{cm}^3$ ion-electron pair
 - impossible to predict trajectories of all particles
 - Fluid model — identity of the individual particle is neglected, only the motion of fluid elements is taken into account.
 - In an ordinary fluid, frequent collisions between particles keep the particle in a fluid element moving together
 - It is surprising that the model works for plasma, which generally have in frequent collisions.

Q. 3.1. Relation of Plasma physics to ordinary electromagnetics.

Q. 3.1.1 Maxwell's Eq.

In vacuum:

$$\epsilon_0 \nabla \cdot \vec{E} = \sigma \quad \leftarrow \text{free charge}$$

$$\nabla \times \vec{E} = -\dot{\vec{B}}$$

$$\nabla \cdot \vec{B} = 0$$

$$\nabla \times \vec{B} = \mu_0 (\vec{J} + \epsilon_0 \dot{\vec{E}})$$

current density

In a medium:

$$\nabla \cdot \vec{D} = \sigma$$

$$\nabla \times \vec{E} = -\dot{\vec{B}}$$

$$\nabla \cdot \vec{B} = 0$$

$$\nabla \times \vec{H} = \vec{J} + \dot{\vec{D}}$$

$$\vec{D} = \epsilon \vec{E} \quad \leftarrow \text{"bound" charge \&}$$

$$\vec{B} = \mu \vec{H} \quad \leftarrow \text{current density arising from polarization and magnetization of the medium.}$$

In plasma, ions & electrons comprising the plasma are the equivalent of the "bound" charge & current. However, they move in a complicated way, it's hard to lump their effect into ϵ, μ . Consequently, in plasma physics, one generally work w/ the vacuum eq. where σ_0 and \vec{J} include All the charges and currents, both external and internal.

3.1.2 Classical Treatment of Magnetic Materials p55

- Each gyrating particle has magnetic moment
 \rightarrow plasma is a magnetic material w/ a permeability μ_m (use μ later)
- Ferromagnetic domains:

$$\vec{M} = \frac{1}{V} \sum_i \vec{\mu}_i \quad \leftarrow \text{magnetization per unit volume.}$$

bound current density: $\vec{j}_b = \nabla \times \vec{M}$

$$\frac{1}{\mu_0} \nabla \times \vec{B} = \vec{j}_f + \vec{j}_b + \epsilon_0 \dot{\vec{E}}$$

Make it into a simple form:

$$\nabla \times \vec{H} = \vec{j}_f + \epsilon_0 \dot{\vec{E}} \quad \text{where}$$

$$\vec{H} = \mu_0^{-1} \vec{B} - \vec{M}$$

Assuming $\vec{M} \propto \vec{B}$ or \vec{H}

$$\vec{M} = \chi_m \vec{H}$$

$$\Rightarrow \vec{B} = \mu_0 (1 + \chi_m) \vec{H} = \mu_m \vec{H}$$

\leftarrow important !!

In plasma, each particle has a magnetic moment μ_B

$$\vec{M} = \frac{1}{V} \sum_i \vec{\mu}_i$$

$$\therefore \mu_B = \frac{m v_{\perp}^2}{2B} \propto \frac{1}{B} \Rightarrow M \propto \frac{1}{B} \quad \text{Not linear!!}$$

\Rightarrow Not useful to consider a plasma as a magnetic medium.

3.1.3 Classical Treatment of Dielectrics p56

polarization \vec{P} per volume:

$$\vec{P} = \frac{1}{V} \sum \vec{P}_i$$

bound charge density:

$$\sigma_b = -\nabla \cdot \vec{P}$$

$$\epsilon_0 \nabla \cdot \vec{E} = (\sigma_f + \sigma_b) \Rightarrow \nabla \cdot \vec{D} = \sigma_f$$

$$\vec{D} = \epsilon_0 \vec{E} + \vec{P} = \epsilon \vec{E}$$

Assuming $\vec{P} = \epsilon_0 \chi_e \vec{E}$ where

$$\epsilon = (1 + \chi_e) \epsilon_0$$

→ We may try to get an expression for ϵ in plasma.

3.1.4 The dielectric Constant of a Plasma

polarization \vec{P} current, from fluctuating $\vec{E}(t)$

Continuity:

$$\frac{\partial \sigma_p}{\partial t} + \nabla \cdot \vec{J}_p = 0$$

$$\left(\frac{\partial \sigma_p}{\partial t} \right) \vec{J}_p$$

* A polarization effect does not arise in plasma unless the electric field is time varying.

* Note that $\vec{J}_p = n_e (\vec{V}_{ip} - \vec{V}_{ep}) = \frac{n_e}{e B^2} (m+m) \frac{d\vec{E}}{dt}$

$$= \frac{\rho}{B^2} \frac{d\vec{E}}{dt}$$

$$v_p = \pm \frac{1}{\omega_{ce}} \frac{d\vec{E}}{dt}, \text{ polarization drift}$$

Also

$$\begin{aligned} \nabla \times \vec{B} &= \mu_0 (\vec{J}_f + \vec{J}_p + \epsilon_0 \dot{\vec{E}}) \\ &= \mu_0 (\vec{J}_f + \epsilon \dot{\vec{E}}) \end{aligned}$$

$$\epsilon = \epsilon_0 + \frac{\vec{J}_p}{\vec{E}} \stackrel{\text{polarization drift}}{=} \epsilon_0 + \frac{\rho}{B^2}$$

P57

$$\text{or } \epsilon_R = \frac{\epsilon}{\epsilon_0} = 1 + \frac{\mu_0 \rho c^2}{B^2} \quad \text{where } c^2 = \frac{1}{\mu_0 \epsilon_0}$$

* Low-frequency plasma dielectric constant for transverse motions.

Note that $\vec{J}_p = \frac{\rho}{B^2} \vec{E}$ only valid for

$$\omega^2 \ll \omega_c^2 \quad \& \quad \vec{E} \perp \vec{B} \quad (\text{polarization drift})$$

* For $\rho \rightarrow 0$, $\epsilon_R \rightarrow 1$, like vacuum.

$B \rightarrow \infty$, $\epsilon_R \rightarrow 1$, \because polarization drift

(rotates crazily)



v_p vanishes and the particles do not move in response to the transverse electric field.

* In usual laboratory, e.g. $n = 10^{16} \text{ m}^{-3}$, $B = 0.1 \text{ T}$

$$\frac{\mu_0 \rho c^2}{B^2} = \frac{4\pi \times 10^{-7} \times 10^{16} \times 1.69 \times 10^{-27} \times (3 \times 10^8)^2}{0.1^2} = 189$$

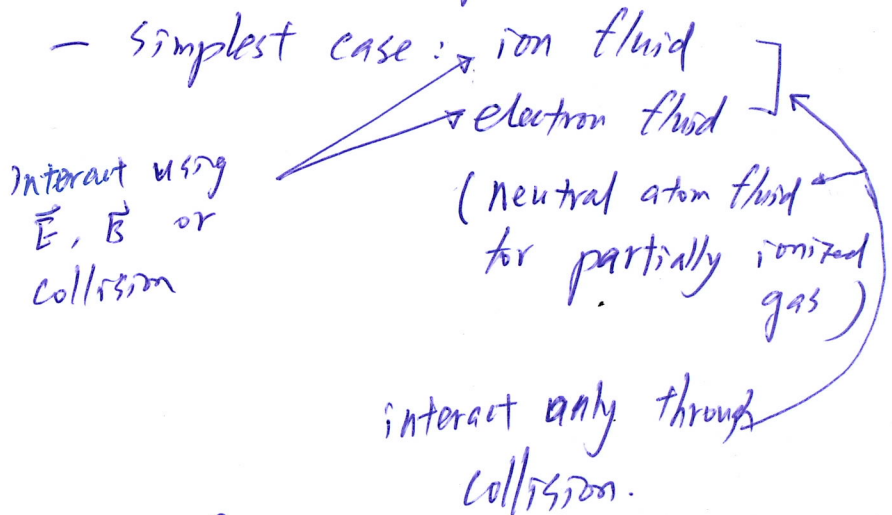
$\epsilon_R \gg 1$, \vec{E} due to the particles in the plasma greatly alter the fields applied externally.

- A plasma with large ϵ shields out altering fields, just as a plasma w/ small λ_D shields out dc fields.

§ 3.2 The fluid equation of motion. P58

$\vec{E}, \vec{B} \rightarrow$ using Maxwell's Eqs.

plasma's response — composed of two or more interpenetrating fluids, one for each species.



§ 3.2.1 The Convective Derivative.

Single particle: $m \frac{d\vec{v}}{dt} = q(\vec{E} + \vec{v} \times \vec{B})$

Assumj: $\left\{ \begin{array}{l} \text{no collisions} \\ \text{no thermal motions} \end{array} \right.$

\rightarrow all particles in a fluid element move together w/ an (average) speed \vec{u} , i.e. $\vec{v} = \vec{u}$

x^n

$m n \frac{d\vec{u}}{dt} = q n (\vec{E} + \vec{u} \times \vec{B})$

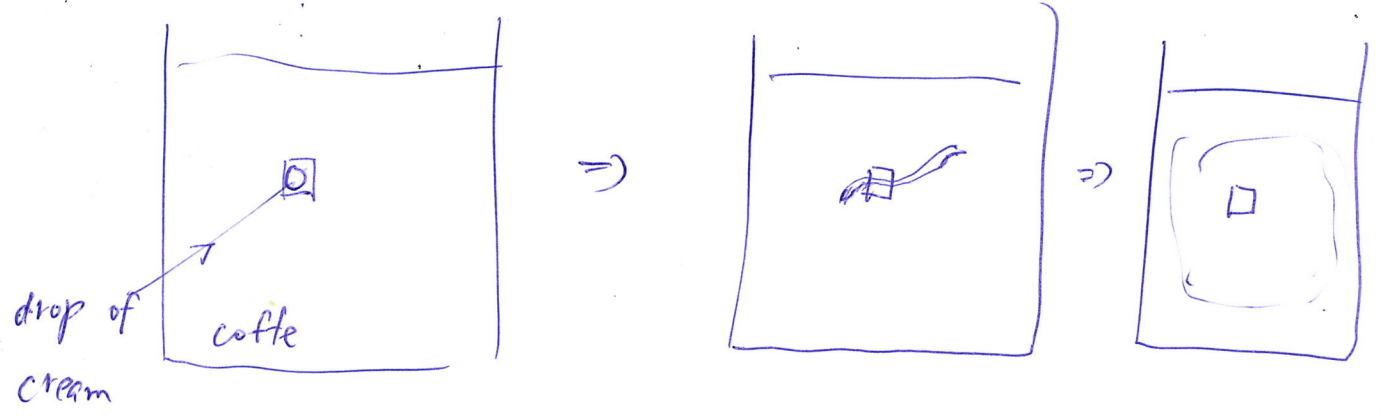


move w/ fluid

\uparrow taken at the position of the particles
 \rightarrow impractical.

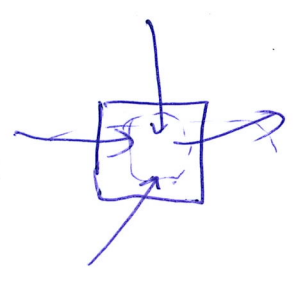
\rightarrow we wish to have an equation for fluid elements ~~to~~ fixed in space.

Example:



It's very hard to follow the drop (fluid element) since it will disperse all over the cup

Instead: the block as shown in the plot.



A fluid element at a fixed spot in the cup, retains its identity although particles continually go in and out of it.

$G(\vec{x}, t)$ - any property of a fluid

In 1D: The change of G with time:

$$\frac{dG(x, t)}{dt} = \frac{\partial G}{\partial t} + \frac{\partial G}{\partial x} \cdot \frac{dx}{dt} = \frac{\partial G}{\partial t} + u_x \frac{\partial G}{\partial x}$$

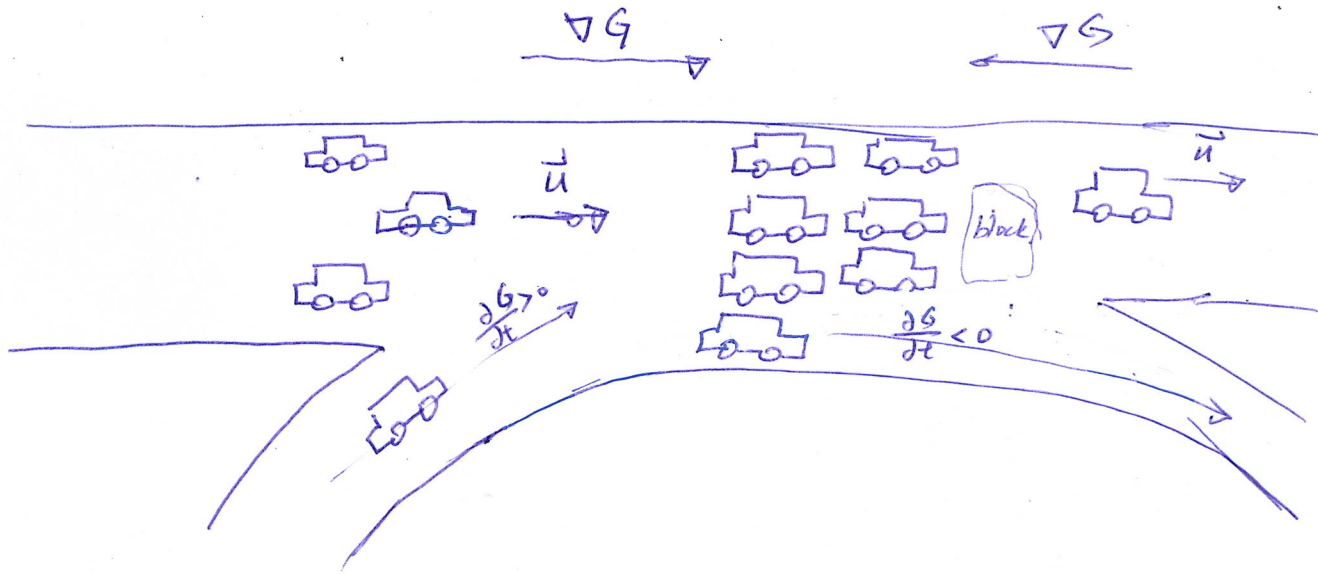
↑
change of G
@ fixed point
in space

change of G as the
observer moves w/
the fluid into a region
in which G is different

$$\Rightarrow \frac{dG}{dt} = \frac{\partial G}{\partial t} + (\vec{u} \cdot \nabla) G \quad \text{— convective derivative.}$$

$\equiv \frac{DG}{Dt}$ scalar differential operator

Example: G : the density of cars on freeway p60



$$\frac{dG}{dt} = \frac{\partial G}{\partial t} + (\vec{u} \cdot \nabla) G$$

↑ enter / exit freeway
 entering a traffic jam.

time derivative in a fixed frame.

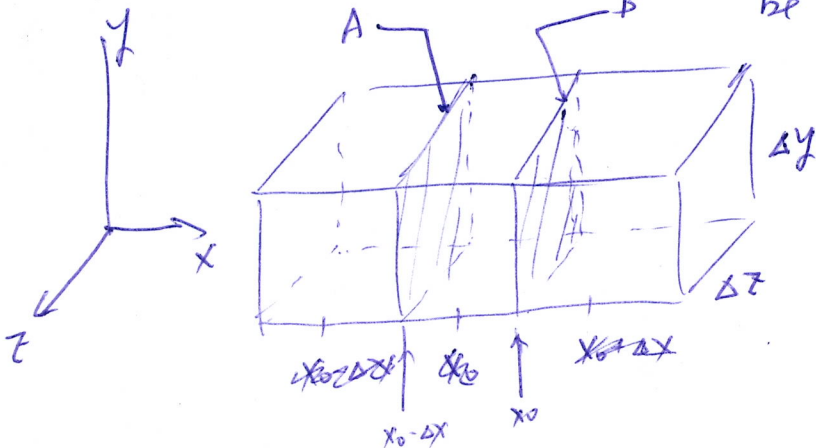
$G \rightarrow$ fluid velocity.

$$m n \frac{d\vec{u}}{dt} = \rho n (\vec{E} + \vec{v} \times \vec{B}) \Rightarrow m n \left[\frac{d\vec{u}}{dt} + (\vec{u} \cdot \nabla) \vec{u} \right] = \rho n (\vec{E} + \vec{v} \times \vec{B})$$

3.2.5 first p66

3.2.2 The stress Tensor.

- Thermal motions \rightarrow a pressure force needs to be added to



* Force arises from the random motion of particles in & out of a fluid element.

* Consider x component through the faces A & B p61

- #/ of particles per second through A w/ v_x

$$\frac{\#}{\Delta t} = \frac{N_v \cdot \Delta x \cdot \Delta y \cdot \Delta z}{\Delta t} = N_v \cdot V_x \cdot \Delta y \cdot \Delta z$$

↑
#/ of particle per m^3 w/ V_x

$$N_v = V_x \iint f(v_x, v_y, v_z) dv_y \cdot dv_z$$

- Momentum of each particle:

$$p_{\text{each}} = m v_x$$

$$\Rightarrow P_{A+} = \sum_v m v_x = \sum_v \left[N_v \cdot V_x \cdot \Delta y \cdot \Delta z \right] \cdot m v_x$$

$$\rightarrow = \sum_v \left[N_v \cdot m v_x^2 \Delta y \Delta z \right]$$

$$= \Delta y \cdot \Delta z \cdot \left[\frac{1}{2} n \cdot m \overline{v_x^2} \right]_{x_0 - \Delta x}$$

$$\rightarrow P_{B+} = \Delta y \cdot \Delta z \cdot \left[\frac{1}{2} n \cdot m \overline{v_x^2} \right]_{x_0}$$

$$P_{A+} - P_{B+} = \Delta y \cdot \Delta z \cdot \frac{1}{2} m \left(\left[n \overline{v_x^2} \right]_{x_0 - \Delta x} - \left[n \overline{v_x^2} \right]_{x_0} \right)$$

$$\rightarrow = \Delta y \cdot \Delta z \cdot \frac{1}{2} m (-\Delta x) \frac{\partial}{\partial x} (n \overline{v_x^2})$$

$$\left[n \overline{v_x^2} \right]_{x_0 - \Delta x} = \left[n \overline{v_x^2} \right]_{x_0} - \Delta x \frac{\partial}{\partial x} (n \overline{v_x^2}) + \dots$$

$$\rightarrow P_{A-} = \sum_v \left(N_v \cdot V_x \cdot \Delta y \cdot \Delta z \right) (-m v_x) = -\Delta y \Delta z \left[\frac{1}{2} n \cdot m \overline{v_x^2} \right]_{x_0 - \Delta x}$$

$$\rightarrow P_{B-} = -\Delta y \Delta z \left[\frac{1}{2} n \cdot m \overline{v_x^2} \right]_{x_0}$$

$$P_{B-} - P_{A-} = \Delta y \Delta z \cdot \frac{1}{2} m \left(\left[n \overline{v_x^2} \right]_{x_0} - \left[n \overline{v_x^2} \right]_{x_0 - \Delta x} \right)$$

$$= \Delta y \cdot \Delta z \cdot \frac{1}{2} m$$

$$P_{B-} - P_{A-} = -\Delta y \Delta z \left(\frac{1}{2}m\right) \left([n\overline{v_x^2}]_{x_0} - [n\overline{v_x^2}]_{x_0-\Delta x} \right) \quad p62$$

$$= -\Delta y \Delta z \frac{1}{2}m (\Delta x) \frac{d}{dx} (n\overline{v_x^2})$$

$$\Rightarrow \Delta P_{total} = (P_{A+} - P_{B+}) + (P_{B-} - P_{A-})$$

$$= -m \frac{d}{dx} (n\overline{v_x^2}) \Delta x \Delta y \Delta z$$

$$\Rightarrow \frac{d}{dt} (n \cdot m \cdot u_x) \Delta x \cdot \Delta y \cdot \Delta z = -m \frac{d}{dx} (n\overline{v_x^2}) \Delta x \Delta y \Delta z$$

~~total P~~

Any $v_x = u_x + v_{xr}$

\parallel
 $\overline{v_x}$ \uparrow
 \uparrow random thermal velocity
 fluid velocity

In 1D: $\frac{1}{2} m \overline{v_{xr}^2} = \frac{1}{2} kT$

$$\Rightarrow \overline{v_x^2} = \overline{(u_x + v_{xr})^2} = \overline{u_x^2 + 2u_x v_{xr} + v_{xr}^2}$$

$$= u_x^2 + 2u_x \overline{v_{xr}} + \overline{v_{xr}^2} = u_x^2 + \frac{kT}{m}$$

\parallel
0

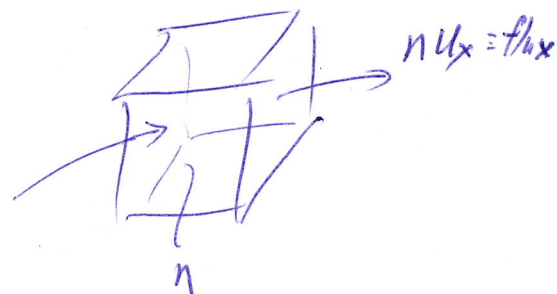
$$\Rightarrow \frac{d}{dt} (n m \cdot u_x) = -m \frac{d}{dx} \left[n \left(u_x^2 + \frac{kT}{m} \right) \right]$$

$$m n \frac{du_x}{dt} + m u_x \frac{dn}{dt} = -m u_x \frac{d(n u_x)}{dx} - m n u_x \frac{du_x}{dx} - \frac{d}{dx} (n kT)$$

~~cont~~ mass conservation (continuity)

$$\frac{\partial n}{\partial t} + \frac{d}{dx} (n u_x) = 0$$

$$\Rightarrow \text{Also, } p = n kT$$



$$mn \frac{\partial u_x}{\partial t} + mn u_x \frac{\partial u_x}{\partial x} = \underbrace{-\frac{\partial P}{\partial x}}_{\text{pressure gradient force.}} \quad p63$$

$$\Rightarrow mn \left[\frac{\partial \vec{u}}{\partial t} + (\vec{u} \cdot \nabla) \vec{u} \right] = \frac{1}{\rho} n (\vec{E} + \vec{u} \times \vec{B}) - \nabla P$$

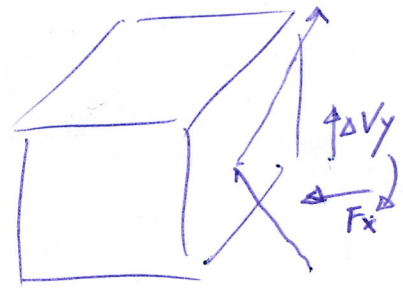
- It is a special case where ∇ transfer of x momentum by motion in the x direction.

③ the fluid is isotropic, i.e., same results in y & z .

* Shear stress \rightarrow can't be represented by a scalar p . but by a tensor \vec{P} , the stress tensor. whose component.

$$P_{ij} = mn \overline{u_i u_j}$$

in this case $-\nabla P \rightarrow -\vec{\nabla} \cdot \vec{P}$



* Two simple cases:

- isotropic Maxwellian:

$$\vec{P} = \begin{pmatrix} P & 0 & 0 \\ 0 & P & 0 \\ 0 & 0 & P \end{pmatrix}$$

$$-\vec{\nabla} \cdot \vec{P} = (dx, dy, dz) \begin{pmatrix} P & 0 & 0 \\ 0 & P & 0 \\ 0 & 0 & P \end{pmatrix} = (dxP, dyP, dzP) = -\nabla P$$

- Two temperature T_{\perp}, T_{\parallel} w/ \vec{B} .

isotropy $\perp \vec{B} \Rightarrow P_{\perp} = nkT_{\perp}, P_{\parallel} = nkT_{\parallel}$

$$\Rightarrow \vec{P} = \begin{pmatrix} P_{\perp} & 0 & 0 \\ 0 & P_{\perp} & 0 \\ 0 & 0 & P_{\parallel} \end{pmatrix} \quad -\vec{\nabla} \cdot \vec{P} = -\nabla_{\perp} P_{\perp} - \nabla_{\parallel} P_{\parallel}$$

$$\vec{P} = \begin{pmatrix} P_{11} & P_{12} & P_{13} \\ P_{21} & P_{22} & P_{23} \\ P_{31} & P_{32} & P_{33} \end{pmatrix}$$

Off-diagonal elements \rightarrow viscosity
 \Rightarrow The resistance to shear flow \uparrow

- The longer the mean free path, the farther momentum is carried, and the larger is the viscosity.

3.2.3 Collisions

\rightarrow If there is a neutral gas, the charged fluid will exchange momentum w/ it through collisions.

$$\Delta p \propto \vec{u} - \vec{u}_0$$

\uparrow
neutral fluid

$$F_1 = \frac{\Delta p}{\Delta t} \sim - \frac{mn(\vec{u} - \vec{u}_0)}{\tau}$$

$\leftarrow \Delta p$
 \leftarrow mean free time between collisions

$$\Rightarrow mn \left[\frac{\partial \vec{u}}{\partial t} + (\vec{u} \cdot \nabla) \vec{u} \right] = \underbrace{qn(\vec{E} + \vec{u} \times \vec{B})}_{\text{collisions}} - \nabla \cdot \vec{P} - \frac{mn(\vec{u} - \vec{u}_0)}{\tau}$$

3.2.4 Comparison w/ Ordinary Hydrodynamics

the same as no collisions & no collisions between different species (single specie)

Navier-Stokes eq: $\rho \left[\frac{\partial \vec{u}}{\partial t} + (\vec{u} \cdot \nabla) \vec{u} \right] = -\nabla p + \underbrace{\rho \mu \nabla^2 \vec{u}}_{\text{viscosity coefficient}}$

\uparrow
viscosity

Kinematic

$$\rho \mu \nabla^2 \vec{u} = \nabla \cdot \vec{P} - \nabla p \quad \text{w/o magnetic field.}$$

For fluid eq: frequent collisions between particles pbs

For Plasma eq: w/o explicit statement of the collision rate but still a good approximation:

- ∴ ① velocity distribution is assumed to be Maxwellian ← generally from freq. collision.
used only to take $\overline{v_x^2}$
- ⇒ Any other distribution w/ the same average would give us the same results.
Not very sensitive to Maxwellian.

~~② w/ freq. collisions the particle comes~~

②: B field !!

When a particle is accelerated (by \vec{E})
→ allowed to free stream w/o collisions
→ w/ freq. collisions, the particle comes to a limiting velocity $\propto \vec{E}$
i.e. $\vec{v} = \mu \vec{E}$ ← mobility.

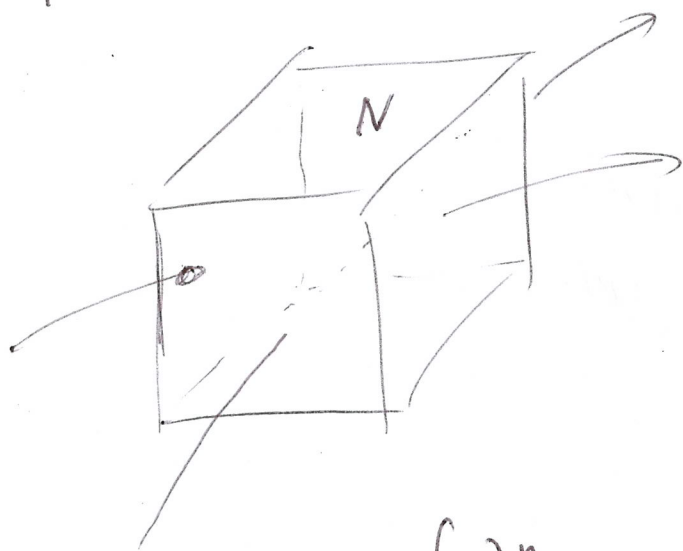
w/ B, → gyromotion, drift ($\vec{v}_E = \frac{\vec{E} \times \hat{k}}{B}$)
→ a collisionless plasma behaves like a collisional fluid.

Nevertheless, particles do free-stream along B,
→ fluid picture is not suitable.

Conclusion: For motion perpendicular to \vec{B} , the fluid theory is a good approximation.

§ 3.2.5 Equation of Continuity

Port.



$$\frac{\partial N}{\partial t} = \underbrace{\int_V \frac{\partial n}{\partial t} dV}_{\# \text{ in the box}} = - \oint \underbrace{n \vec{u}}_{\text{flux}} \cdot d\vec{s} = - \int_V \underbrace{\nabla \cdot (n \vec{u})}_{\text{divergence theory}} dV$$

⇒ it holds for any volume V.

$$\Rightarrow \frac{\partial n}{\partial t} + \nabla \cdot (n \vec{u}) = 0 \quad (+ \text{ source} - \text{ sink})$$

↓ equation of continuity or continuity equation.

§ 3.2.6 Equation of State.

- to closed the system of equations

$$\frac{\partial n}{\partial t} + \nabla \cdot (n \vec{u}) = 0$$

$$m n \left[\frac{\partial \vec{u}}{\partial t} + (\vec{u} \cdot \nabla) \vec{u} \right] = \rho n (\vec{E} + \vec{v} \times \vec{B}) - \nabla \cdot \vec{P} - \frac{m n (\vec{u} - \vec{u}_0)}{\tau}$$

① thermodynamic equation: (Adiabatic compression) $p67$

$PV^\gamma = \text{const}$ $p = C_0 n^\gamma$ $\rho = \frac{m}{V} \rightarrow \rho = \rho_0 = \text{const.}$ $\gamma \equiv \frac{C_p}{C_v}$

$\frac{\Delta p}{p} = \gamma \frac{\Delta n}{n}$ ($\rho = n \cdot m$) $\frac{\Delta p}{p} = \frac{C_0 \gamma n^{\gamma-1} \Delta n}{C_0 n^\gamma} = \gamma \frac{\Delta n}{n}$

② Isothermal compression:

$\Delta p = \Delta(nkT) = kT \Delta n + n \Delta(kT)$
 $\frac{\Delta p}{p} = \frac{kT \Delta n}{nkT} = \frac{\Delta n}{n}$
 treat as $\gamma = 1$

③ Adiabatic compression: kT will also change

$\gamma = \frac{2+N}{N}$ $\frac{\Delta p}{p} = \frac{kT \Delta n + n \Delta(kT)}{nkT}$

* The validity of the equation of state requires that heat flow be negligible, i.e., thermal conductivity is low.

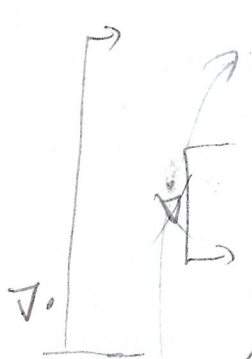
→ more likely to be true in directions $\perp \vec{B}$ than $\parallel \vec{B}$

$= \frac{\Delta n}{n} + \frac{\Delta(kT)}{kT}$
 $\equiv \gamma \frac{\Delta n}{n}$

7 3.2.7 The complete set of Fluid equation p68

$$\sigma = n_i q_i + n_e q_e$$

$$\vec{J} = n_i q_i \vec{v}_i + n_e q_e \vec{v}_e$$



$$\epsilon_0 \nabla \cdot \vec{E} = n_i q_i + n_e q_e \quad \leftarrow \quad \nabla \cdot \vec{E} = \frac{\sigma}{\epsilon_0}$$

$$\nabla \times \vec{E} = -\dot{\vec{B}} \quad \nabla \cdot (\nabla \times \vec{E}) = -\nabla \cdot \dot{\vec{B}} = -\frac{d}{dt} (\nabla \cdot \vec{B})$$

$$\nabla \cdot \vec{B} = 0 \quad \nabla \cdot \vec{B} = 0$$

$$\mu_0^{-1} \nabla \times \vec{B} = n_i q_i \vec{v}_i + n_e q_e \vec{v}_e + \epsilon_0 \dot{\vec{E}}$$

$$\nabla \cdot \left(\frac{\nabla \times \vec{B}}{\mu_0} \right) = \nabla \cdot [n_i q_i \vec{v}_i + n_e q_e \vec{v}_e] + \epsilon_0 \nabla \cdot \dot{\vec{E}}$$

$$m_i n_i \left[\frac{\partial \vec{v}_i}{\partial t} + (\vec{v}_i \cdot \nabla) \vec{v}_i \right] = q_i n_j (\vec{E} + \vec{v}_i \times \vec{B}) - \nabla P_j$$

$$= q_i \nabla \cdot (n_i \vec{v}_i) + q_e \nabla \cdot (n_e \vec{v}_e) + \epsilon_0 \nabla \cdot \dot{\vec{E}} \quad (\text{neglect collisions \& viscosity})$$

$$= -q_i \frac{\partial n_i}{\partial t} - q_e \frac{\partial n_e}{\partial t} \quad \frac{\partial n_j}{\partial t} + \nabla \cdot (n_j \vec{v}_j) = 0$$

$$+ \epsilon_0 \nabla \cdot \dot{\vec{E}} \quad P_j = \zeta_j n_j^{\gamma_j}$$

$$= \frac{\partial}{\partial t} [-q_i n_i + q_e n_e] + \epsilon_0 \nabla \cdot \dot{\vec{E}} = 0 \quad j = \text{ion, electron.}$$

Unknown: $n_i, n_e, P_i, P_e, \vec{v}_i, \vec{v}_e, \vec{E}, \vec{B}$ (16 unknown)

⇒ Eqs: 18 → 16 eqs

⇒ Self-consistent set of fluids.

§ 3.3 Fluid drifts perpendicular to \vec{B}

- Fluid - composed of many individual particles
 \rightarrow the fluid has drifts perpendicular to \vec{B} if the individual guiding centers have such drift.

- $\nabla p \rightarrow$ appears only in fluid.

$$m \frac{d\vec{v}}{dt} = q(\vec{E} + \vec{v} \times \vec{B})$$

$$\text{v.s. } mn \left(\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla \vec{u} \right) = qn (\vec{E} + \vec{v} \times \vec{B}) - \nabla p$$

\Rightarrow a drift associated w/ ∇p which does not occur for individual particles

$$mn \left[\frac{\partial \vec{u}}{\partial t} + (\vec{u} \cdot \nabla) \vec{u} \right] = qn (\vec{E} + \vec{v} \times \vec{B}) - \nabla p$$

① ② ③ ④

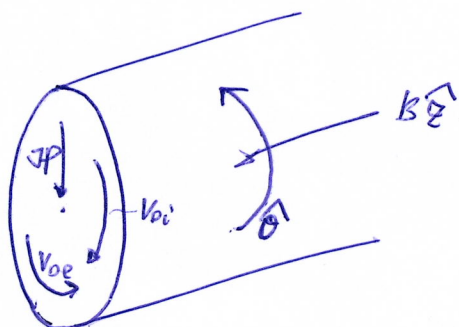
$$\frac{\text{①}}{\text{③}} \approx \left| \frac{mn i \omega v_{\perp}}{qn v_{\perp} B} \right| \approx \frac{\omega}{\omega_c}$$

where $\frac{\partial}{\partial t} \rightarrow i\omega$
 concerned only w/ v_{\perp}

$\therefore \omega \ll \omega_c, \therefore$ ① is neglected.

* neglect $(\vec{u} \cdot \nabla) \vec{u}$ for now and verify later.

* let \vec{E} & \vec{B} uniform, but $\nabla n, \nabla p \neq 0$
 this is a usual condition in MCFI



$$0 = qn [\vec{E} + \vec{v} \times \vec{B}] - \nabla P \quad \times \vec{B}$$

$$0 = qn [\vec{E} \times \vec{B} + (\vec{v}_L \times \vec{B}) \times \vec{B}] - \nabla P \times \vec{B}$$

~~$$(\vec{A} + \vec{B}) \times \vec{C} = (\vec{A} \cdot \vec{C}) \vec{B} - (\vec{A} \cdot \vec{B}) \vec{C}$$

$$(\vec{v}_L \times \vec{B}) \times \vec{B} = (\vec{v}_L \cdot \vec{B}) \vec{B} - (\vec{v}_L \cdot \vec{B}) \vec{B}$$~~

$$\vec{A} \times (\vec{B} \times \vec{C}) = (\vec{A} \cdot \vec{C}) \vec{B} - (\vec{A} \cdot \vec{B}) \vec{C}$$

$$(\vec{v}_L \times \vec{B}) \times \vec{B} = -\vec{B} \times (\vec{v}_L \times \vec{B}) = - [(\vec{B} \cdot \vec{B}) \vec{v}_L - (\vec{B} \cdot \vec{v}_L) \vec{B}]$$

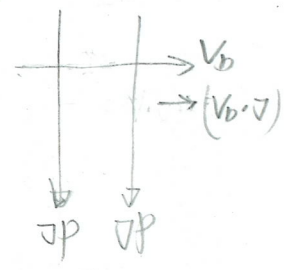
$$= -B^2 \vec{v}_L + (\vec{B} \cdot \vec{v}_L) \vec{B} = -B^2 \vec{v}_L$$

$$\Rightarrow 0 = qn [\vec{E} \times \vec{B} - B^2 \vec{v}_L] - \nabla P \times \vec{B}$$

$$\text{or } \vec{v}_L = \frac{\vec{E} \times \vec{B}}{B^2} - \frac{\nabla P \times \vec{B}}{qn B^2} \equiv \underbrace{\vec{v}_E}_{\vec{E} \times \vec{B} \text{ drift}} + \underbrace{\vec{v}_D}_{\text{Diamagnetic drift}}$$



For $\vec{E} = 0$, $\vec{v} = \vec{v}_D \perp \nabla P$



to verify neglecting ∇v should $\parallel \nabla P$
 $\therefore (\vec{v} \cdot \nabla) \vec{v} = 0$
 $(\vec{v}_D \cdot \nabla) \vec{v}_D \perp \vec{v}_D$
 $\nabla P \parallel \nabla \phi$ For $\vec{E} \neq 0$, $\vec{E} = -\nabla \phi \neq 0$

$\vec{v}_E = \frac{\vec{E} \times \vec{B}}{B^2}$
 $\vec{v}_E \perp \nabla P \parallel \nabla \phi$

if $\nabla \phi$ & ∇P are in the same direction.
 $\Rightarrow \vec{E} \parallel \nabla P \Rightarrow \vec{v} = \vec{v}_E + \vec{v}_D \perp (\vec{E}, \nabla P)$

$$\vec{v} \perp \nabla P \Rightarrow (\vec{v} \cdot \nabla) \vec{v} = 0$$

$\nabla P \perp \nabla v$ $(\vec{v}_E \cdot \nabla) \vec{v}_E \perp \vec{v}_E$
 $p = nkT$

Note that $p = C n^{\gamma} \Rightarrow \frac{\nabla P}{p} = \gamma \frac{\nabla n}{n}$

$$\vec{v}_D = - \frac{\nabla P \times \vec{B}}{qn B^2} = -\gamma \frac{p}{n} \frac{\nabla n \times \vec{B}}{qn B^2} = \gamma \frac{p}{n} \frac{(\frac{\vec{B}}{B}) \times \nabla n}{qn B} = \gamma \frac{p}{n} \frac{\vec{E} \times \nabla n}{qn B}$$

isothermal
 $p = nkT$

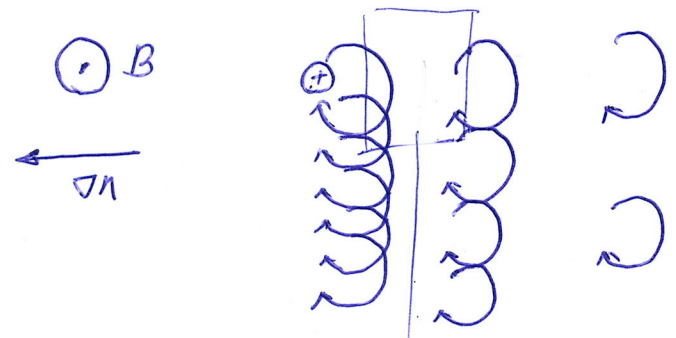
$$\vec{v}_D = \gamma \frac{kT}{qB} \frac{\vec{E} \times \nabla n}{n} = \pm \gamma \frac{kT}{eB} \frac{\vec{E} \times \nabla n}{n}$$

$$\nabla n = n' \hat{r}$$

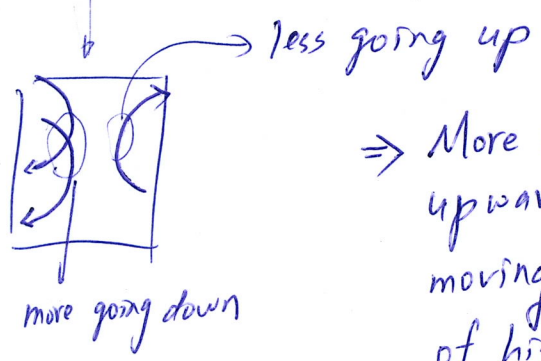
$$\Rightarrow v_{Di} = \frac{kT_i}{eB} \frac{n'}{n} \hat{\theta} \quad (n' \equiv \frac{dn}{dr} < 0)$$

$$v_{De} = -\frac{kT_e}{eB} \frac{n'}{n} \hat{\theta} \rightarrow \frac{1}{n} \frac{dn}{dr} \equiv \frac{1}{\lambda} \text{ density scale length}$$

$$\Rightarrow v_D = \frac{kT(\text{ev})}{B} \cdot \frac{1}{\lambda} \left(\frac{m}{\text{sec}} \right)$$

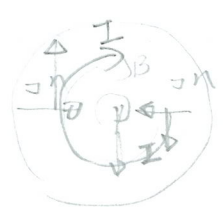
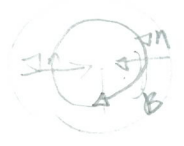


$$\hat{z} \times \hat{B} \times \nabla n$$



⇒ More ions moving downward than upward, since the downward-moving ions come from a region of higher density.

Therefore, a fluid drift perpendicular to ∇n and B even though the guiding centers are stationary.



* v_D dependent on δ

~~independent on ω~~

→ diamagnetic current, for $\delta = Z = 1$

$$\vec{J}_D = ne (\vec{v}_{Di} - \vec{v}_{De}) = (kT_i + kT_e) \frac{\vec{B} \times \nabla n}{B^2}$$

→ { particle picture : NO current if the guiding center do NOT drift
 fluid picture : current \vec{J}_D flows whenever there is a ∇n

Curvature drift, gravitational drift P12

∇B drift

(nonuniform E field drift) \rightarrow complicated

* Curvature drift:

The curvature drift exists in the fluid ~~model~~ picture since the centrifugal force is felt by all particles in a fluid element as they move around a bend in the magnetic field.

$$F_{cf} = \frac{mv_{\perp}^2}{Rc} \hat{r} \rightarrow \bar{F}_{cf} = \frac{nmv_{\perp}^2}{Rc} \hat{r}$$

$$= \frac{nkT_{\perp}}{Rc} \hat{r}$$

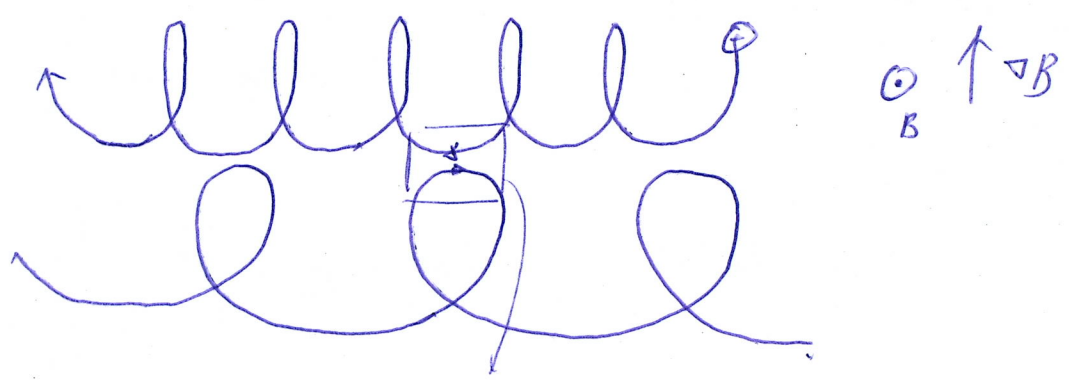
$$\Rightarrow \vec{V}_R = \frac{nkT_{\perp}}{nq} \frac{\vec{R} \times \vec{B}}{R^2 B^2} = \frac{kT_{\perp}}{q} \frac{\vec{R} \times \vec{B}}{R^2 B^2}$$

* Gravitational drift

$$\vec{F}_g = m\vec{g} \rightarrow \bar{F}_g = Mn\vec{g}$$

$$\vec{V}_g = \frac{\bar{F}_g}{nq} \times \frac{\vec{B}}{B^2} = \frac{Mn}{nq} \frac{\vec{g} \times \vec{B}}{B^2} = \frac{M}{q} \frac{\vec{g} \times \vec{B}}{B^2}$$

* ∇B drift: does NOT exist for fluids



\rightarrow Curvature are different but V_{\perp} are the same (\because No \vec{E}) & # of particles are the same \Rightarrow perfect cancellation.

3.4 Fluid drifts parallel to \vec{B}

$$\vec{B} = B \hat{z}$$

$$mn \left[\frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \nabla) \vec{v} \right] = qn (\vec{E} + \vec{v} \times \vec{B}) - \nabla p$$

For \hat{z} component, (v_z)

$$mn \left[\frac{\partial v_z}{\partial t} + (\vec{v} \cdot \nabla) v_z \right] = qn \left(E_z + (\vec{v} \times \vec{B})_z \right) - \frac{\partial p}{\partial z}$$

for $\ll \frac{\partial v_z}{\partial t}$, e.g., v_z is spatially uniform $\therefore (\vec{v} \times \vec{B})_z = 0$

$$\Rightarrow \frac{\partial v_z}{\partial t} = \frac{q}{m} E_z - \frac{1}{mn} \frac{\partial p}{\partial z}$$

For $\frac{\partial p}{p} = \gamma \frac{\partial n}{n}$ & $p = n k T$

$$\frac{1}{mn} \frac{\partial p}{\partial z} \frac{p}{p} = \frac{1}{mn} \frac{\partial n}{\partial z} \gamma \frac{p}{n} = \gamma \frac{1}{mn} \frac{\partial n}{\partial z} \gamma \frac{kT}{n}$$

$$= \frac{\gamma k T}{mn} \frac{\partial n}{\partial z}$$

$$\Rightarrow \frac{\partial v_z}{\partial t} = \frac{q}{m} E_z - \frac{\gamma k T}{mn} \frac{\partial n}{\partial z}$$

\Rightarrow The fluid is accelerated along \vec{B} under the combined electrostatic and pressure gradient forces.

for $m \rightarrow 0$ (me) $\frac{\partial v_z}{\partial t} \ll$ the other two terms $\Rightarrow \gamma E_z \sim \frac{\gamma k T}{n} \frac{\partial n}{\partial z}$

with $q = -e$, $E = -\nabla \phi = -\frac{\partial \phi}{\partial z}$

$$\Rightarrow e \frac{\partial \phi}{\partial z} = \frac{\gamma k T_e}{n} \frac{\partial n}{\partial z}$$

- Electrons are so mobile that their heat conductivity is almost infinite. \Rightarrow assume isothermal electron, $\gamma = 1$

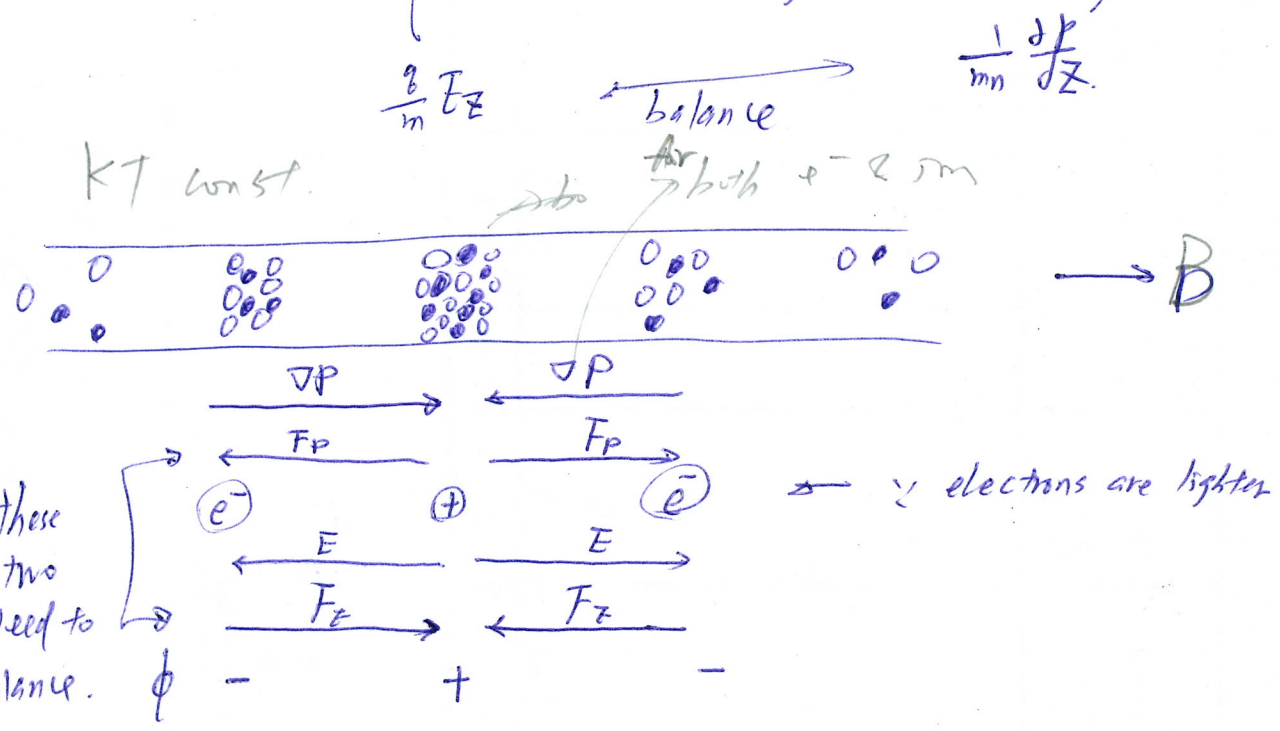
$$e \int d\phi = \int_{n_0}^n \frac{dn}{n} \Rightarrow e\phi = \int_{n_0}^n kT_e \ln \frac{n}{n_0} = kT_e \ln n + C - kT_e \ln n_0$$

$$\Rightarrow n = n_0 \exp\left(\frac{e\phi}{kT_e}\right)$$

Boltzmann relation for electrons !)

→ physical picture:

- electron → accelerated to high energy very quickly w/ time $\frac{\partial v_z}{\partial t} \approx 0$
- ions are left behind
- electrostatic & pressure gradient forces on e^- must be closely in balance



$$\Rightarrow n = n_0 \exp\left(\frac{e\phi}{kT_e}\right)$$

* There is enough charge to set up the \vec{E} field required to balance the forces on the electrons.

3.5 The plasma approximation. p75

For Poisson's eq: $\nabla \cdot \vec{E} = \frac{\sigma}{\epsilon_0}$ ← from $\sigma \rightarrow \vec{E}$

In plasma: $\left\{ \begin{array}{l} \vec{E} \text{ is found from the equations of motion} \\ \text{Poisson's eq is to find } \sigma \end{array} \right.$

∴ Plasma has tendency to remain neutral.

\vec{E} must adjust itself so that the orbits of the electrons and ions preserve neutrality.

$E \rightarrow \sigma$ → $\left\{ \begin{array}{l} \text{The charge density is of secondary importance; it} \\ \text{will adjust itself so that Poisson's eq. is satisfied.} \end{array} \right.$

— True only for low-freq. motions.

— In plasma, $n_i = n_e$ and $\nabla \cdot \vec{E} \neq 0$!!

↳ plasma approximation.

↳ Do NOT use Poisson's eq. to obtain \vec{E} unless it is unavoidable.

↳ as long as motions are slow enough, that both ions and electrons have time to move, it is a good approximation to $\nabla \cdot \vec{E} = \frac{\sigma}{\epsilon_0} \rightarrow n_e = n_i$

↳ If only one species can move and the other can NOT follow, e.g., high freq. electron waves, plasma approximation is not valid. \Rightarrow find \vec{E} from Maxwell's eq. not from eq. of motion.

↳ For ion waves, low freq. \Rightarrow good for plasma approximation.